

# The Oklo Phenomenon and Nuclear Data

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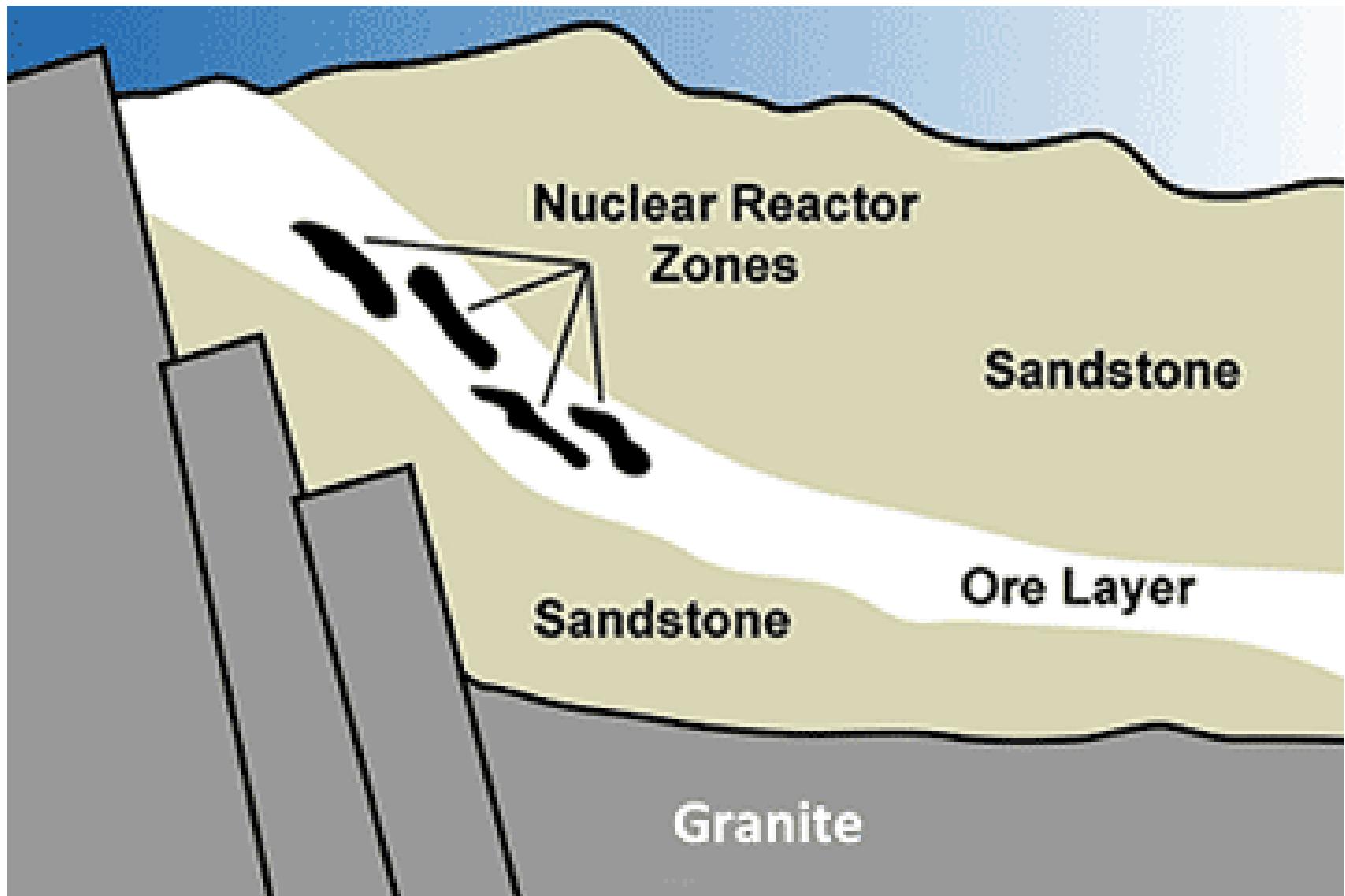
Joint Institute for Nuclear Research

of the Oklo uranium  
revealed more  
es where nuclear  
taken place.





# Oklo reactor zones



# Characteristic Data of the Oklo Phenomenon

age	( $1,75 \pm 0,10$ ) Ga
duration	200 ka – 500 ka
fluence	max. $1,6 \cdot 10^{21} \text{ cm}^{-2}$
flux density	$10^7 \text{ cm}^{-2} \text{ s}^{-1}$ – $10^8 \text{ cm}^{-2} \text{ s}^{-1}$
burn-up	10 000 MW·d/tU – 25 000 MW·d/tU
conversion factor	0,4 – 0,6
fission events	0 % - 4 % fast fission of U-238 4 % - 10 % thermal fission of Pu-239

for comparison a modern power reactor:

burn-up	33 000 MW·d/tU
flux density	$10^{13} \text{ cm}^{-2} \text{ s}^{-1}$ – $10^{14} \text{ cm}^{-2} \text{ s}^{-1}$

# Nuclear Data used for OKLO

- Neutron cross sections: S.Mughabghab, S.Suchoruchkin
- Mass distribution of fission products: JEFF-3.1.1
- Oklo mass spectroscopy isotopic ratios: Hidaka,Holliger
- Life times, beta- and  $\gamma$ -decay radiation data: ENSDF
- Decay heat: sublibrary of the base ENDF/B-VII.1
- Reactor Physics: IRPhE project, 2011-Handbook
- Photoexcitation cross sections: latest Astrophysics Refs

H Hidaka P Holliger, Geochimica et Cosmochimica Acta 62 1998 89  
R. Naudet, Oklo: des Reacteurs Nucleaires Fossiles, Eyrolles 1991

# **Plan of the talk**

- Nuclear data for OKLO study
- OKLO phenomenon and variation of  $\alpha = e^2/hc$
- Temperature problem for OKLO reactor zones
- Lutetium thermometry with 175/176 Lu-isotopes
- Analysis and suggestions on Lu-176 isomeric ratio
- Lu-176 burning through  $(\gamma, \gamma')$  channel
- Gamma ray fluxes in Oklo zone RZ10

**Co-authors: C Gould, A. Sonzogni,**  
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# OKLO and time variation of $\alpha = e^2/\hbar c$

## *Current Limits on $d/dt(\Delta\alpha/\alpha)$*

$$\Delta\alpha = (\alpha_{\text{past}} - \alpha_{\text{present}})$$

Review: J.-P. Uzan, *Liv. Rev. Relativity*, v. 14 (2011), 2

Oklo over 2 Gyr (2006)  $< 4 \times 10^{-18} \text{ yr}^{-1}$

Atomic clocks: Hg, Al (2008)  $< 4 \times 10^{-17} \text{ yr}^{-1}$

Quasars over 10 Gyr (2011)\*  $\sim 2 \times 10^{-16} \text{ yr}^{-1}$

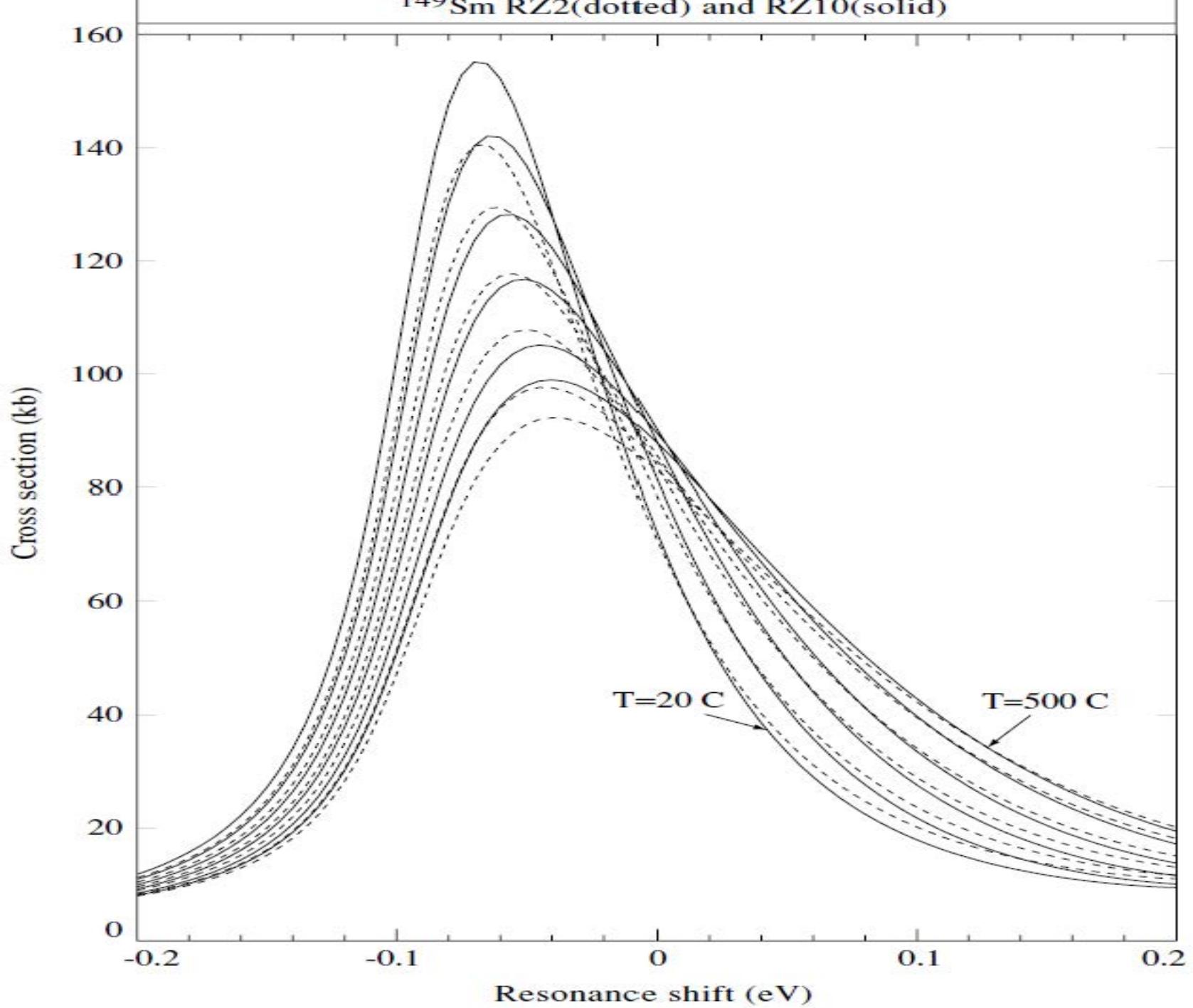
\*(smaller in the past)

C. Gould, E. Sharapov, S. Lamoreaux, *PRC* 83 (2006)

205618 Yu.Petrov et al.: G. Onegin, *PRC* 84 (2006)

205888

$^{149}\text{Sm}$  RZ2(dotted) and RZ10(solid)



# Oklo temperature problem

*Temperatures cited for Oklo vary widely*

*Indirect arguments:*

- *Damour and Dyson (96)*                    450 - 1000 C
- *Fujii et al (02)*                                180 - 400
- *Gould et al (06)*                                200- 300
- *Petrov et al (06)*                                400- 500

*Lutetium thermometry – analyze abundances of  $^{175}\text{Lu}$  and  $^{176}\text{Lu}$ ,  
Sensitivity is due to 141 eV resonance in  $^{176}\text{Lu}$*

- *Holliger and Devillers (81)*                    260-280
- *Onegin (10)*                                        100-260
- *Hidaka and Holliger (98)*                        380, or >1000 ?!

# Lutetium properties relevant for Oklo

## *Lu abundances and cross sections*

	175	176
Natural	97.401%	2.599%
Oklo (meta sample)	99.6%	0.4%

*Complications!*

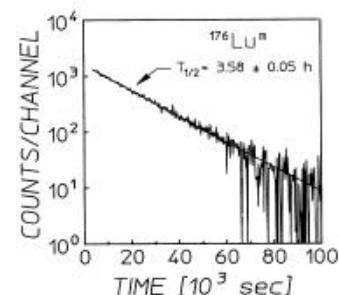
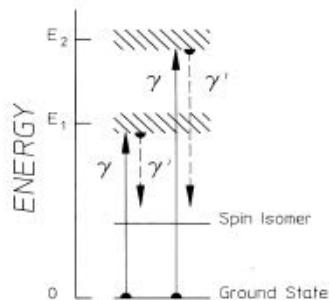
a) huge difference in capture

$$175 \sim 20b$$

$$176 \sim 2000b \quad Er=0.14$$

b) 175 capture is to  $7^{+}$  g.s as well as to 4 hr  $1^{-}$  isomer

c) 176 isomer can be photo excited in  $(\gamma, \gamma')$  reactions



# Definitions specific for Oklo study

'Effective' neutron Xs:  $\sigma_{ef} = \int \sigma(v) n(v, T) v dv / \int n(v, T) v_0 dv$

'Effective' neutron Flux:  $\Phi_{ef} = \int n(v, T) v_0 dv, v_0 = \text{thermal}$

Standard:  $\langle \Phi \rangle = \int n(v, T) v dv, \langle \sigma \rangle = \int \sigma(v) n(v, T) v dv / \int n(v, T) v dv$

$$\sigma_{ef} \Phi_{ef} = \langle \sigma \rangle \langle \Phi \rangle$$

Cross section branching:  $B_g = \sigma_g / (\sigma_g + \sigma_m)$

Burning constant:  $\lambda = \sigma_{ef} \Phi_{ef}$

Shorthand for numbers of atoms for A=176

(Lutetium):  $N6(t)$

Shorthand for neutron Xs of Lu-175, 176:  $\sigma_{ef-5}, \sigma_{ef-6}$

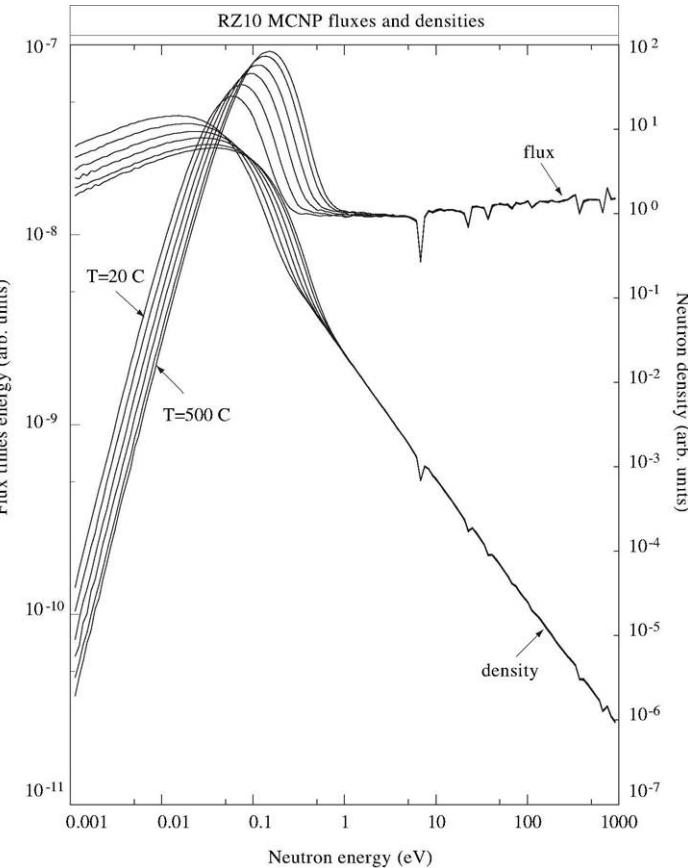
# Lutetium thermometry

$$\frac{N_6}{N_5}(\text{now}) = \frac{N_6^0}{N_5^0} \exp(-(\hat{\sigma}_6 - \hat{\sigma}_5)\hat{\Phi}t_1) + B^g \frac{\hat{\sigma}_5}{\hat{\sigma}_6 - \hat{\sigma}_5} \\ \times [1 - \exp(-(\hat{\sigma}_6 - \hat{\sigma}_5)\hat{\Phi}t_1)] \exp\left(-\frac{D\ln 2}{t_{1/2}}\right).$$

$\sigma_{\text{ef-6}}$  depends on  $T$ , so the isotopic ratio  $N_6/N_5(\text{now})$  does!

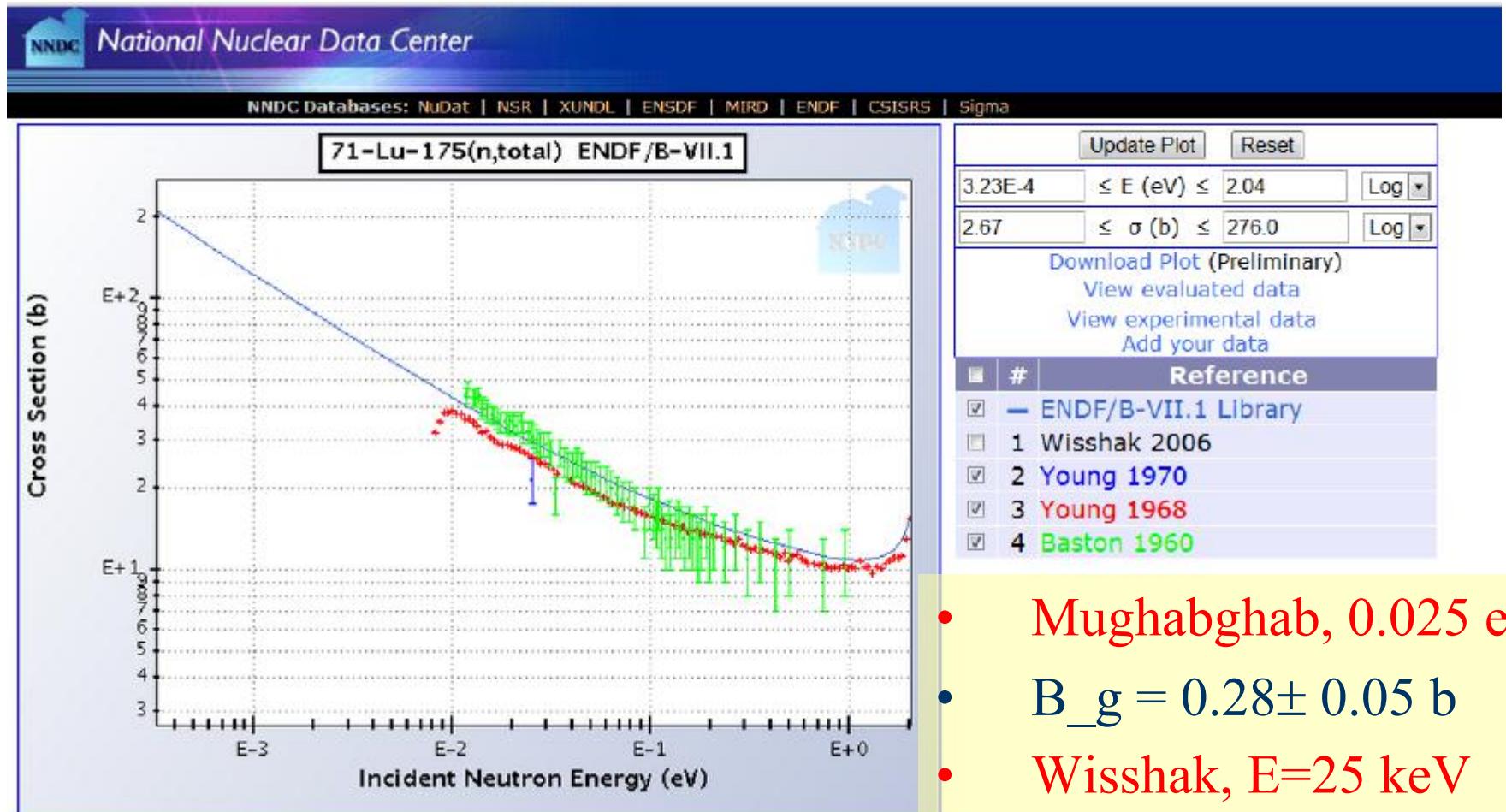
TABLE I. Lutetium cross sections  $\hat{\sigma}_5$  and  $\hat{\sigma}_6$  for the Oklo RZ1 reactor at temperatures  $T_O$  from 0 °C to 600 °C (see text).

$T_O$ (°C)	$\hat{\sigma}_5$ (kb)	$\hat{\sigma}_6$ (kt)
0	0.115	4.216
20	0.115	4.487
100	0.115	5.359
200	0.115	6.310
300	0.114	7.013
400	0.114	7.544
500	0.114	7.715
600	0.114	7.750



This slide illustrates Lu-176 burning through neutron channel. There exists however problem with the  $B_g$  parameter.

# Lu-175 branching ratio problem

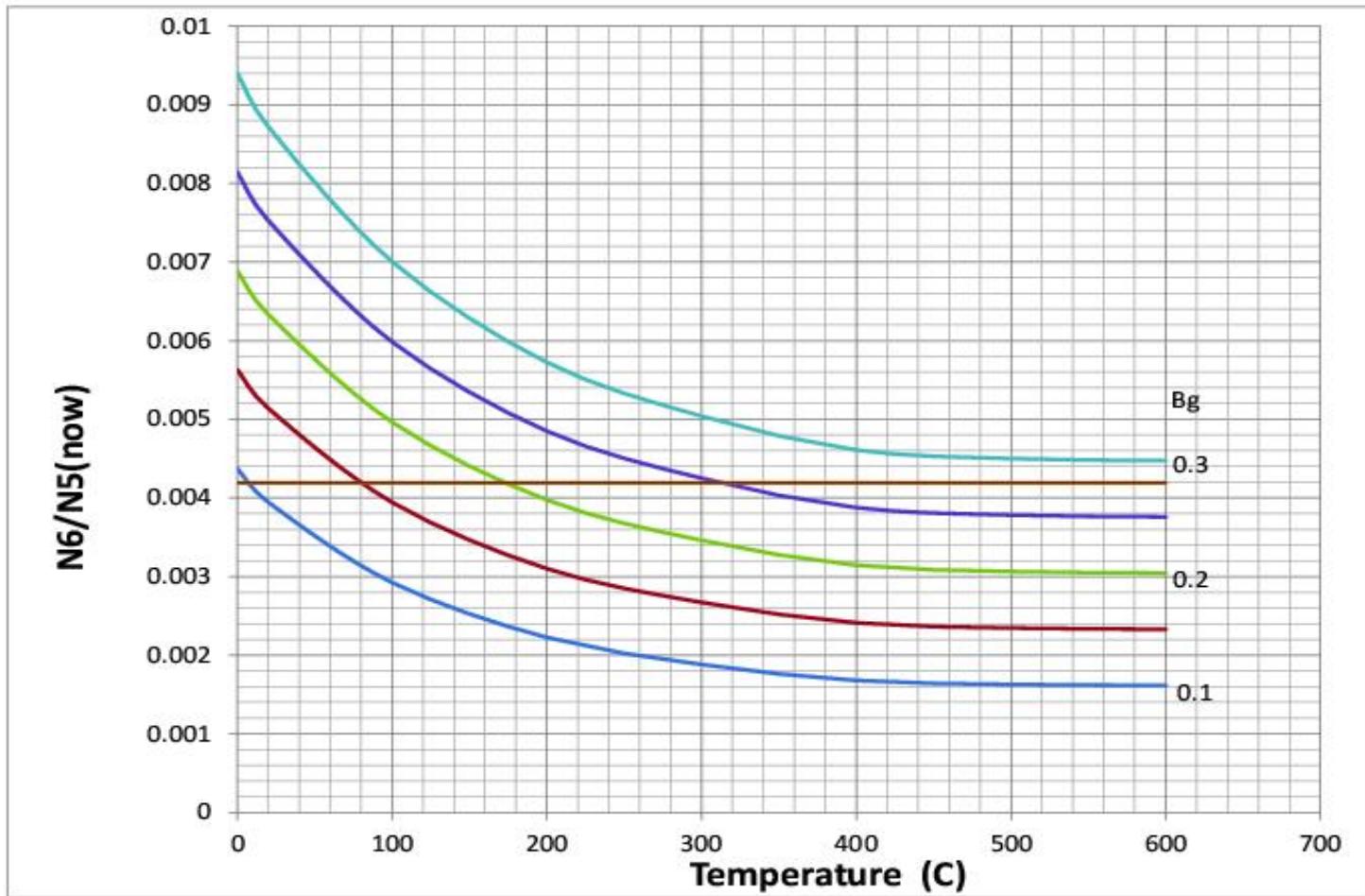


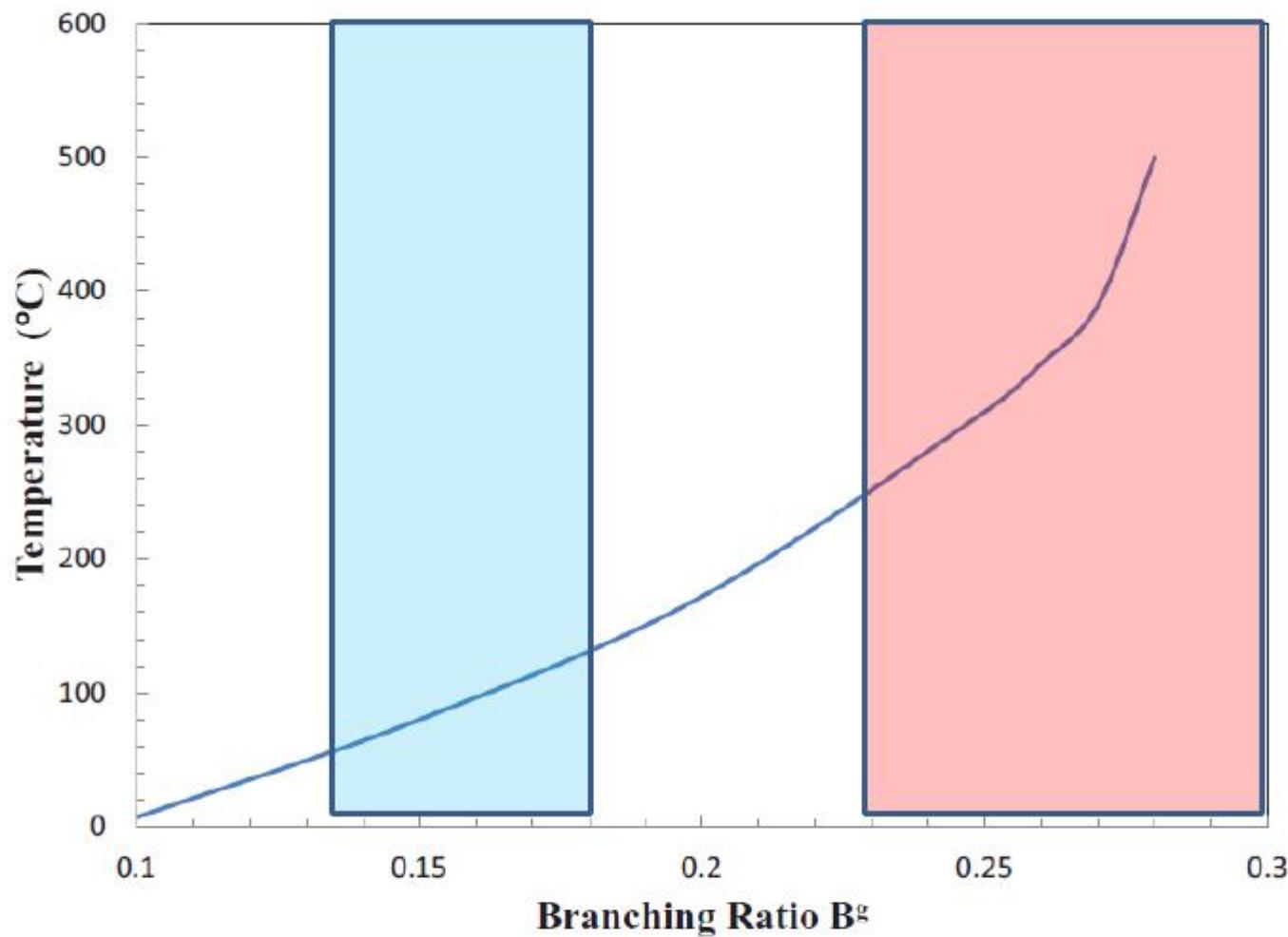
Activation cross section to Lu-176m  $\sigma_m$

$$\sigma_m = 17 \pm 2 \text{ b}; \sigma_m + \sigma_g = 23 \pm 2 \text{ b}$$

- Mughabghab, 0.025 eV
- $B_g = 0.28 \pm 0.05 \text{ b}$
- Wissak,  $E=25 \text{ keV}$
- $B_g = 0.14 \pm 0.02 \text{ b}$
- A. Baston revisited
- $B_g = 0.21 \pm 0.11 \text{ b}$

# Lutetium thermometry: continue





We suggest to perform new, Improved Measurements of the Lu-175 thermal capture and activation cross sections in order to obtain a precision value of the branching ratio  $B_g$  for thermal neutrons

# Lu-176 burning through $(\gamma, \gamma')$ reaction

$$\begin{aligned}\lambda_{\gamma\gamma'}(E_k) &= \int \Phi_\gamma(E) \sigma_{\gamma\gamma'}(E, E_k) dE \\ &= \Phi_\gamma(E_k) \sigma_{\gamma\gamma'}^{\text{int}}(E_k)\end{aligned}$$

J. Carroll, *Astrophys. J.* (1989) v.34

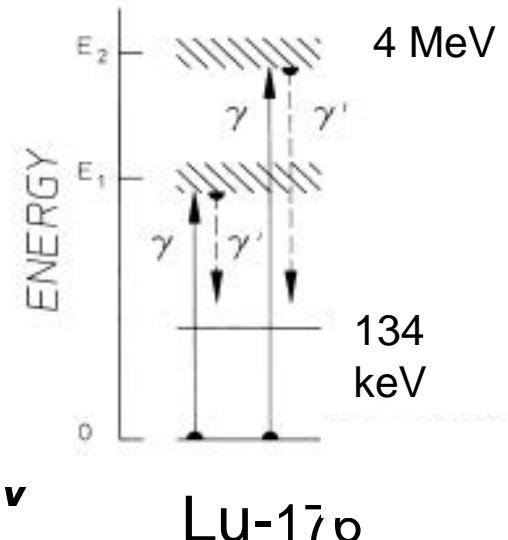
$$E_k = 839 \text{ keV}, \sigma_{\gamma\gamma'}^{\text{int}} = 33.4 \text{ mb} \bullet \text{eV}$$

$$E_k = 4.0 \text{ MeV}, \sigma_{\gamma\gamma'}^{\text{int}} = 140 \text{ mb} \bullet \text{keV}$$

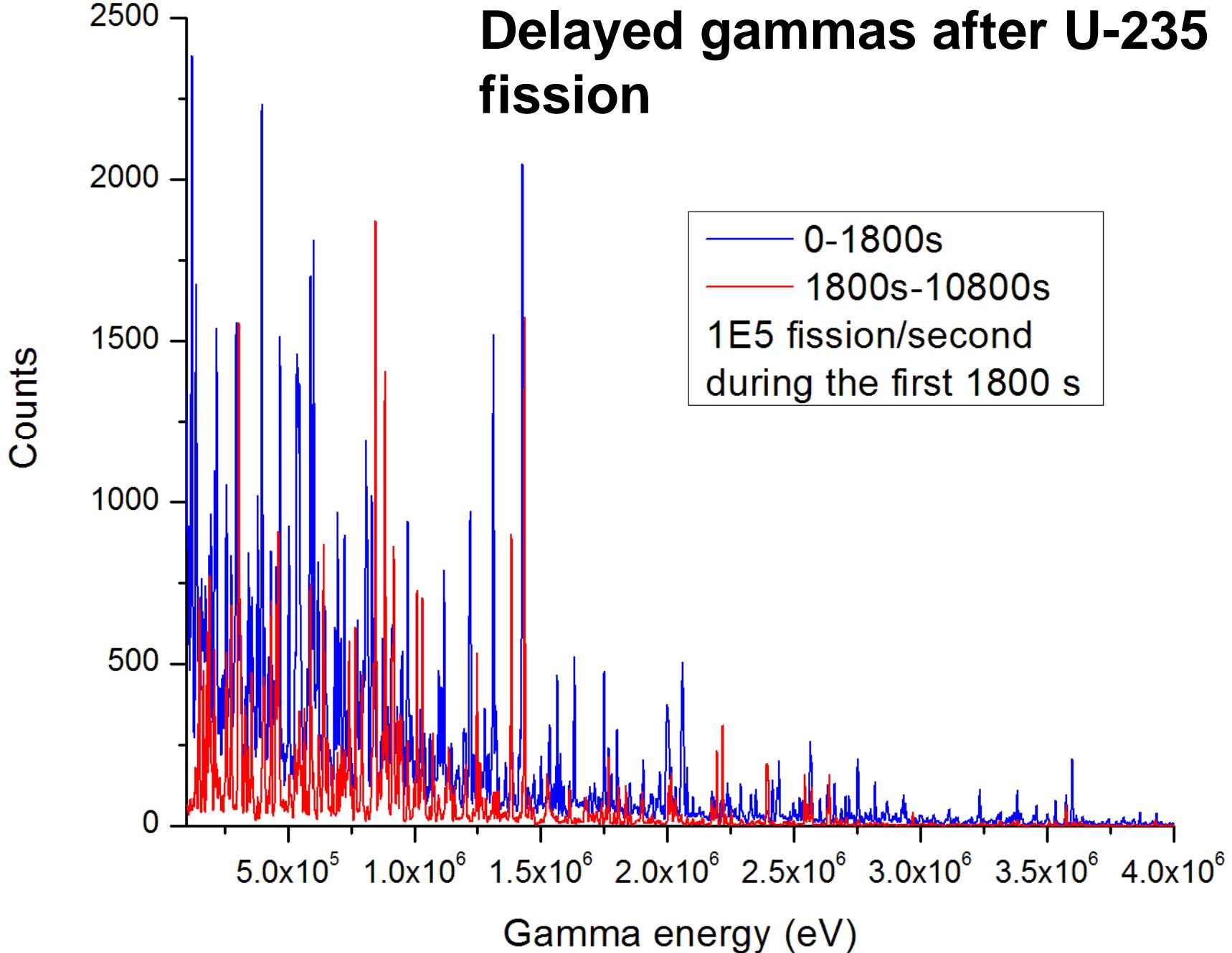
$\Phi_\gamma(E_k)$  = see calculations on next slides

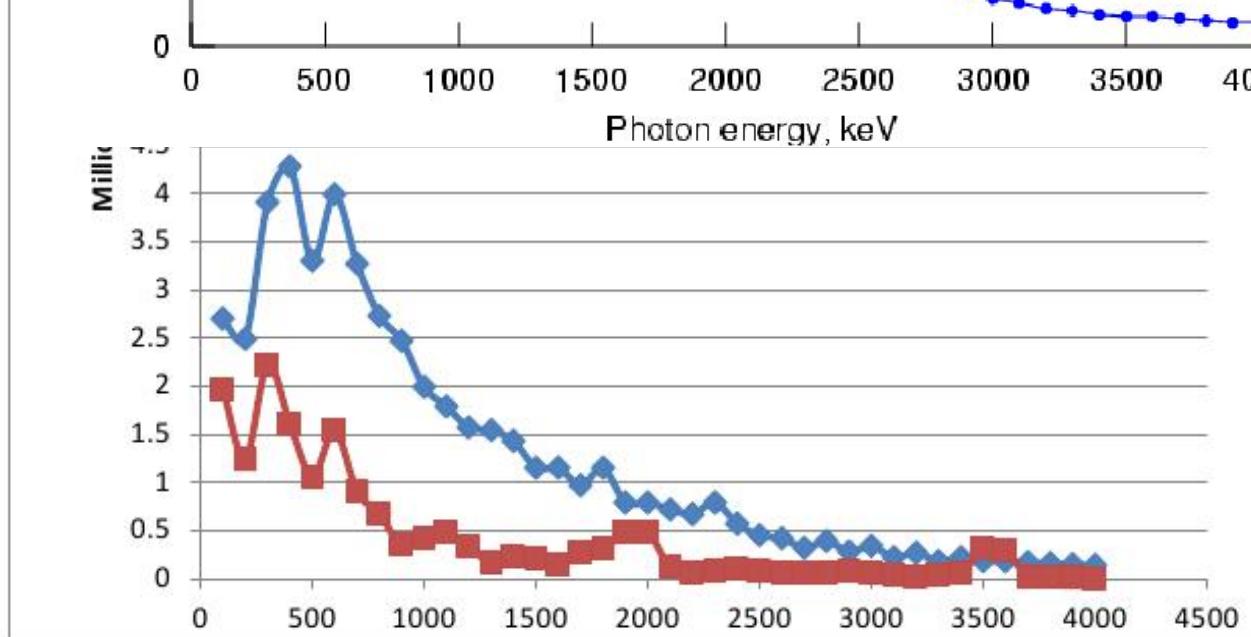
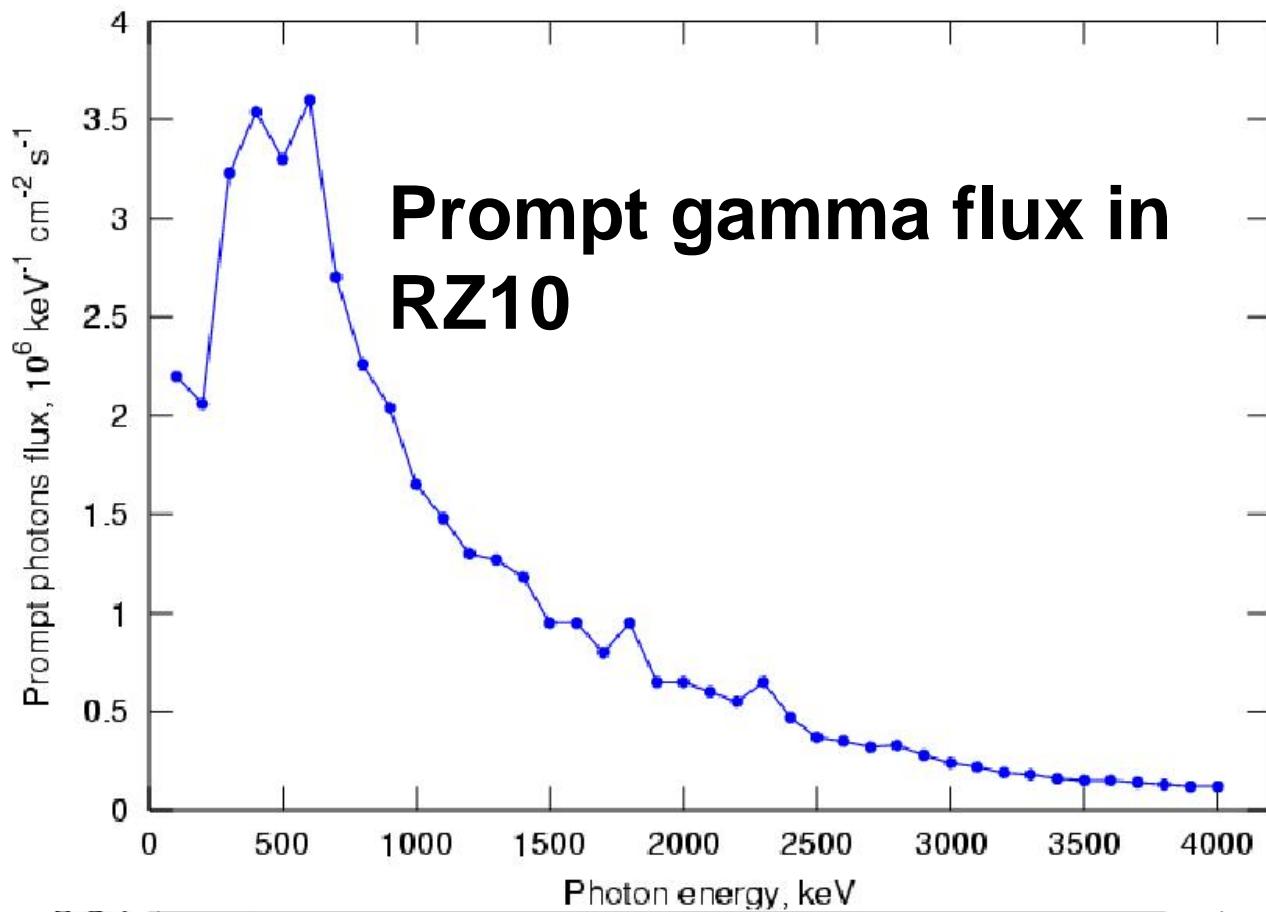
Then, for  $E_k = 4.0 \text{ MeV}$ :  $(T \text{ of the Universe} = 1.2 \bullet 10^{10} \text{ yr})$

$\lambda_{\gamma\gamma'} < 1.4 \bullet 10^{-20} \text{ s}^{-1}$ ,  $T = \ln 2 / \lambda_{\gamma\gamma'} > 1.7 \bullet 10^{12} \text{ yr}$ , This is huge!



# Delayed gammas after U-235 fission



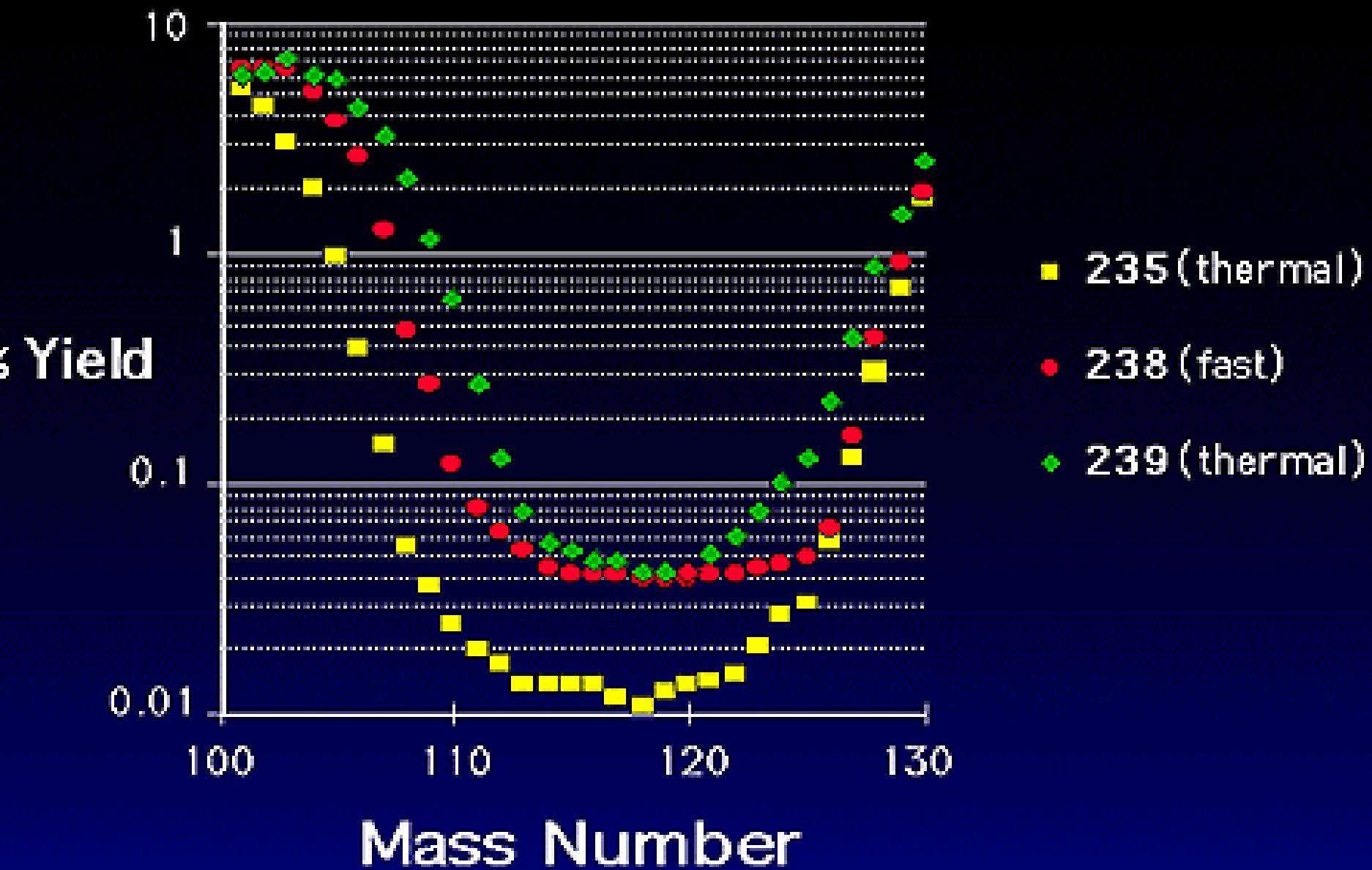


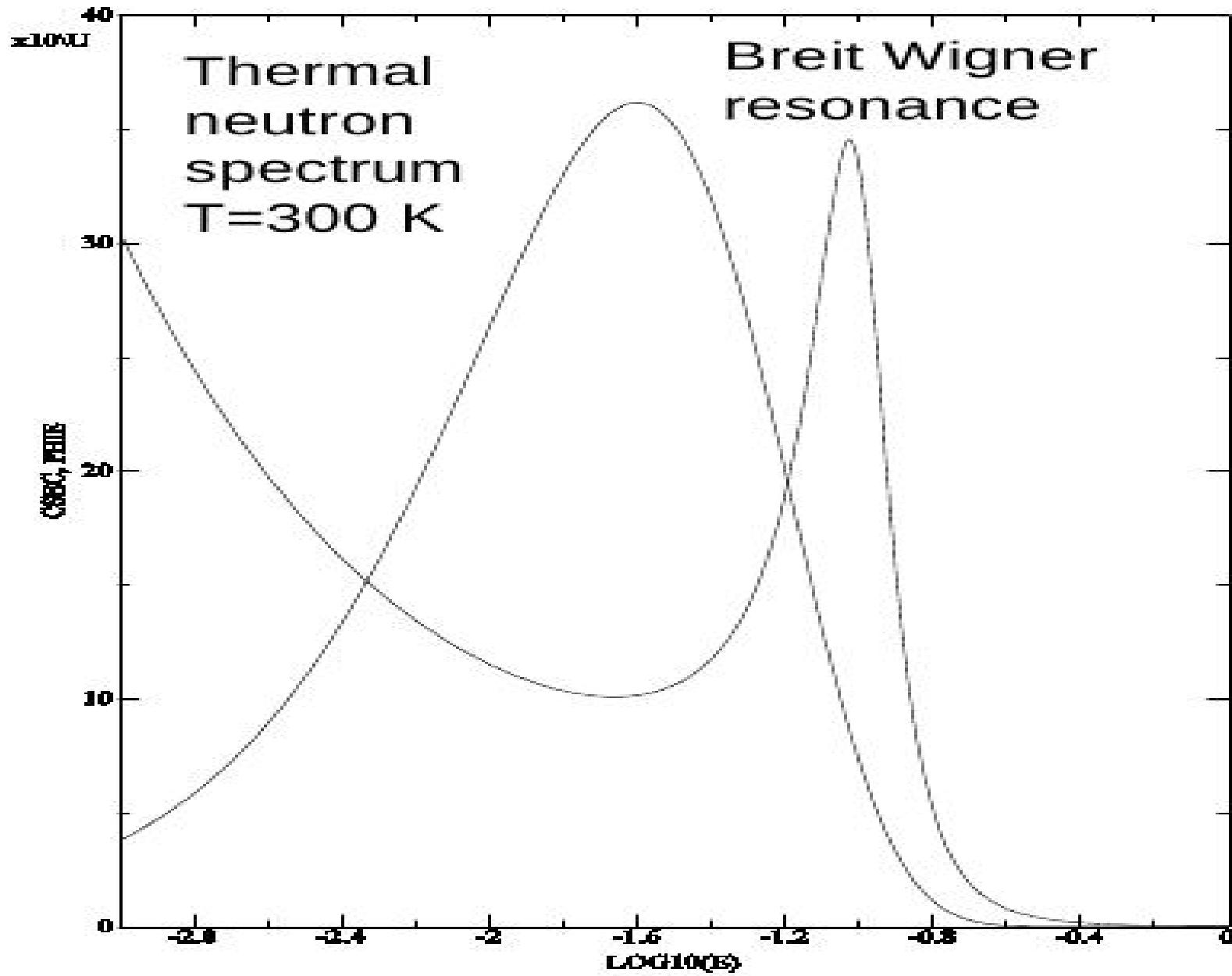


- In view of the Lu-thermometry, the final word on limits of  $\alpha$  variability from Oklo phenomenon might not have been said
- In contrast to astrophysics, the Lu-176/175 isotopic ratios in Oklo are not influenced by the  $(\gamma, \gamma')$  photoexcitation
- Applicability of the Lu-thermometry to Oklo reactors studies requires improved neutron cross sections measurements for Lu-175 in the thermal energy region

**THE END**  
**THANK YOU**

# Absolute Cumulative Fission Yields





# RZ10 MCNP fluxes and densities

