

Fundamental Physics

with

(Very Cold &) Ultra-Cold Neutrons (UCNs)

at the

Institut Laue Langevin (ILL) in Grenoble, France

Peter Geltenbort

with the help of many colleagues and friends (as you can see on the numerous slides)

Very Hot (fission) Neutrons

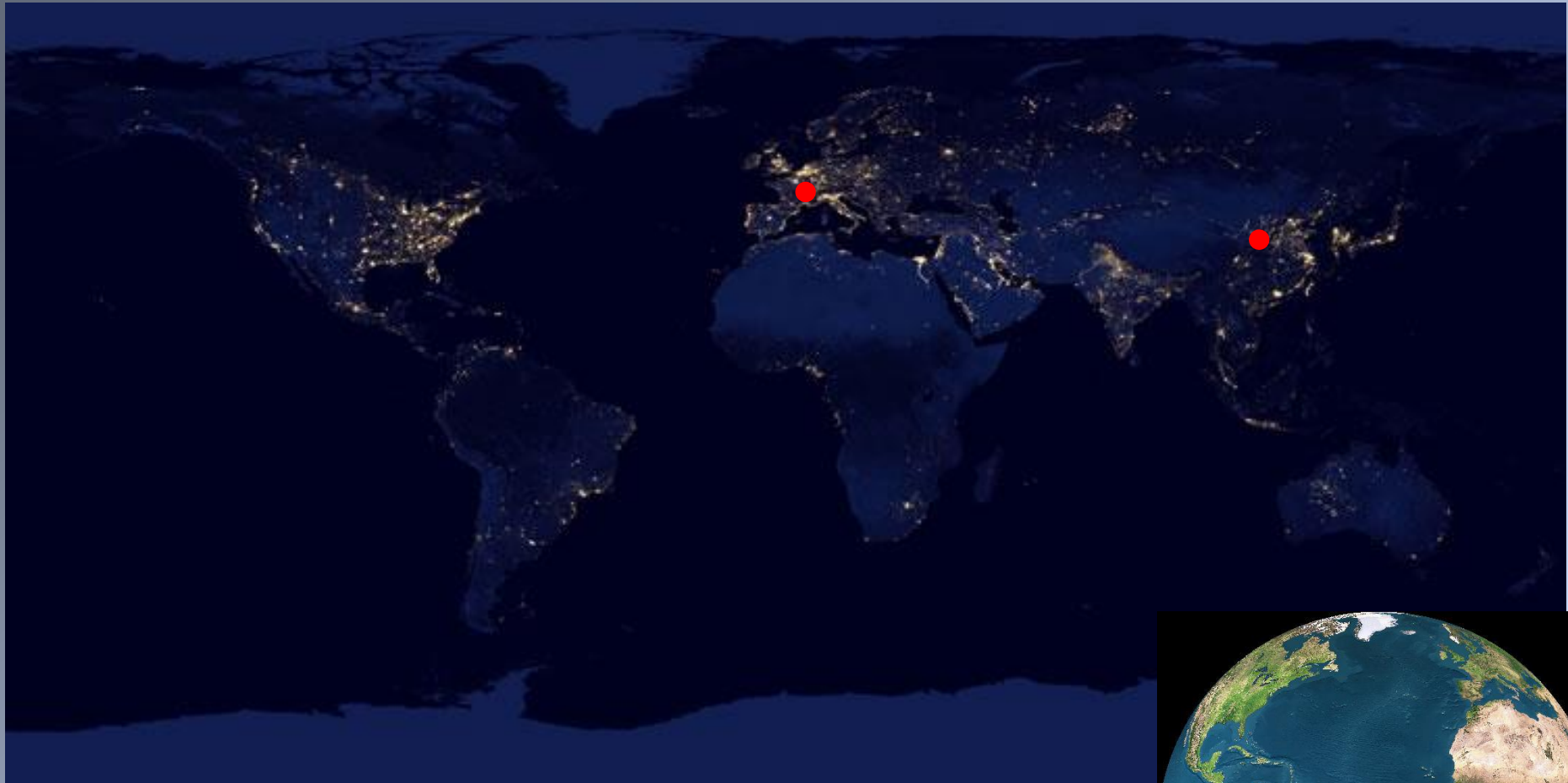
Ultracold Neutrons

10^7 eV



10^{-7} eV

Setting the scene: Satellite photo of **Earth** at night (NASA)



door to door : ~ **8 300 km** (22 hours from door to door)

This is my **first ever** visit to China! Thanks for the invitation to give a plenary talk.



Setting the scene: Europe and France

- Grenoble, France
2 650 km
- Dubna, Russia

ILL - CERN (~150 km): 1.5 hours by car
ILL - GANIL (~800 km): 7 hours by car
ILL - GSI (~750 km): 7 hours by car



Setting the scene: Grenoble (Capital of the French Alps)

(Host City of X Olympic Winter Games 1968)



Population: 180 000
380 000 (metro area)

Elevation: 214 m
Pic de Belledonne: 2 977 m

Amongst the flattest cities in France!

Xi'an: 8 700 000 (2015)
Metro: 12 900 000

Elevation: 405 m



- The city benefits from the **highest concentration of strategic jobs in France after Paris**, with 14% of the employments, 35,186 jobs, 45% of which specialized in design and research.
- Grenoble is also the **largest research center in France after Paris** with 22,800 jobs (11,800 in public research, 7,500 in private research and 3,500 PhD students)

[Figures from Wikipedia, June 2015]

Outline

- ILL (Institut Laue Langevin) and its high flux reactor
- NPP (Nuclear and Particle Physics group)
- Ultra-Cold Neutrons (UCNs)
- **Flagship experiments**
 - Neutron Electric Dipole Moment (nEDM)
 - Neutron lifetime (nTau)
 - Gravitational Levels (GRS)

Everything started a bit more than 50 years ago

- Proposed in 1964 (Grenoble had knowledge + inclination)
- Laboratory agreed upon in 1967 by France and Germany
 - (Scientific) Founding fathers
 - L. Néel (NP 1970, antiferromagnetism) and H. Maier-Leibnitz



Interstate treaty: 19 January 1967



Traité de l'Elysée: 22 January 1963



Louis Néel
1904 – 2000
Nobel Prize 1970
Antiferromagnetism



Heinz Maier-Leibnitz
1911 – 2000

Ideas become concrete

- Reactor first critical in 1971
- Operational for researchers (user service) in 1972 – 58 MW
- UK joined in 1973

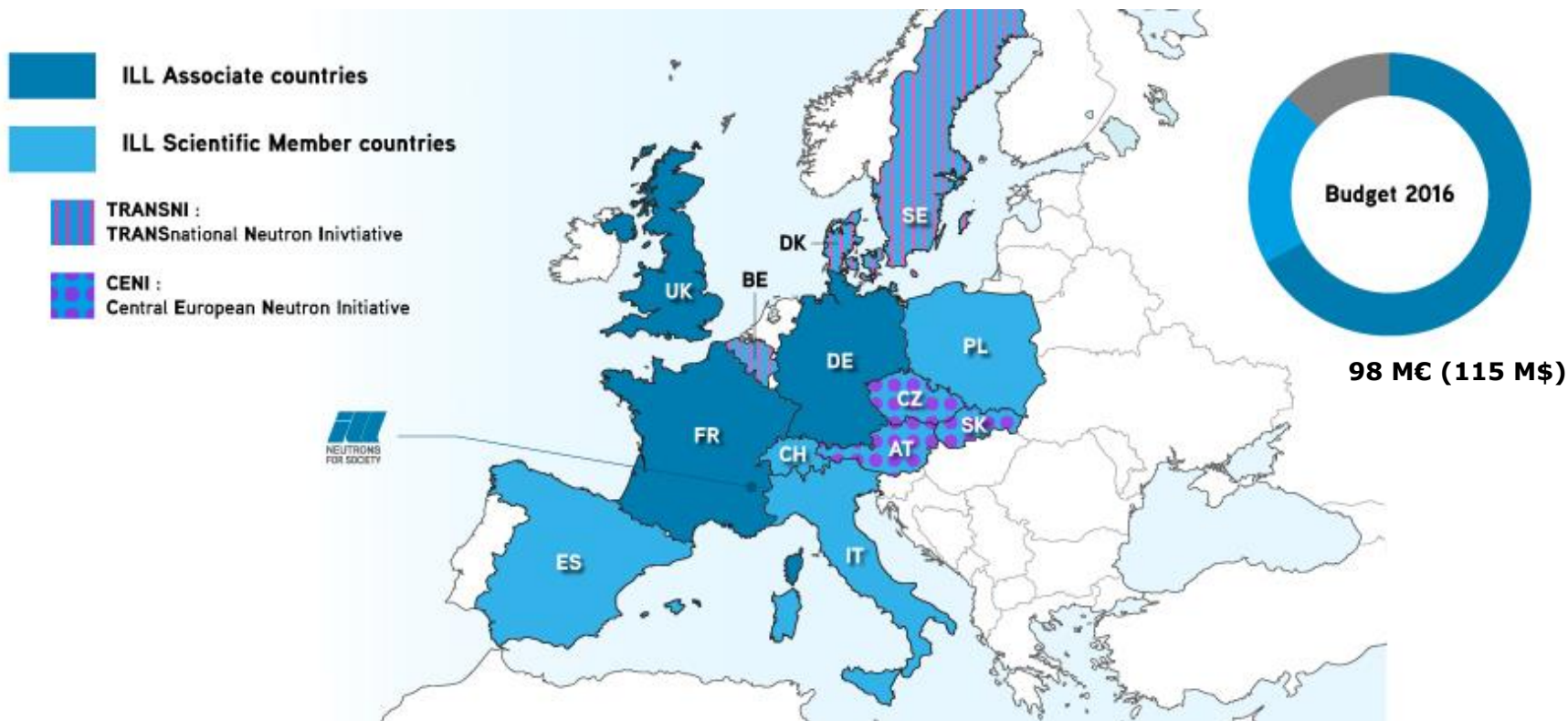
The Institut Laue-Langevin is an international research centre at the leading edge of neutron science and technology.

As the world's flagship centre for neutron science, the ILL provides scientists with a very high flux of neutrons feeding some **40 state-of-the-art instruments**, which are constantly being developed and upgraded.

As a **service institute** the ILL makes its facilities and expertise available to visiting scientists. Every year, some 1400 researchers from over 40 countries visit the ILL. More than 800 experiments selected by a scientific review committee are performed annually. Research focuses primarily on fundamental science in a variety of fields: condensed matter physics, chemistry, biology, nuclear physics and materials science, etc.



THE ILL MEMBER COUNTRIES AND THEIR FINANCIAL PARTICIPATION



KEY FIGURES ABOUT THE ILL



1400 users from an active community of 12 000 scientists



850 experiments/year



**600 publications/year
>21000 pubs since 1973**



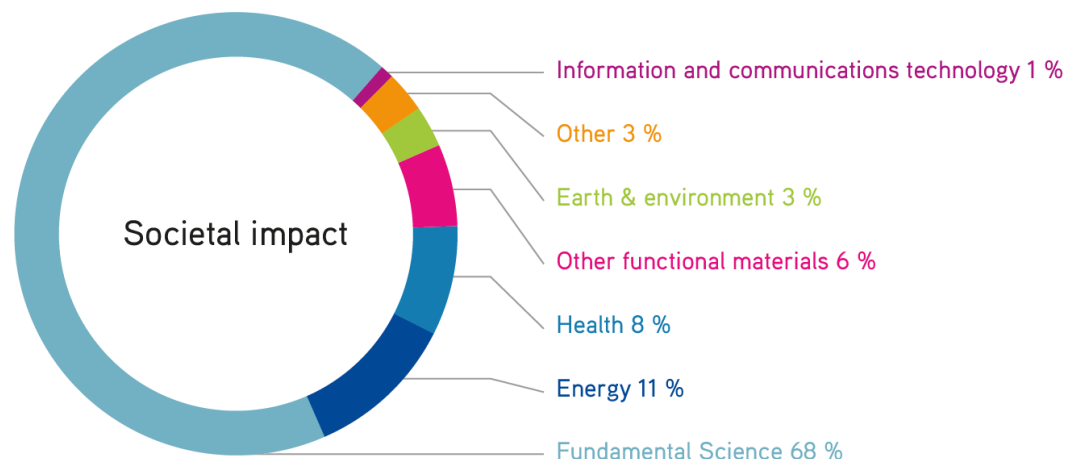
38 countries



28 instruments + 10 CRG



4 cycles of 50 days/year



Science Campus



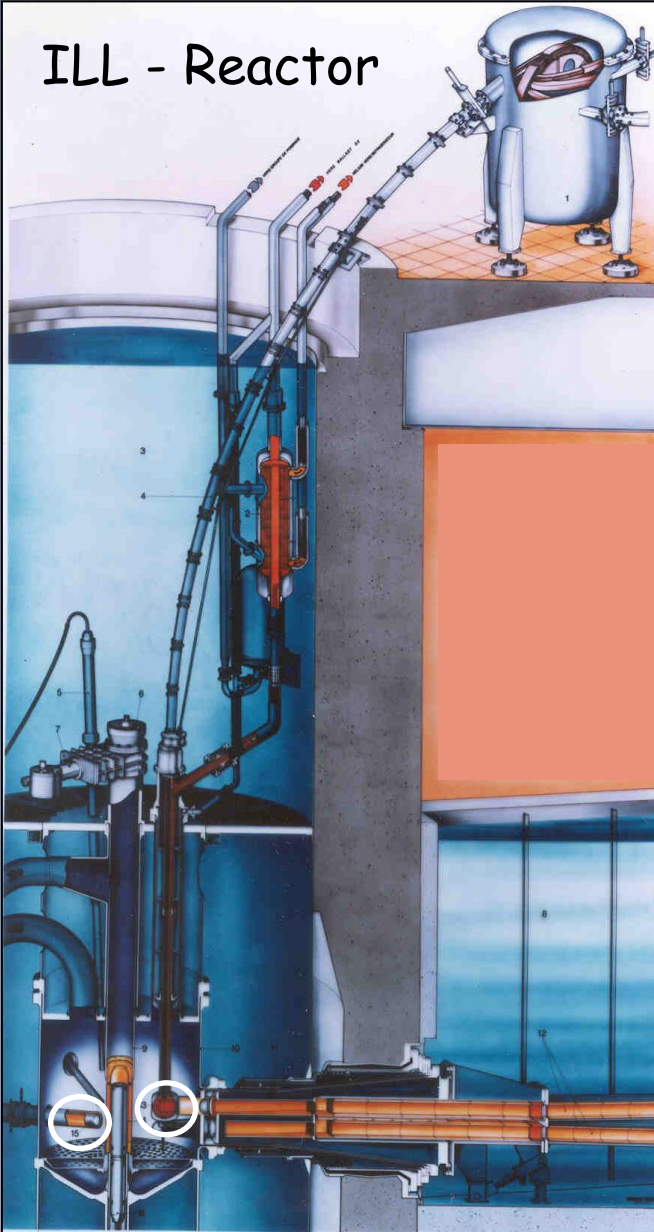
Institut Laue-Langevin (ILL) operates the most intense (reactor) neutron source in the world, feeding a suite of 40 high-performance instruments

European Synchrotron Radiation Facility (ESRF) is a world-leading synchrotron radiation source hosting 41 cutting-edge experimental stations

The Institut de Biologie Structurale (IBS) is a research centre in structural biology. The IBS possesses cutting edge facilities and is a partnership between CEA, CNRS and UJF

European Molecular Biology Laboratory (EMBL) Grenoble is an outstation of the EMBL organisation (HQ in Heidelberg), specialising in research in structural biology (in very close proximity to the ILL and the ESRF)

Neutron sources at the ILL



ILL - Reactor

Fuel (chain reaction): $^{235}\text{U}(n_{\text{th}},f) \rightarrow$ fission neutrons

Moderator: D_2O at 300K \rightarrow thermal neutrons

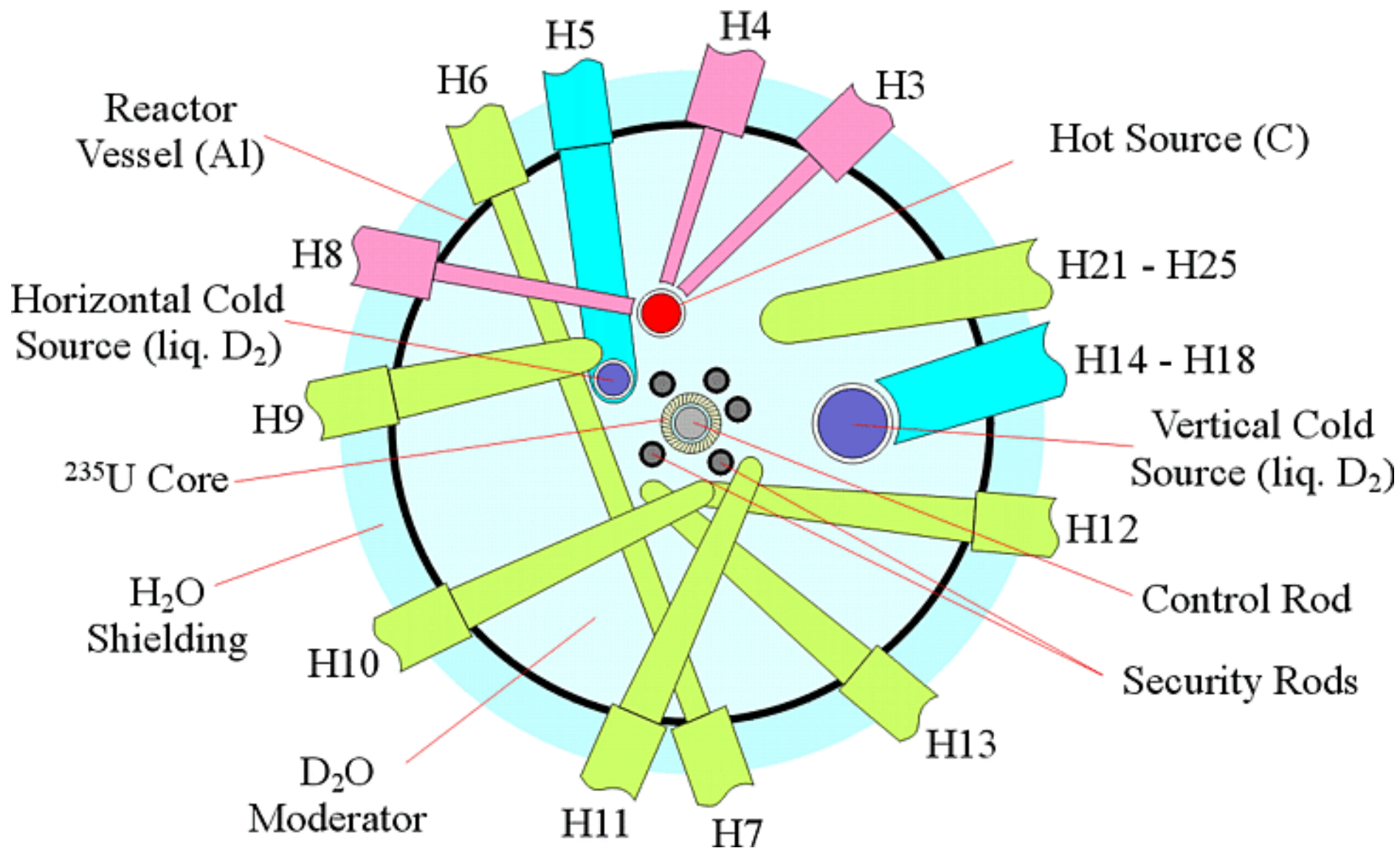
Hot source: 10 dm³ of graphite at 2400 K

Cold source^h (horizontal): 6 dm³ of liquid D_2 at 25 K

Cold source (vertical): 20 dm³ of liquid D_2 at 25 K

	<div style="border: 1px solid black; border-radius: 15px; padding: 5px; display: inline-block;"> Ultracold Neutrons </div>	<div style="border: 1px solid black; border-radius: 15px; padding: 5px; display: inline-block;"> Cold Neutrons </div>	<div style="border: 1px solid black; border-radius: 15px; padding: 5px; display: inline-block;"> Reactor Neutrons </div>
Temperature (K)	10^{-8}	10^1	10^3
Energy (eV)	10^{-7}	10^{-3}	10^{-1}
Velocity (m/s)	5	800	2200

Thermal, cold and hot neutrons



Nuclear physics

Particle physics

ILL-funded

PN1 (LOHENGRIN)
Recoil mass spectrometer for fission fragments

PF1B
Facility for cold neutrons

PN3 (GAMS)
Ultra-high resolution gamma ray spectrometer

PF2
Facility for ultracold and very cold neutrons

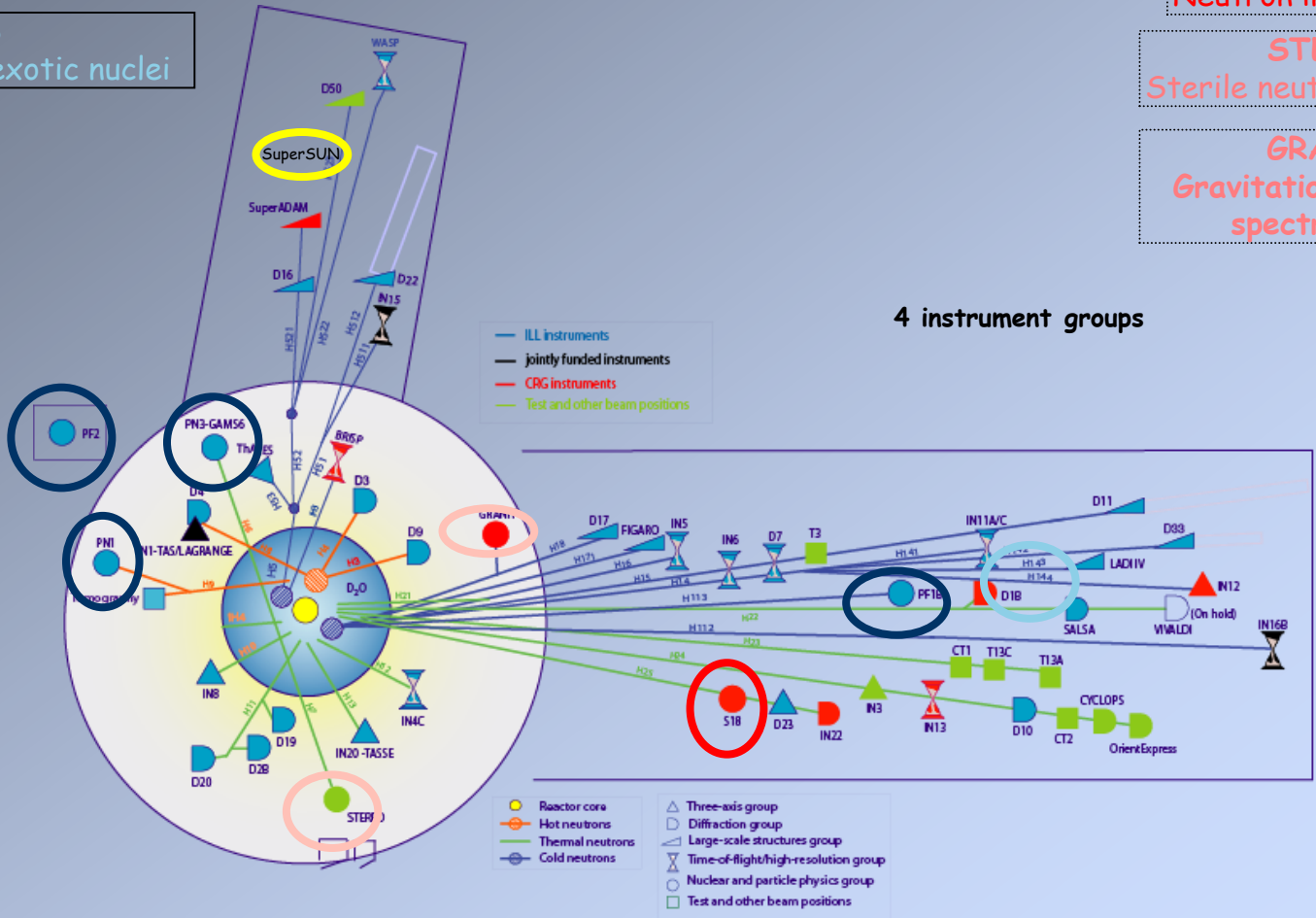
FIPPS
Spectroscopy of exotic nuclei

Collab. Research Group
(former PFI)
cryoEDM-experiment
SuperSUN

S18 - perfect crystal
Neutron interferometer

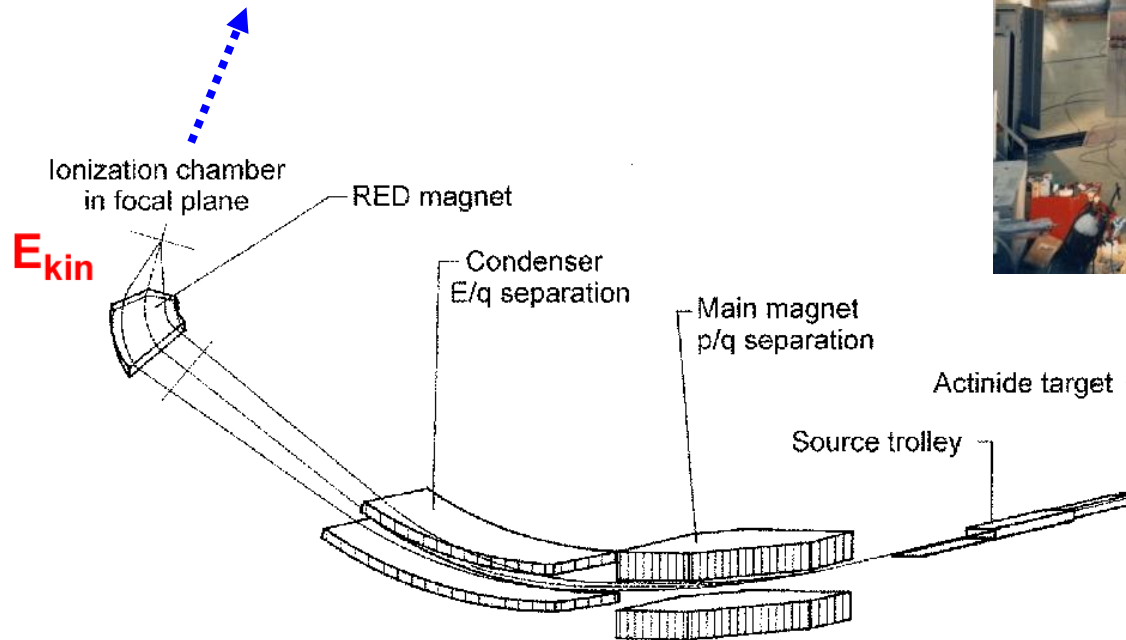
STEREO
Sterile neutrino research

GRANIT
Gravitational neutron
spectrometer



U. Koester, Y.-H. Kim, N. Laurens

mass-separated fission fragments,
up to 10^5 per second, $T_{1/2} \geq \mu\text{s}$



- n-flux $5.5 \times 10^{14} \text{ cm}^{-2}/\text{s}$
- few mg fission target (various materials)
- several 10^{12} fissions/s

$$m v^2 / r_{el} = q E$$

$$E_{kin} / q = E / 2 r_{el}$$

$$m v^2 / r_{magn} = q v B$$

$$m v / q = B r_{magn}$$

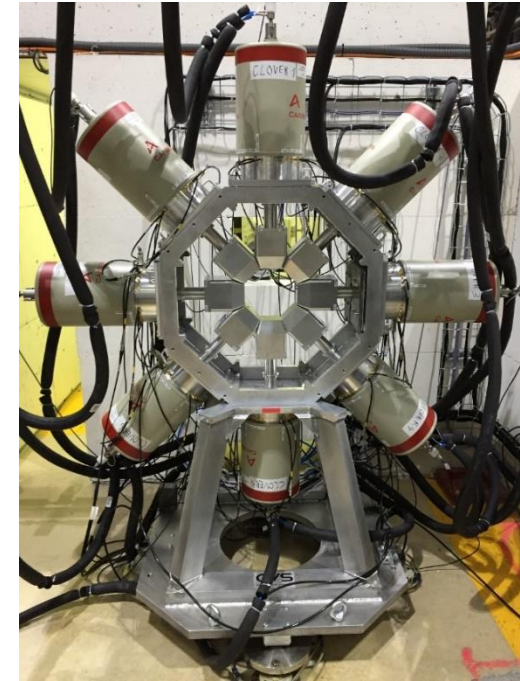
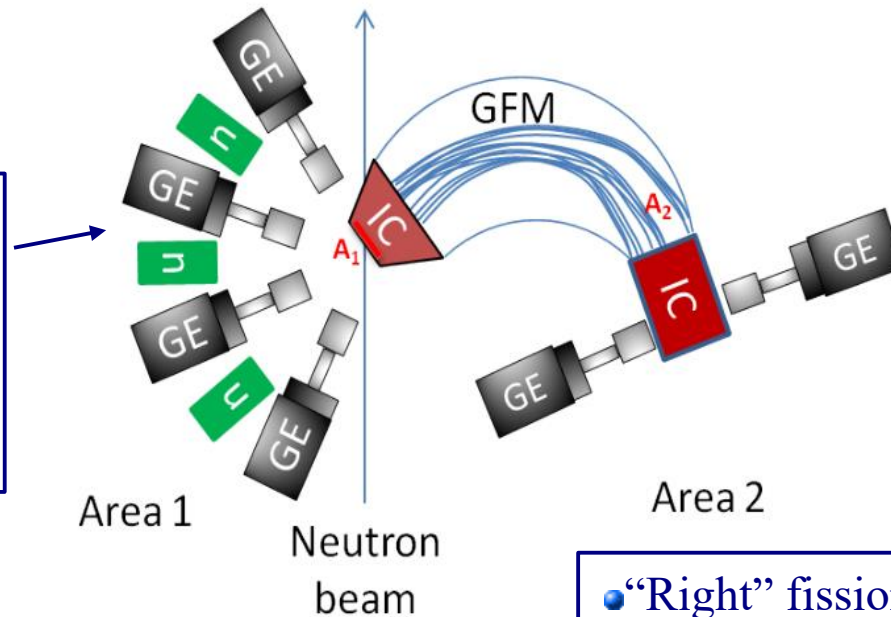
P. Armbruster et al., Nucl. Instr. Meth. 139 (1976) 213.

FIPPS: layout

Fission Product Prompt gamma-ray Spectrometer

+4 more Ge detectors
out of plane

IC: Ionization chamber
n: neutron detectors
Ge: Ge clovers



• “Left” fission fragment: stopped in backing
→ Doppler free γ detection by Ge-array

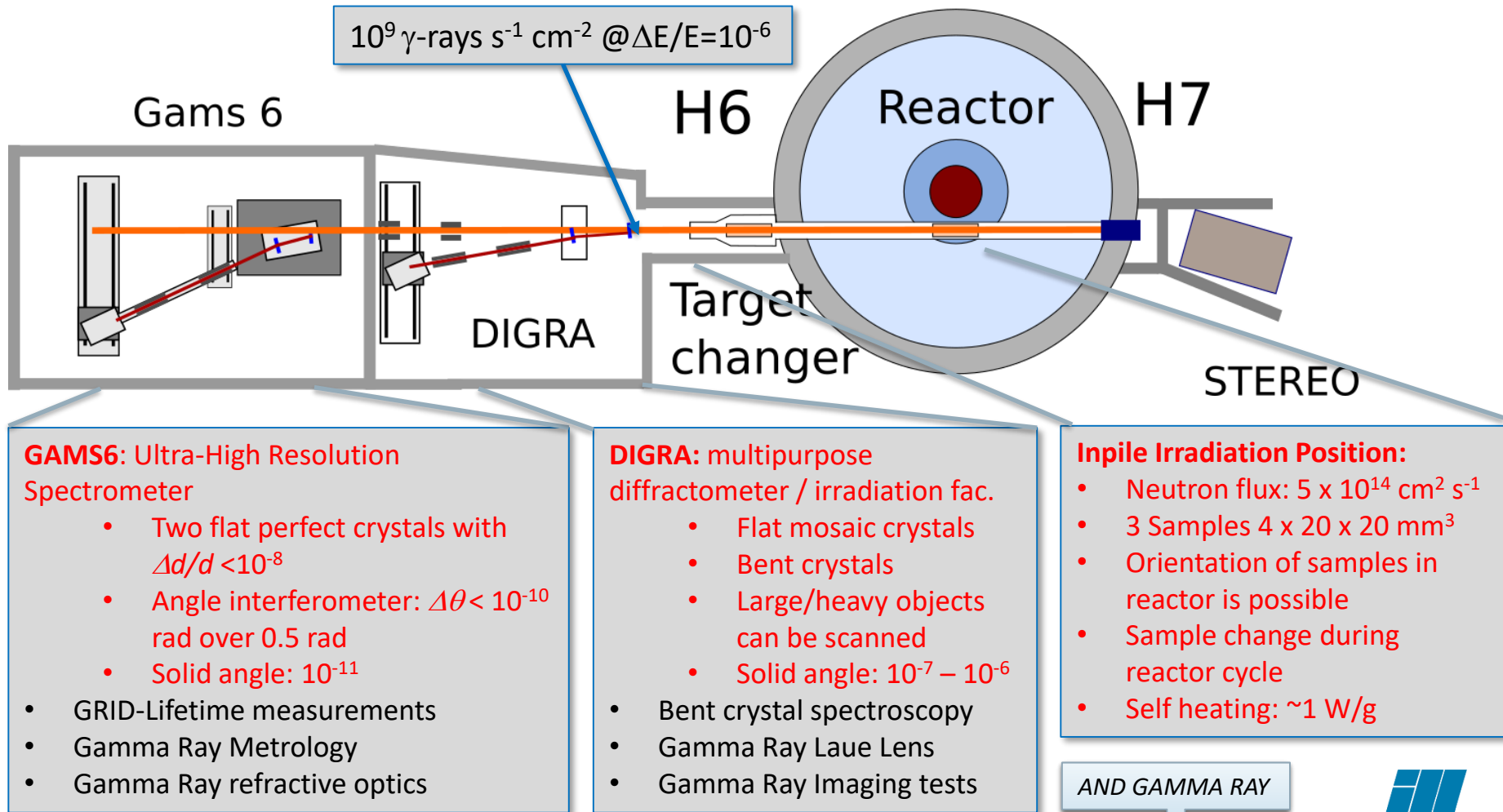
• “Right” fission fragment:
→ mass identification with a Gas-Filled magnet for filtering

Ancillary detectors:

- neutron detectors for fission studies
- + → LaBr₃(Ce) for short lifetime (10 ps → 1 ns)
- low energy Ge detectors
- ...

PN3: The high resolution gamma ray facility

General Layout and Parameters (PN3 since end 2014)

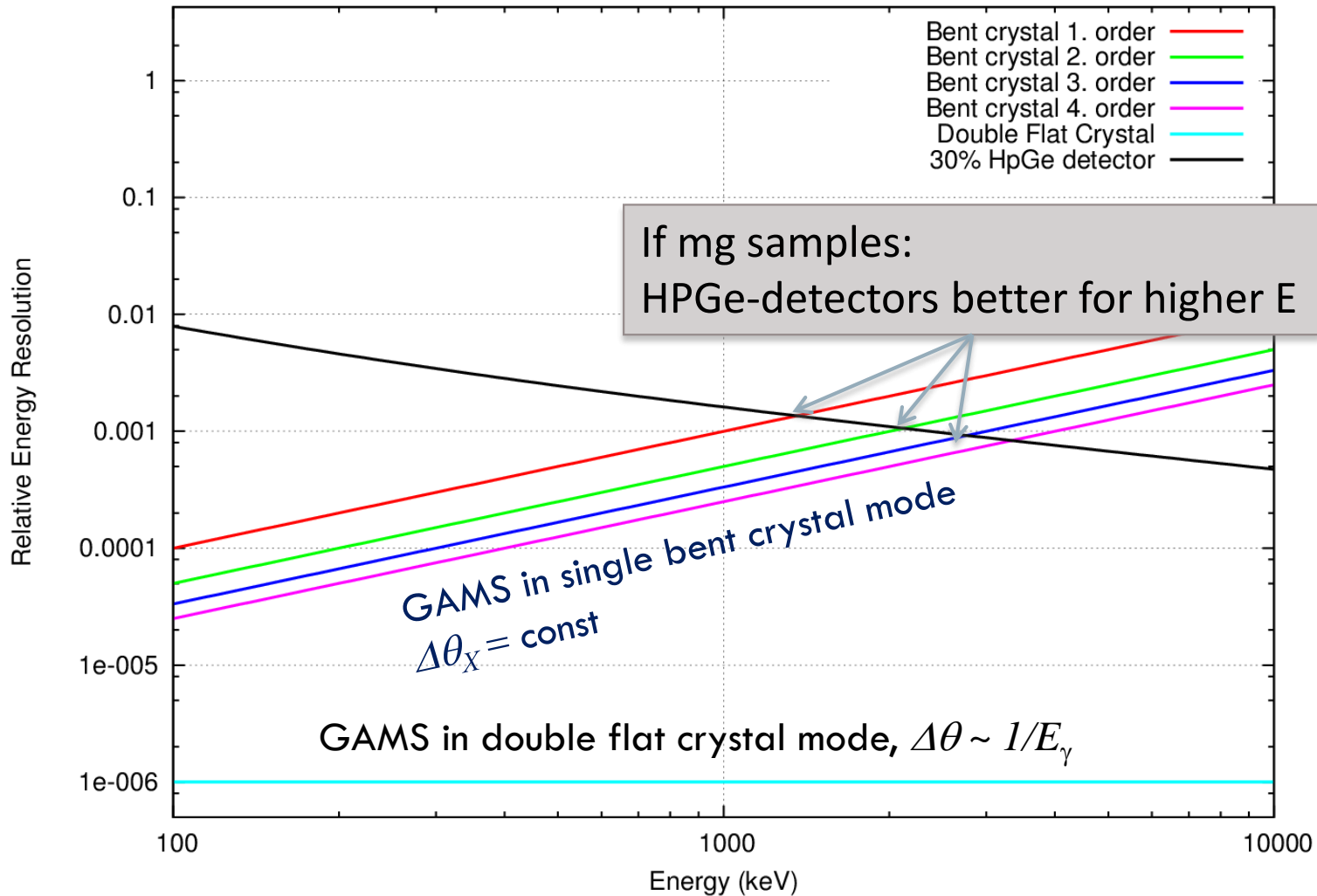


THE EUROPEAN NEUTRON SOURCE



NEUTRONS
FOR SOCIETY

Overview on Energy resolution

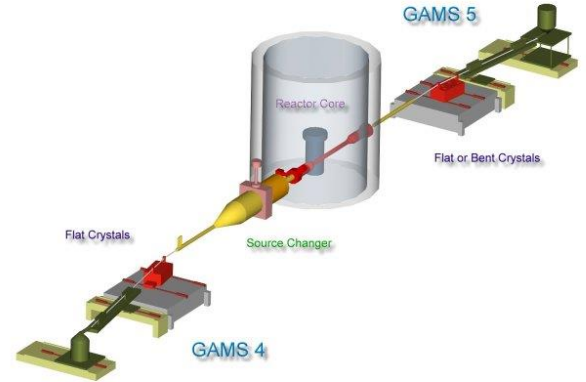
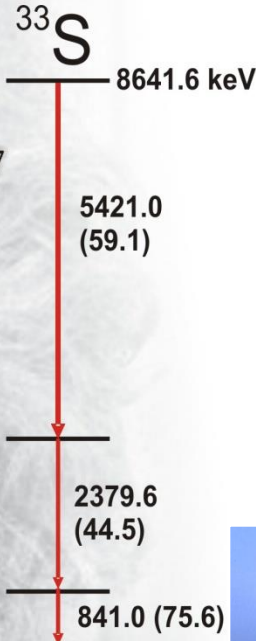


Direct test of mass/energy relationship $E = mc^2$

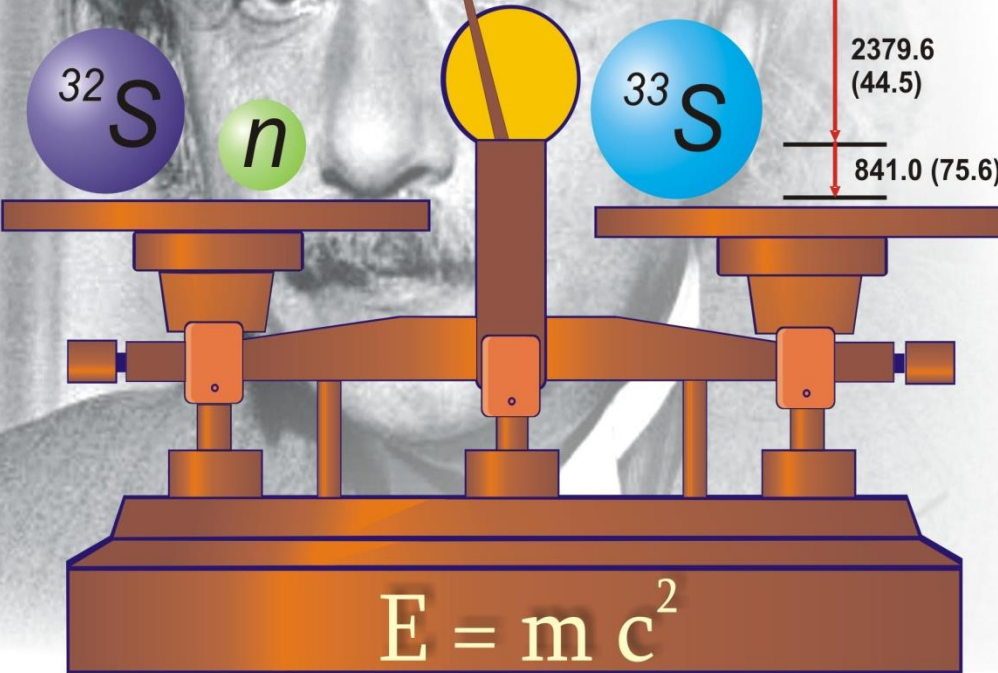
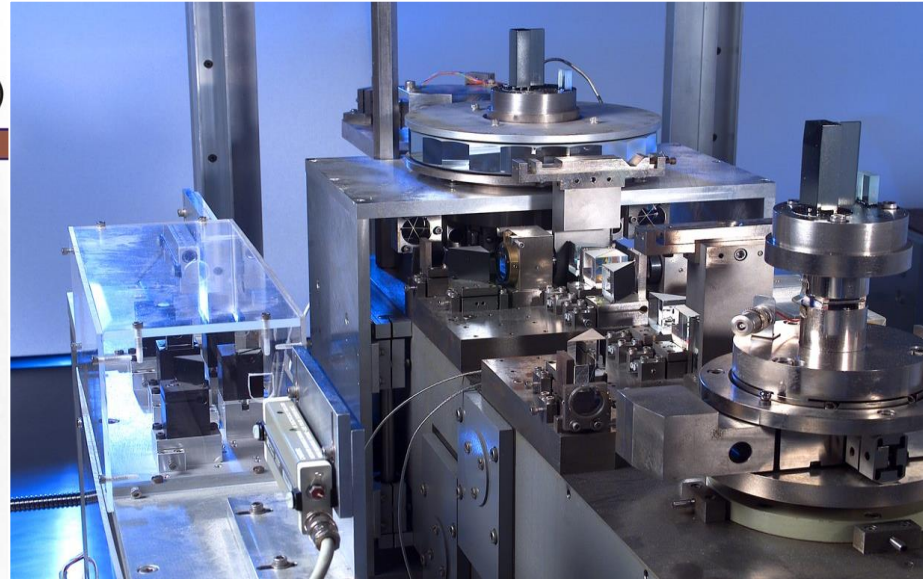
2005 World Year of Physics - Einstein's "Miraculous year" 1905

ILL-MIT-NIST, Nature 430, 58 (2005)

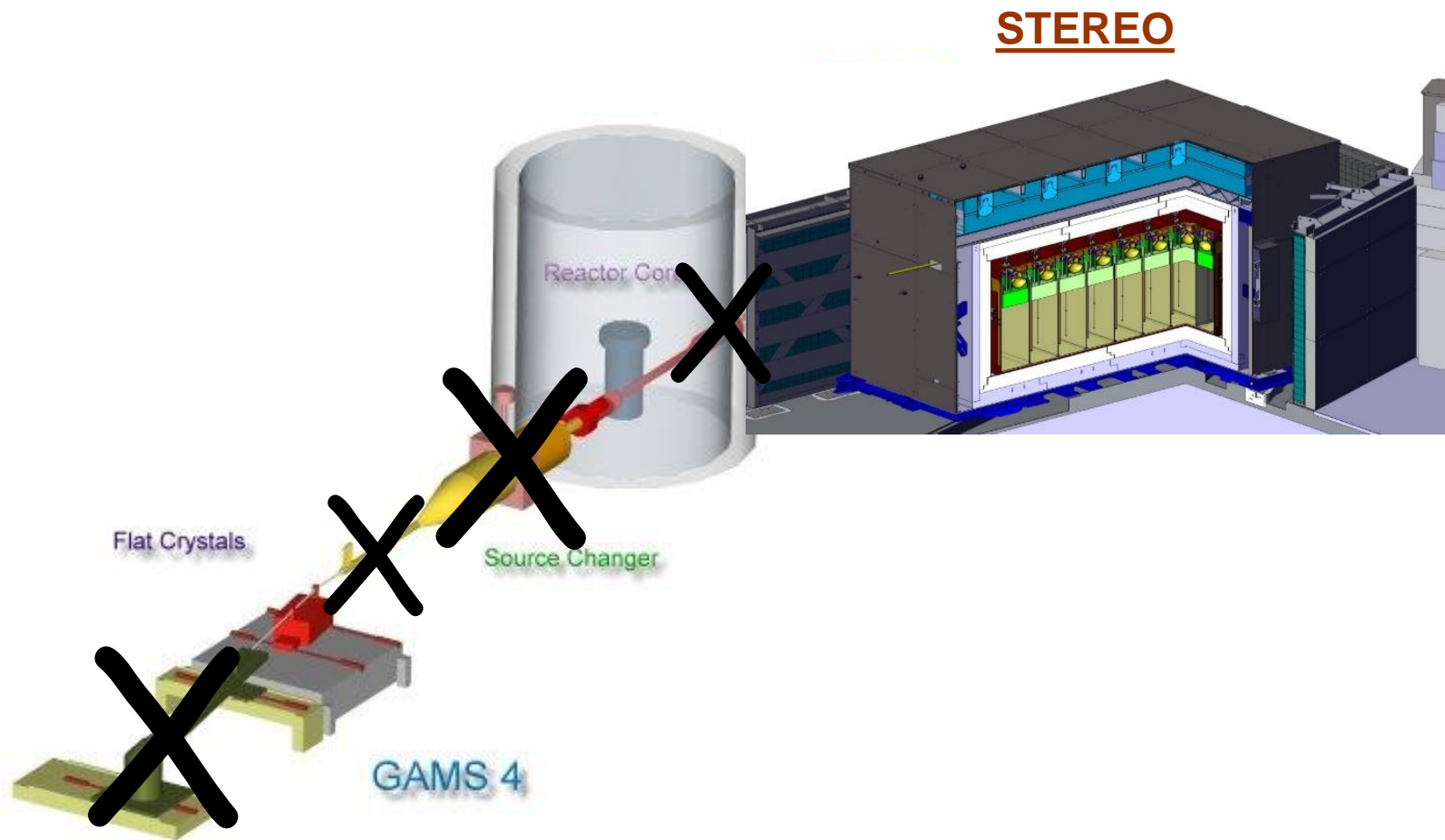
$$1 - \Delta mc^2 / E = 1.4(4.4) \cdot 10^{-7}$$



GAMS Interferometer ILL

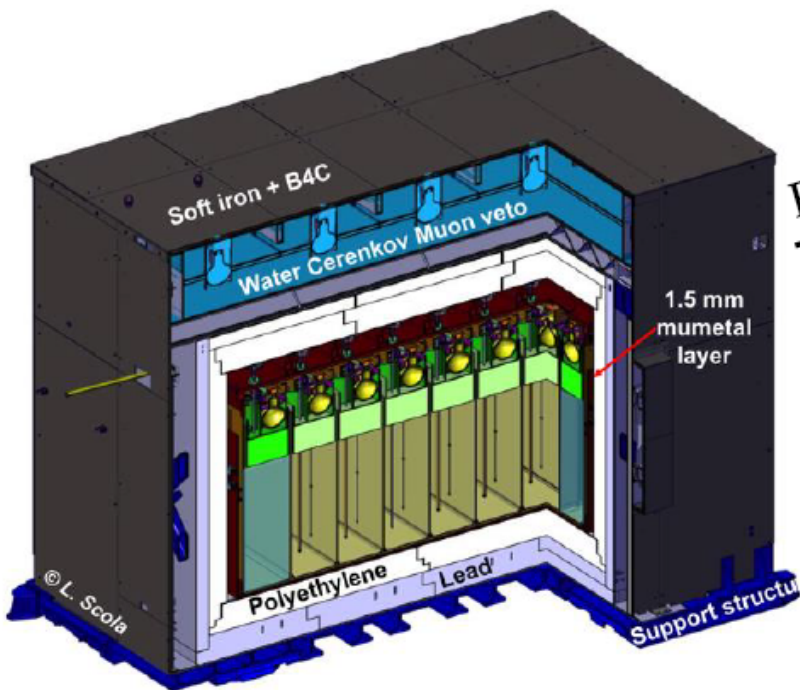


Antineutrino detector **STEREO** at the sealed end of beam tube H7



In April 2017 through beam tube H6/H7 removed and sealed

STEREO Detector



Reactor (~10 m)

20 cm thick acrylic buffers for homogeneous detector response. PMT coupling via oil bath.

Liquid Scintillator

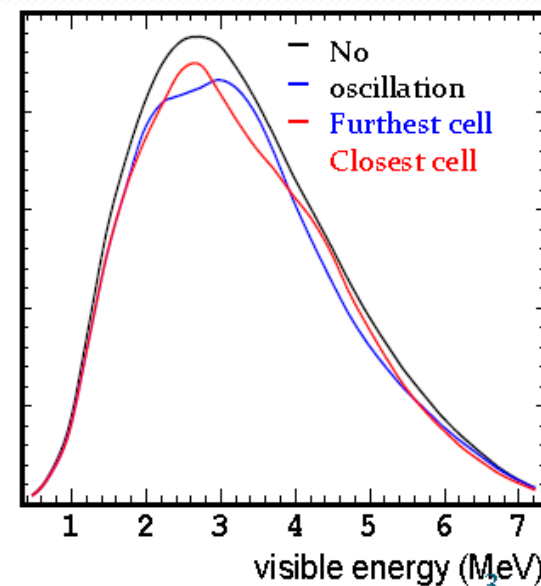
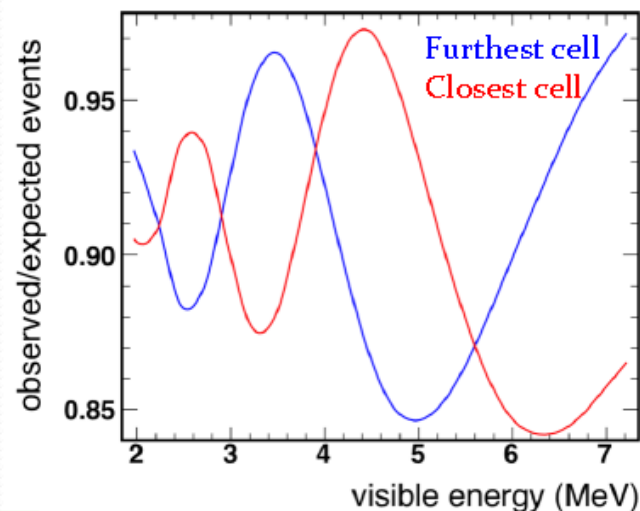
Target: 6 identical cells

- Gd-loaded (0.2% in mass)
- $V_{\text{tot}} = 2.2 \times 0.9 \times 0.9 \text{ m}^3$

Gamma catcher (unloaded):

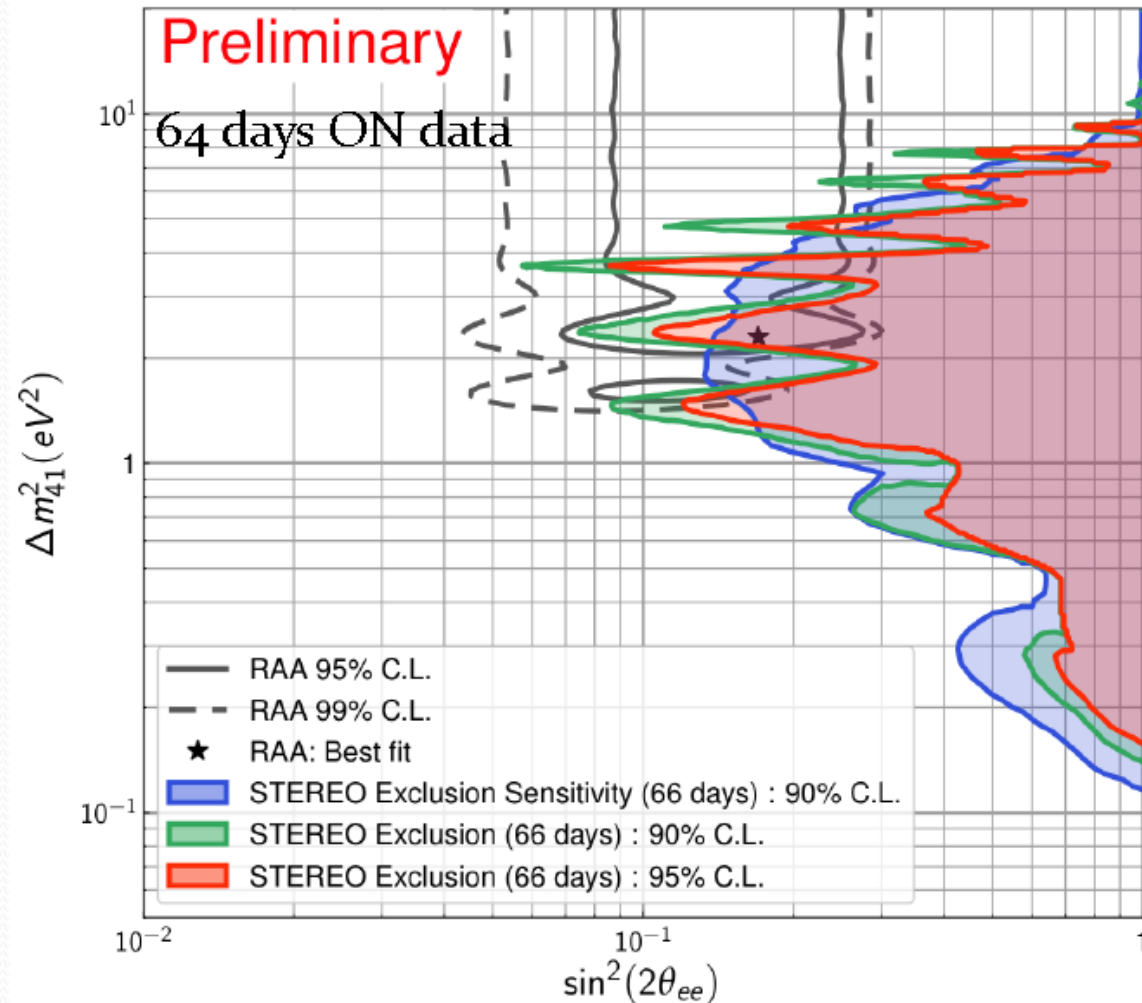
- Vetos ext. background
- Captures escaping γ 's

$$\sin^2(2\theta) = 0.14, \Delta m^2 = 2.4 \text{ eV}^2$$



STEREO exclusion contours

- Raster scan approach.
- $\Delta\chi^2$ law simulated in each Δm^2 bin.
- Reject oscillation amplitudes larger than statistical fluctuations for a given C.L.

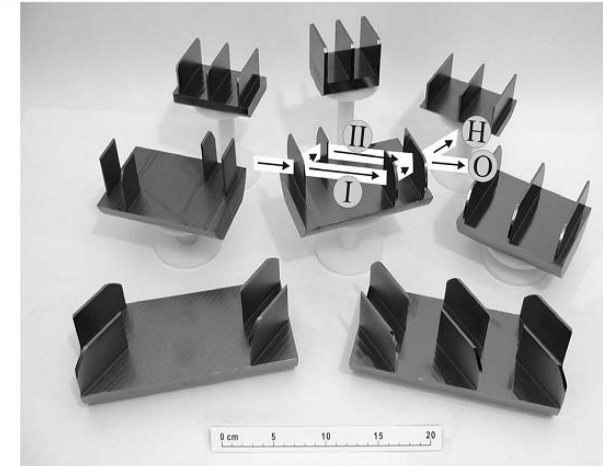
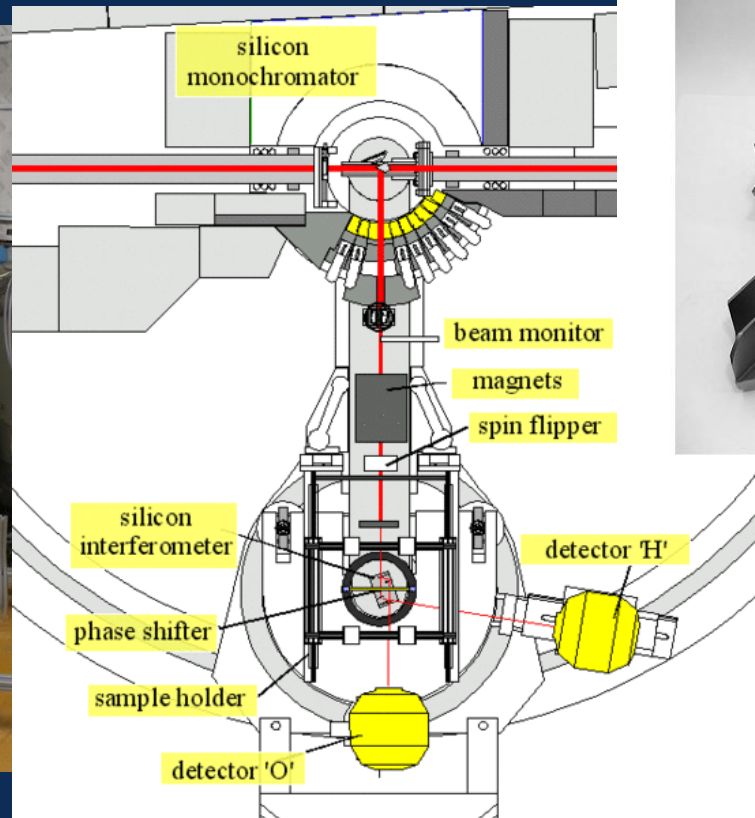
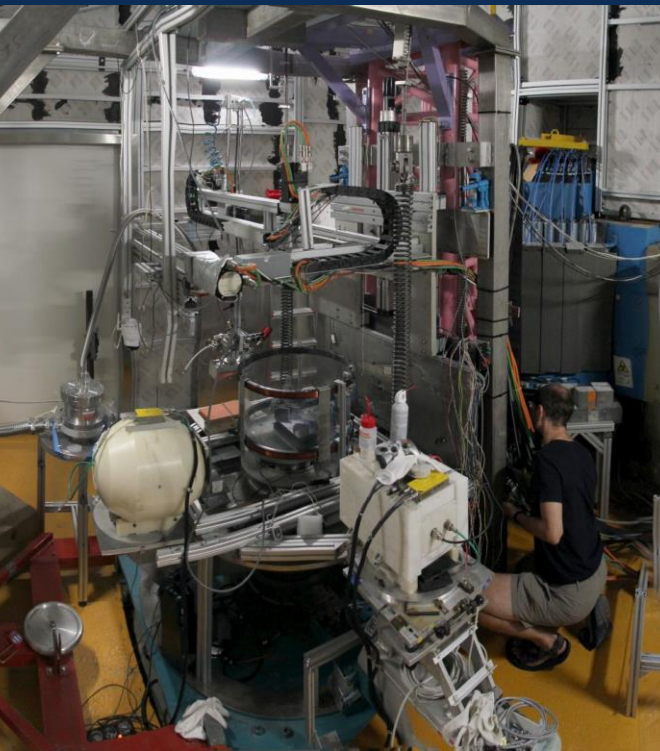


→ Reject at 98.8% C.L. the RAA oscillation best fit value
→ Contour sensitivity will be extended with more statistic

S18 - CRG instrument (Atominstitut, TU Vienna, Austria [H. Lemmel])

interferometer (perfect Si crystals) for basic neutron quantum optics, fundamental tests of quantum physics, neutron scattering lengths and USANS (ultra-small angle neutron scattering)

Neutron interferometer family



Observation of a quantum Cheshire Cat in a matter wave interferometer experiment

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(Dated: December 16, 2013)

From its very beginning quantum theory has been revealing extraordinary and counter-intuitive phenomena, such as wave-particle duality [1], Schrödinger cats [2] and quantum non-locality [3–6]. In the study of quantum measurement, a process involving pre- and postselection of quantum ensembles in combination with a weak interaction was found to yield unexpected outcomes [7]. This scheme, usually referred to as "weak measurements", can not only be used as an amplification technique [8–10] and for minimal disturbing measurements [11, 12], but also for the exploration of quantum paradoxes [13–17]. Recently the quantum Cheshire Cat has attracted attention [18–20]: a quantum system can behave as if a particle and its property (e.g. its polarization) are spatially separated. Up to now most

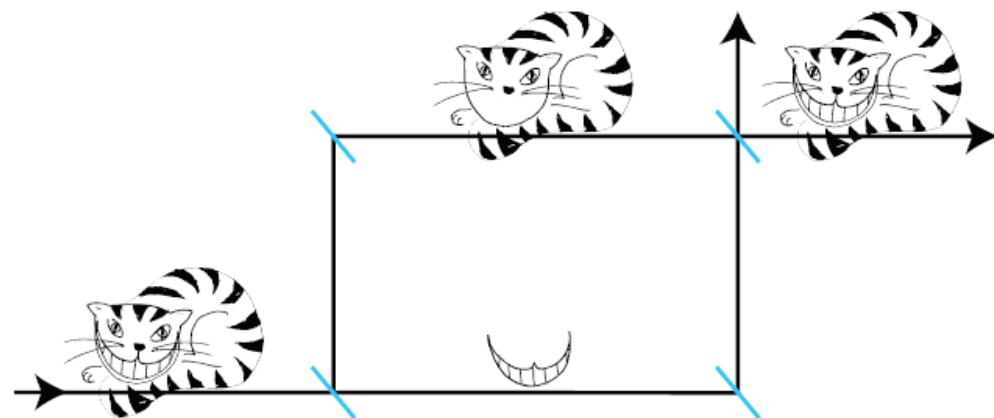


FIG. 1: Artistic depiction of the quantum Cheshire Cat: Inside the interferometer the Cat goes through the upper beam path, while its grin travels along the lower beam path.

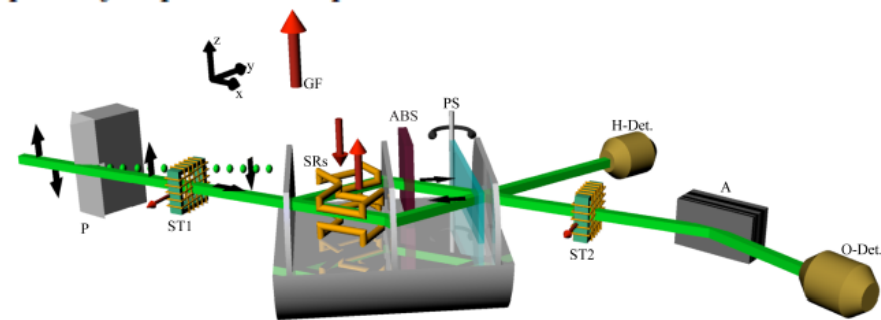


FIG. 2: Illustration of the experimental setup for the observation of a quantum Cheshire Cat in a neutron interferometer: The neutron beam is polarized by passing through magnetic birefringent prisms (P). To prevent depolarization, a magnetic guide field (GF) is applied around the whole setup. A spin turner (ST1) rotates the neutron spin by $\pi/2$. When entering the interferometer the neutron beam splits into two paths. Preselection of the system's wave function $|\psi_i\rangle$ is completed by two spin rotators (SRs) inside the neutron interferometer. These SRs are also used to perform the weak measurement of $\langle \hat{\sigma}_z \hat{\Pi}_{II} \rangle_w$ and $\langle \hat{\sigma}_z \hat{\Pi}_{II} \rangle_w$. The absorbers (ABS) are inserted in the beam paths when $\langle \hat{\Pi}_I \rangle_w$ and $\langle \hat{\Pi}_{II} \rangle_w$ are determined. The phase shifter (PS) makes it possible to tune the relative phase χ between the beams in path I and path II. The two outgoing beams of the interferometer are monitored by the H- and O-detector in reflected and forward directions, respectively. Only the neutrons reaching the O-detector are affected by postselection using a spin turner (ST2) and a spin analyzer (A).

Properties of UCN

Ultracold neutrons, that is, neutrons whose energy is so low that they can be contained for long periods of time in material and magnetic bottles

$$E_{kin} (\sim 5 \text{ ms}^{-1}) = 100 \text{ neV} (10^{-7} \text{ eV})$$

$$\lambda_{UCN} \sim 1000 \text{ \AA}$$

$$T_{UCN} \sim 2 \text{ mK}$$

UCN are totally reflected from suitable materials at **any angle of incidence**, hence **storable!**

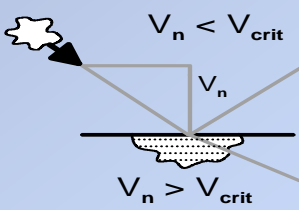
Long storage and observation times possible (up to several minutes!)

High precision measurements of the properties of the free neutron (lifetime, electric dipole moment, gravitational levels, ...)

Interaction with matter:
UCN see a *Fermi-Potential* E_F

$E_F \sim 10^{-7} \text{ eV}$ for many materials, e.g.

- beryllium 252 neV
- stainless steel 200 neV



UCN are furthermore storable by gravity and/or magnetic fields

Fermi potential	$\sim 10^{-7} \text{ eV}$
Gravity $\Delta E = m_n g \Delta h$	$\sim 100 \text{ neV / Meter}$
Magnetic field $\Delta E = \mu_n B$	$\sim 60 \text{ neV / Tesla}$

V. I. Lushchikov, Yu. N. Pokotilovskii, A. V. Strelkov, and F. L. Shu
 Joint Institute for Nuclear Research
 Submitted 18 November 1968
 ZhETF Pis. Red. 9, No. 1, 40 - 45 (5 January 1969)

Ya. B. Zel'dovich showed in 1959 [1] that neutrons with velocities v experience total reflection from the walls at all incidence angles, can be cavity. As was noted recently [2], the idea of storing neutrons points to the accuracy of measurement of the neutron dipole moment, an important fac of the neutron. This has been undertaken to check experimentally the

... by extracting neutrons from the low energy tail of the distribution in the source

the tube. The neutron detectors 11 and 12 were FEU-13 photomultipliers of

MEASUREMENTS OF TOTAL CROSS SECTIONS FOR VERY SLOW NEUTRONS WITH VELOCITIES FROM 100 m/sec TO 5 m/sec

A. STEYERL

Physik-Department, Technische Hochschule München, Munich, Germany

Received 24 February 1969

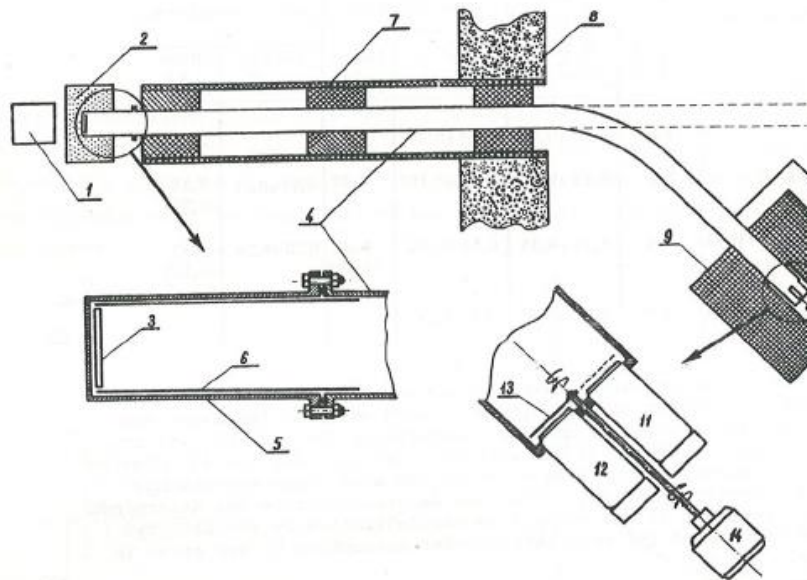


Fig. 1. Diagram of setup. 1- IBR reactor; 2, 3 - moderator (2 - paraffin layer 1 mm thick); 4 - copper tube, 9.4 cm i.d., total length 10.5 m; 5 - copper-foil cylinder; 7 - shield (paraffin with boron carbide); 8 - 2-m c actor chamber; 9 - detector shield (paraffin); 10 - tube filling and evac 12 - detectors (FEU-13 with layers of ZnS or ZnS + Li compound); 13 - cog between shutter and detector < 1 mm); 14 - shutter mechanism; 15 - trap 1

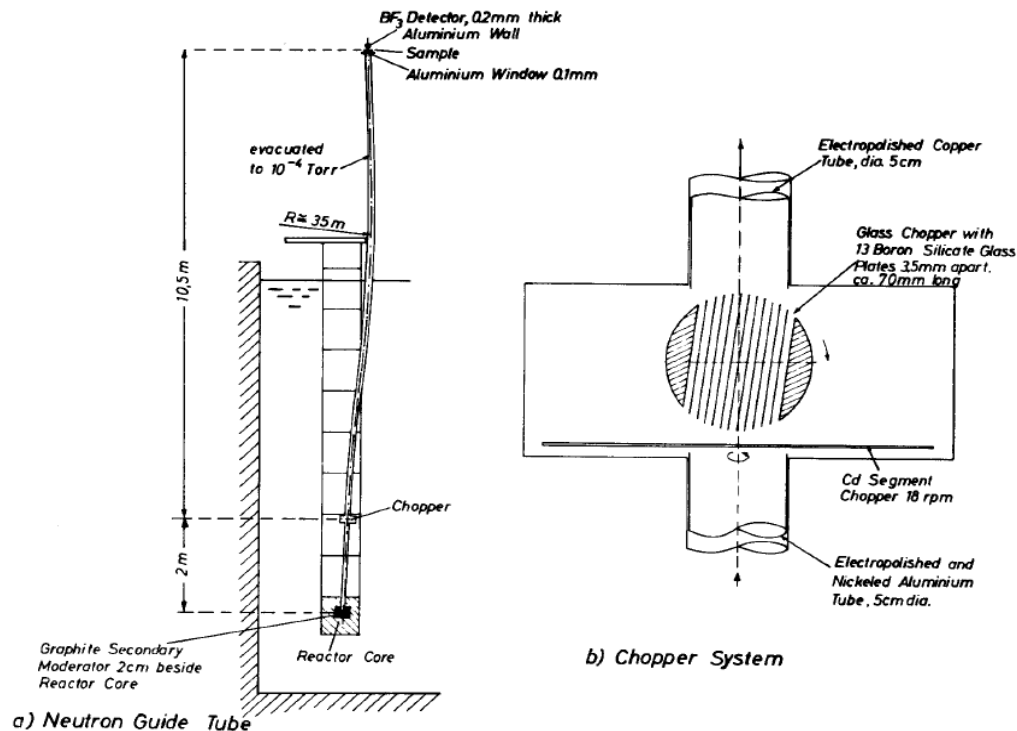
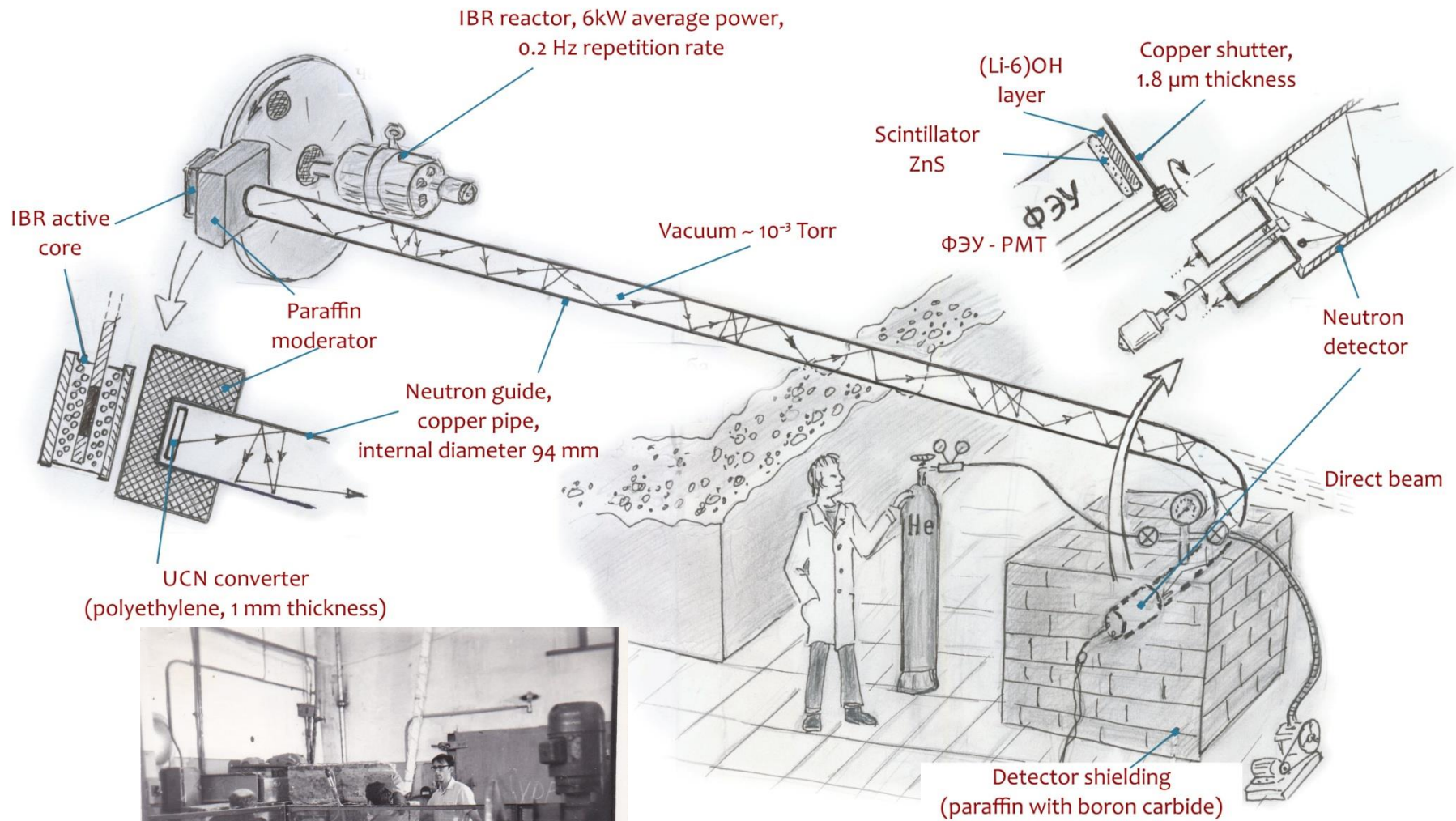


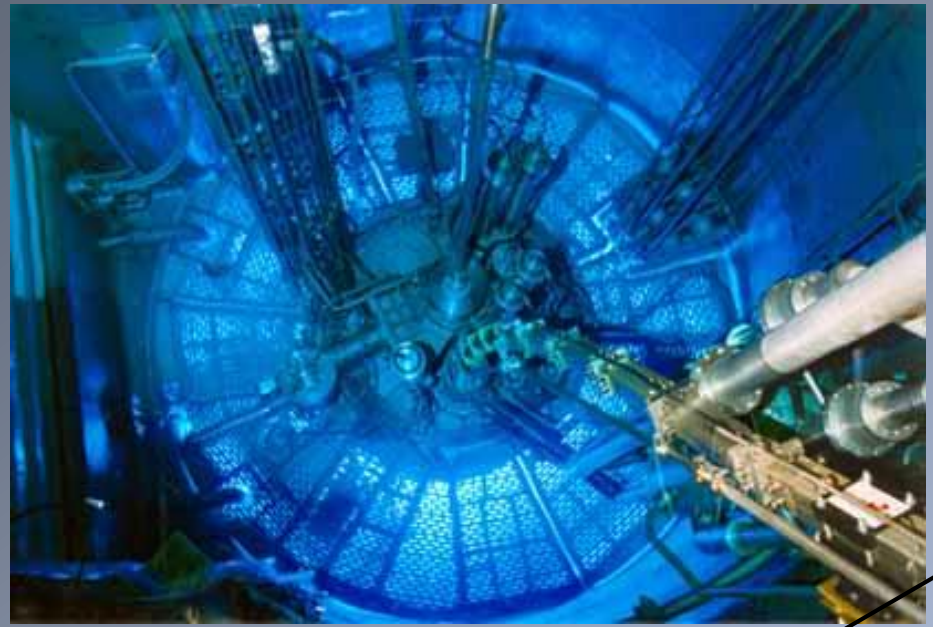
Fig. 1. Vertical beam tube for very slow neutrons.

how UCN were "really" discovered in Dubna

drawing courtesy of A.V. Strelkov



The UCN/VCN facility PF2

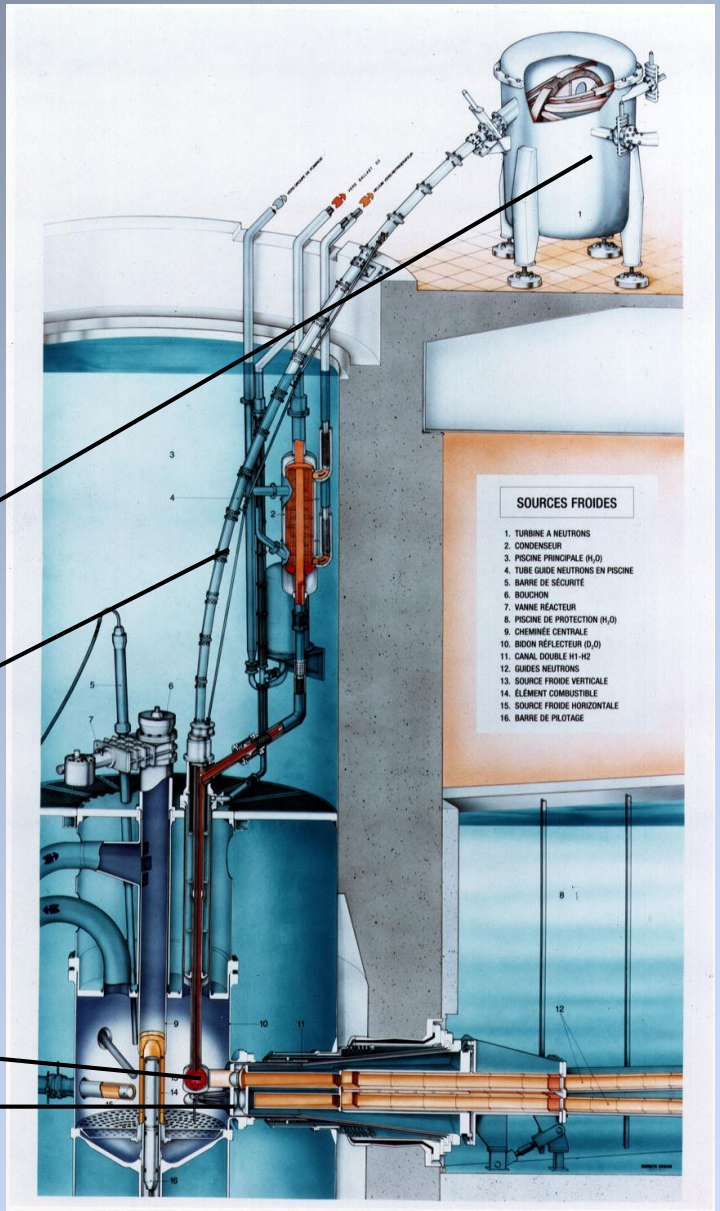


Neutron turbine
A. Steyerl (TUM - 1985)

Vertical guide tube

Cold source

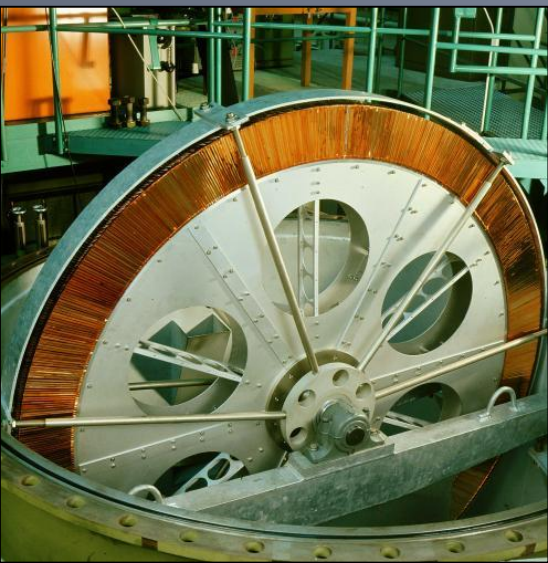
Reactor core



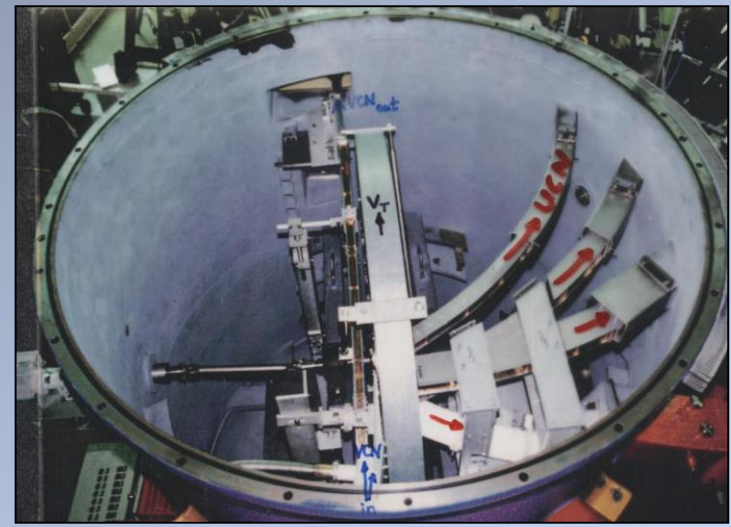
A. Steyerl et al., Phys. Lett. A116 (1986) 347

Generating Ultracold Neutrons (UCN)

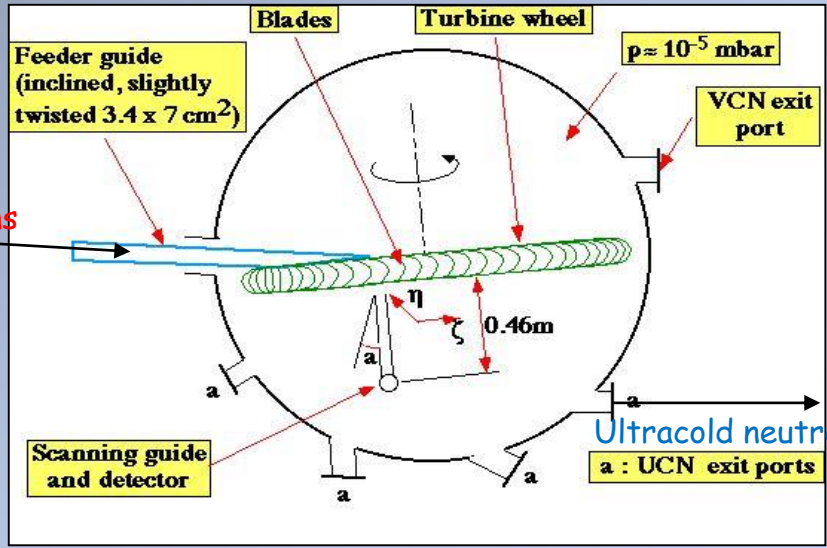
"Steyerl turbine" Doppler shifting device



Steyerl turbine at FRM-I (Munich)



Steyerl turbine (2nd generation) at PF2 / ILL 10 years later



The total UCN current density is $2.6 \times 10^4 \text{ cm}^{-2} \text{ s}^{-1}$ up to $v_z = 6.2 \text{ m/s}$ and $3.3 \times 10^4 \text{ cm}^{-2} \text{ s}^{-1}$ up to $v_z = 7 \text{ m/s}$. The total UCN current amounts to more than a million UCNs/s. Furthermore, we deduce from the TOF data special UCN densities of 87 cm^{-3} (for $v_z < 6.2 \text{ m/s}$) and 110 cm^{-3} for $v_z < 7 \text{ m/s}$.

In a storage bottle experiment 36 UCNs per cm^{-3} (for $v_z < 6.2 \text{ m/s}$) were detected!

A. Steyerl et al.,
Physics Letters A 116 (1986) 347 - 352

The PF2 beam facility



NE
FOR



PF2: Physique Fondamentale 2
2nd installation for fundamental physics

4 positions for Ultracold Neutrons (UCN)

was :

$$v = 5 \text{ ms}^{-1}$$

$$\rho = \sim 50 \text{ cm}^{-3} \text{ (at the experiment)}$$

is :

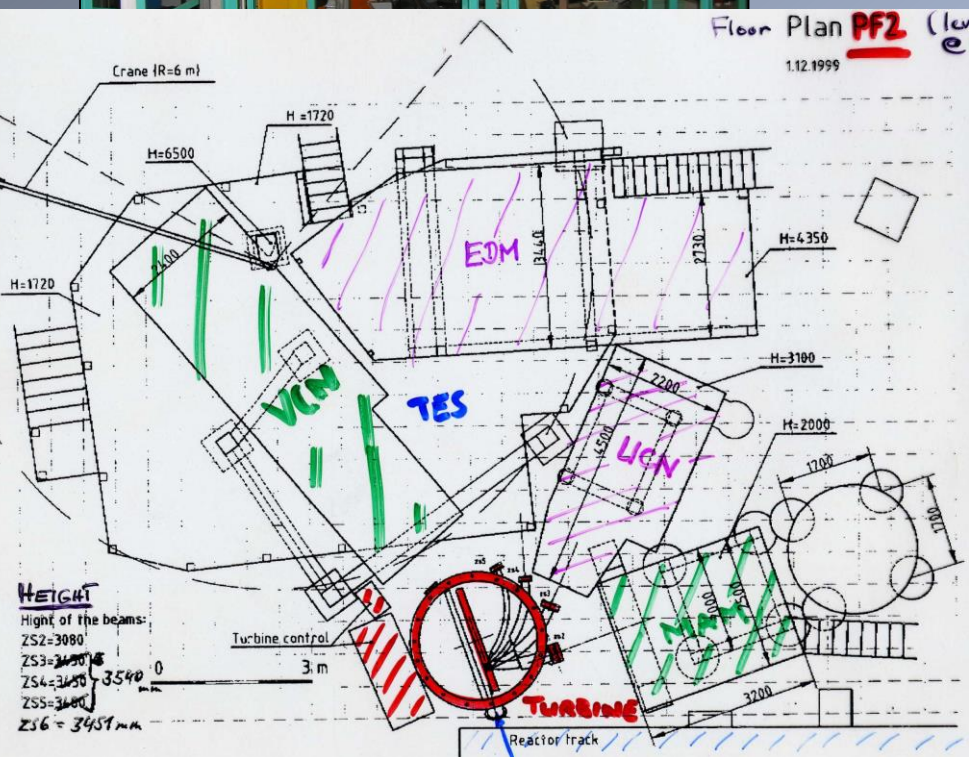
$$v \leq 7 \text{ ms}^{-1}$$

$$\rho = \sim 20 \text{ cm}^{-3} \text{ (at the experiment)}$$

- MAM
- EDM
- UCN
- TES

1 position for Very Cold Neutrons (VCN)

- VCN beam
- $$v = 50 \text{ ms}^{-1}$$
- $$\Phi = 10^8 \text{ cm}^{-2} \text{ s}^{-1}$$



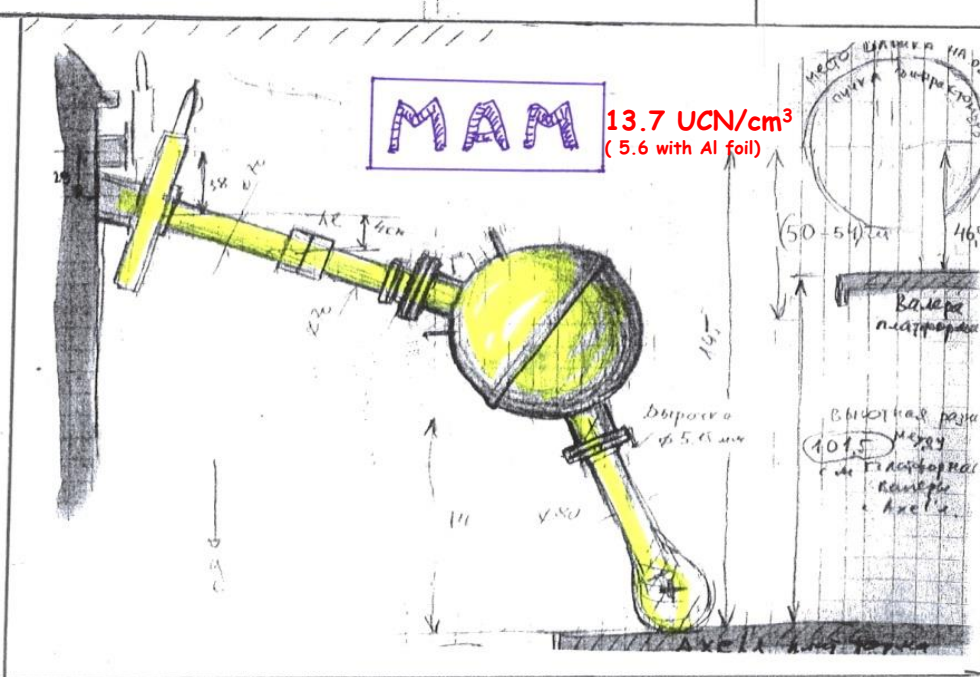
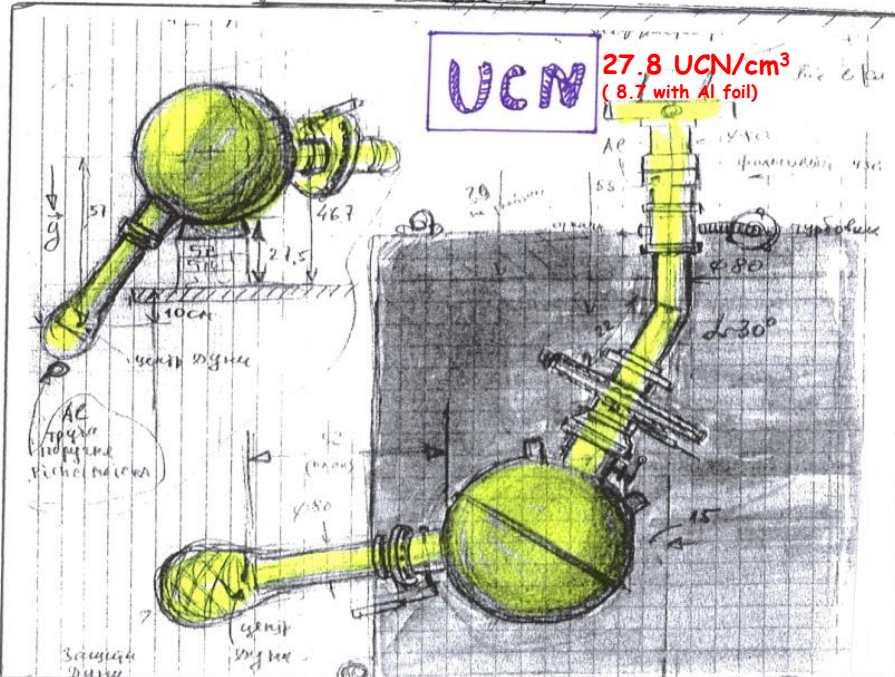
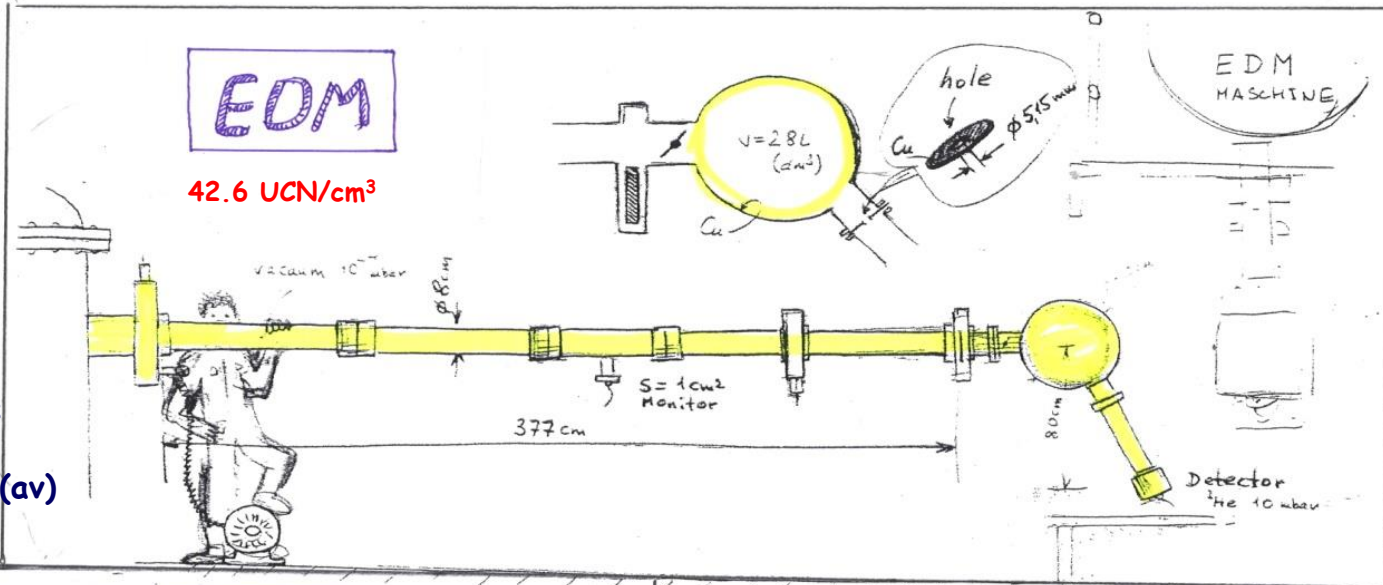
W. Drexel, Neutron News 1 (1990) 23

UCN densities (Cu sphere as UCN container)

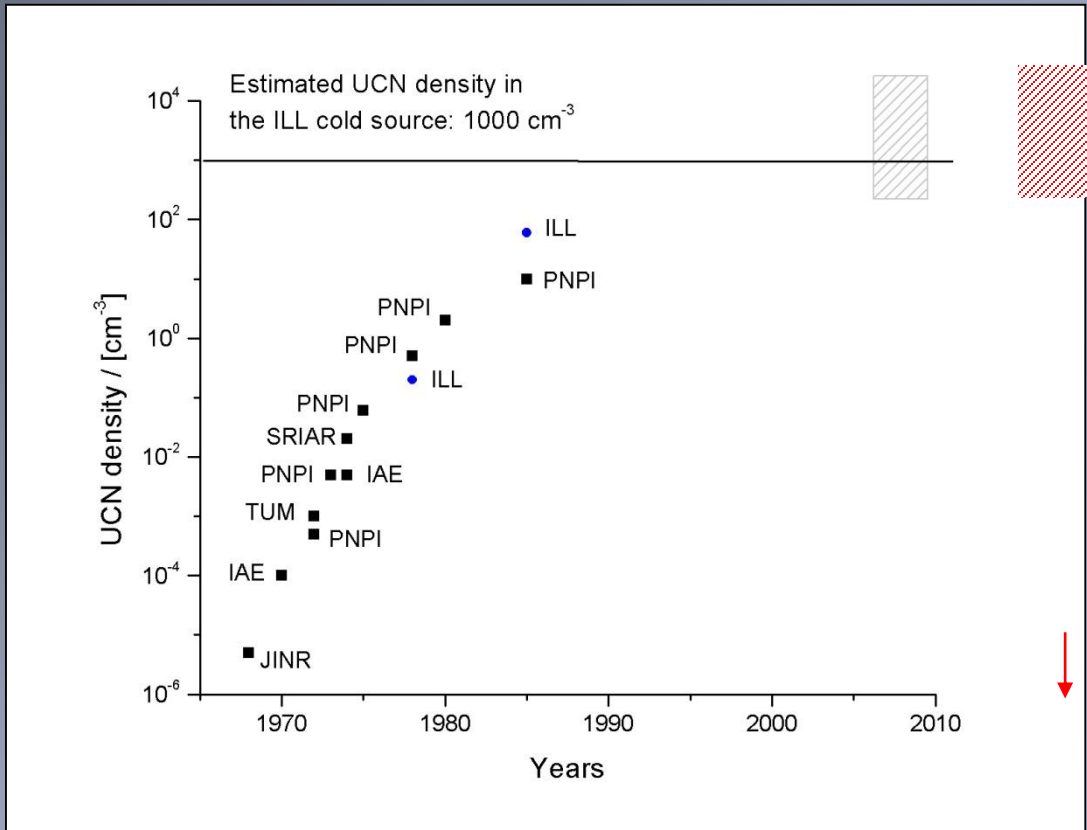
drawing (log book September 1999 on UCN flux measurements at different beam positions of PF2) courtesy of A. Strelkov

measured using
"pin hole" method:

$$\text{flux} = 1/4 * \text{dens} * \text{vel}(\text{av})$$



UCN facilities - Status and Future



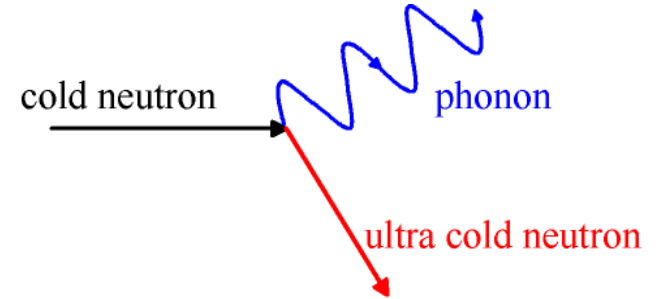
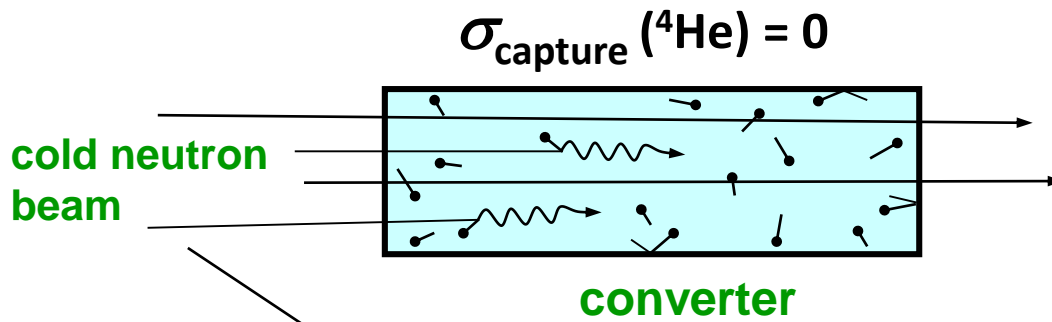
More and stronger UCN facilities in the future worldwide

- **PSI** (CH)
- **Mainz / Munich** (D)
- **ILL** (F)
- **LANL / NIST / SNS / NCSU** (USA)
- **RGNP** (J) now **TRIUMF** (Canada)
- **JPARC** (J)
- **PNPI** (RUS)

Reactor tank and pool are very close



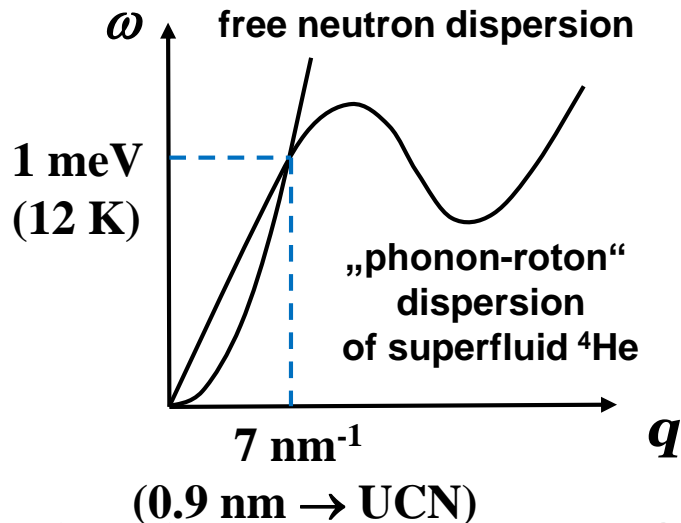
UCN production in He-II



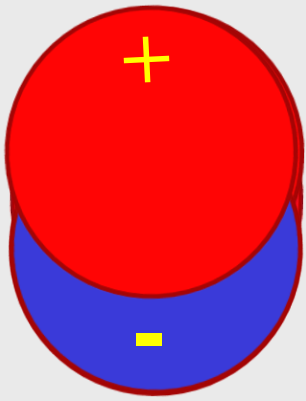
$$\rho_{\text{UCN}} = P\tau$$

$$\tau^{-1} = \tau^{-1}_{\text{decay}} + \tau^{-1}_{\text{upscattering}} + \tau^{-1}_{\text{capture}} + \tau^{-1}_{\text{wall losses}}$$

T [K]	τ_{max} [s]
1	100
0.8	310
0.7	510
0.5	820
0	880

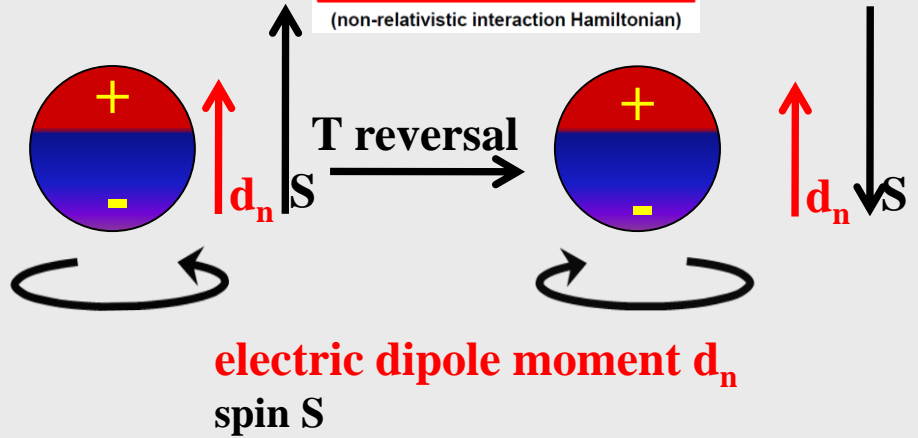


\rightarrow need $T < 0.5 - 0.6 \text{ K}$
and low-loss walls



$$\mathcal{H} = -\mu \cdot \frac{\vec{S}}{|\vec{S}|} \cdot \vec{B} - d \cdot \frac{\vec{S}}{|\vec{S}|} \cdot \vec{E}$$

(non-relativistic interaction Hamiltonian)



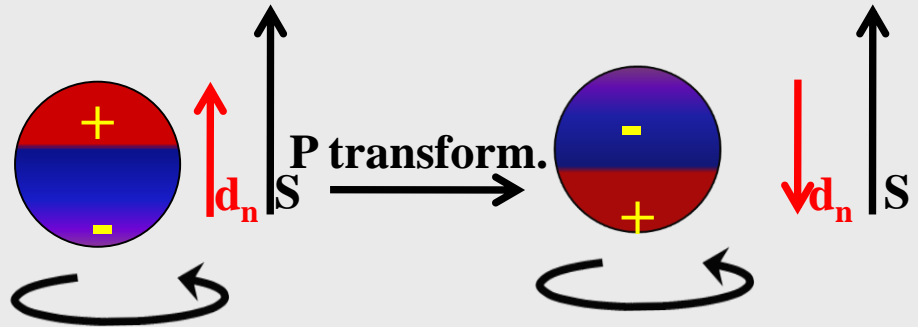
Electric Dipole Moment:

neutron is electrically neutral

If average positions of positive and negative charges do not coincide:



EDM d_n



CPT conservation \rightarrow CP violation

CP violation in Standard Model generates very small neutron EDM
Beyond the Standard Model contributions tend to be much bigger

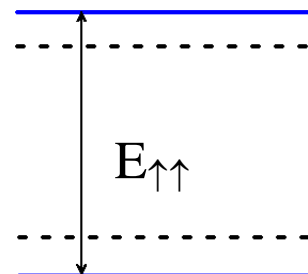
neutron a very good system to look for CP violation beyond the Standard Model

E.M. Purcell and N.F. Ramsey
Phys. Rev. 78, 807 (1950)

Experiments:

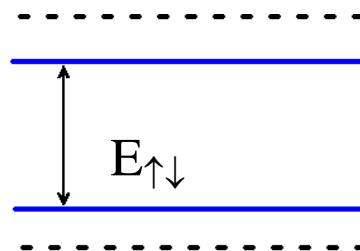
Measurement of Larmor precession frequency of polarised neutrons in a magnetic & electric field

Compare the precession frequency for parallel fields:



$$\nu_{\uparrow\uparrow} = E_{\uparrow\uparrow}/h = [-2B_0\mu_n - 2Ed_n]/h$$

to the precession frequency for anti-parallel fields



$$\nu_{\uparrow\downarrow} = E_{\uparrow\downarrow}/h = [-2B_0\mu_n + 2Ed_n]/h$$

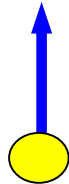
Need to measure change in Larmor precession frequency to a very high degree : $< 1\mu\text{Hz}$
 < 1 turn per month!

The difference is proportional to d_n and E :

$$h(\nu_{\uparrow\uparrow} - \nu_{\uparrow\downarrow}) = 4E d_n$$

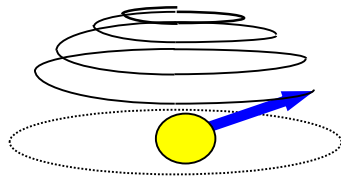
Ramsey method of Separated Oscillatory Fields

1.



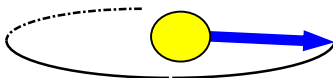
*"Spin up"
neutron...*

2.



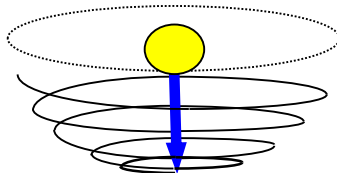
*Apply $\pi/2$
spin
flip pulse...*

3.

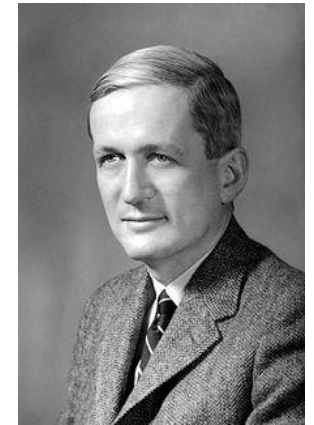
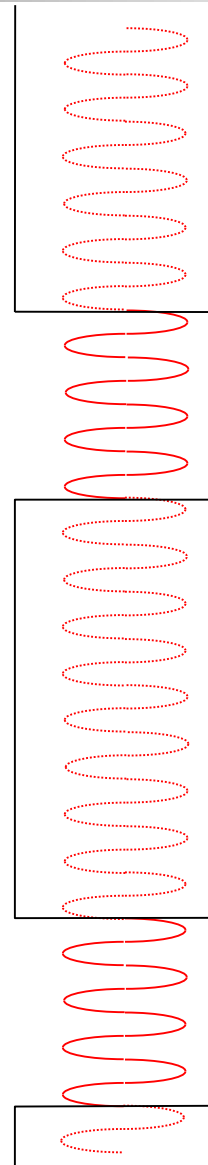


*Free
precession.
..*

4.



*Second $\pi/2$
spin
flip pulse.*



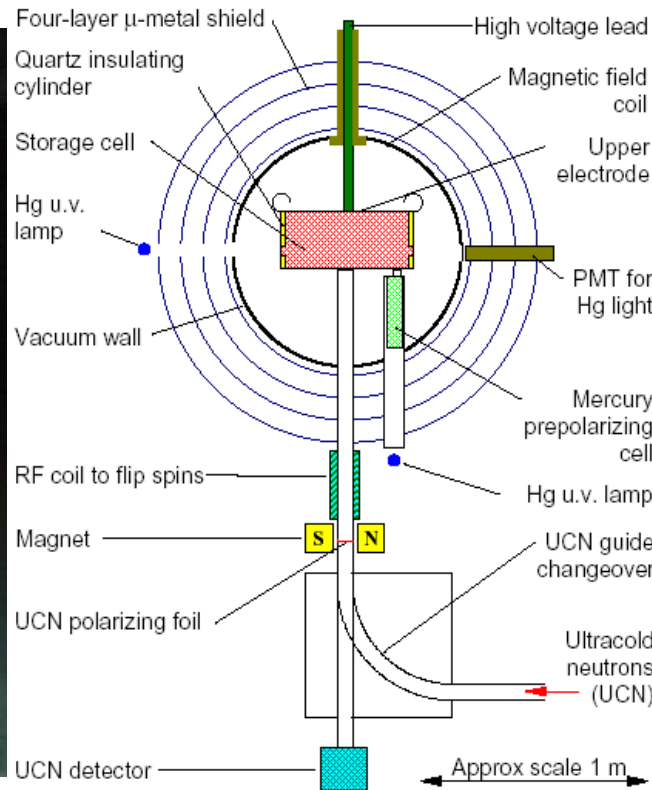
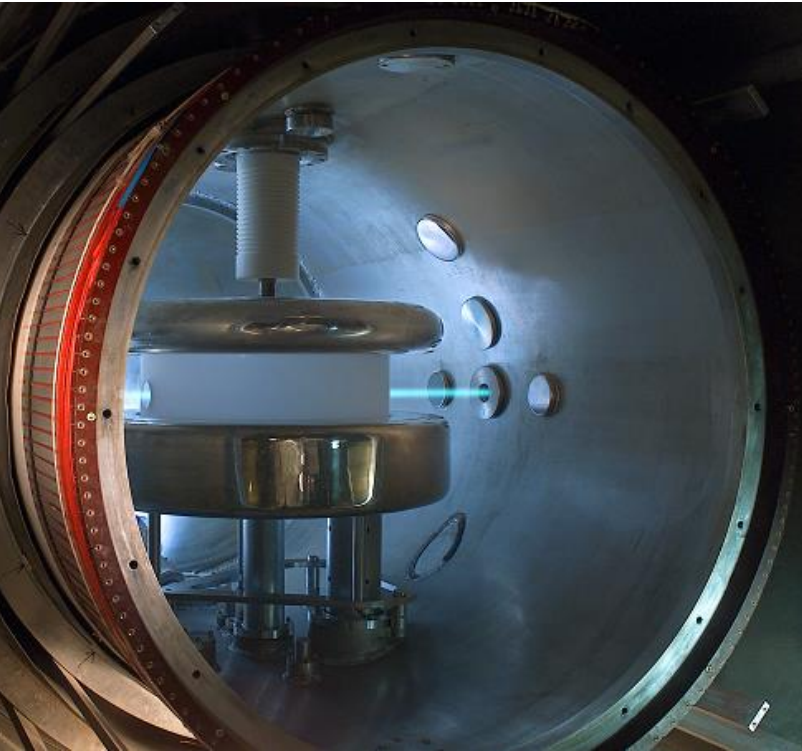
Norman F. Ramsey

1915 – 2011

Nobel Prize in Physics 1989

*"for the invention of the
separated oscillatory fields
method and its use in the
hydrogen maser and other
atomic clocks. ...*

Room Temperature Results



US University of Sussex

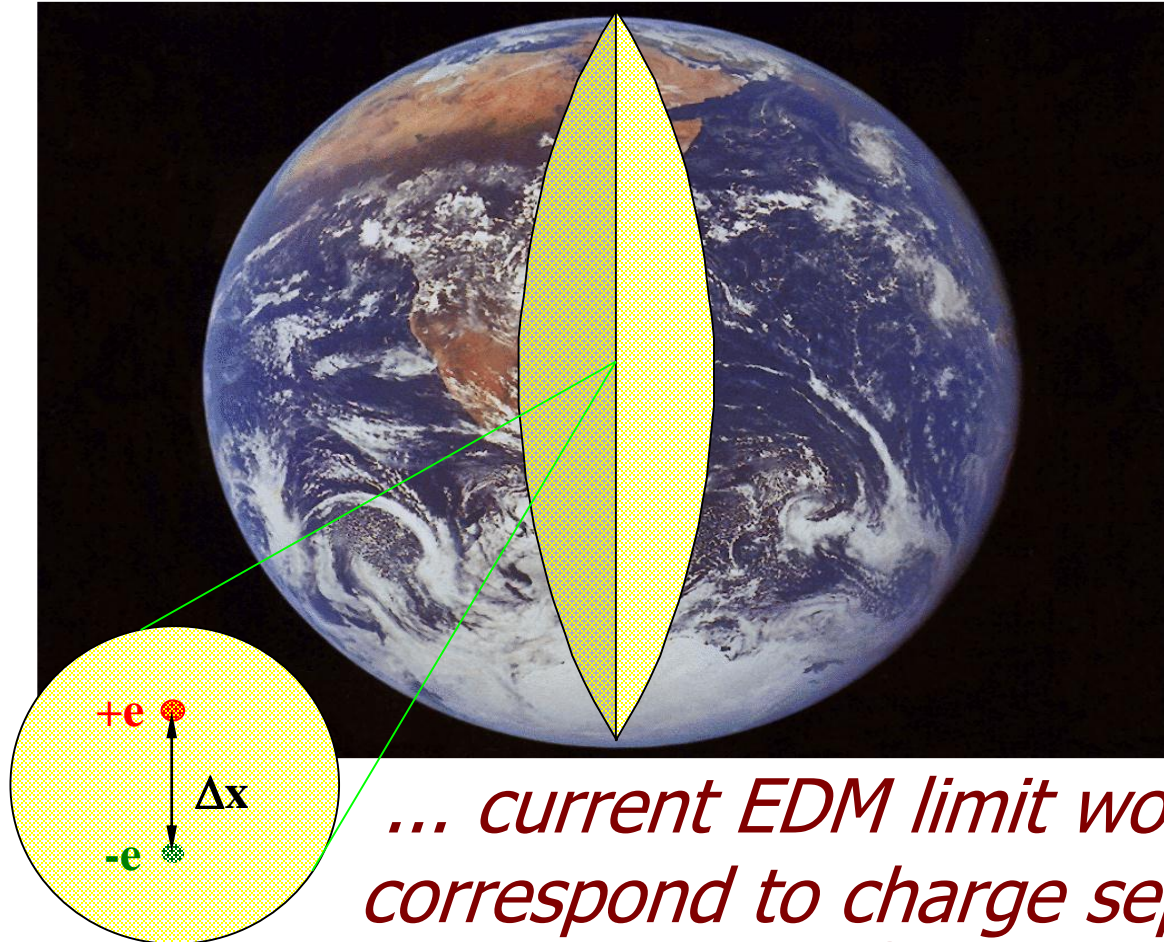


Room temperature neutron EDM result:
C.A. Baker et al., Phys. Rev. Lett. **97**, 131801 (2006)
 $|d_n| < 2.9 \times 10^{-26}$ e.cm (90% C.L.)

Reanalysis: J.M. Pendlebury et al., Phys. Rev. D **92**, 092003 (2015)
 $|d_n| < 3.0 \times 10^{-26}$ e.cm (90% C.L.)

Reality check

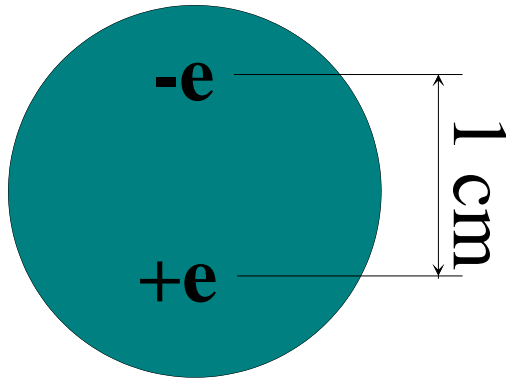
If neutron were the size of the Earth...



... current EDM limit would correspond to charge separation of
 $\Delta x \approx 3\mu$

The neutron EDM: exp. vs theory

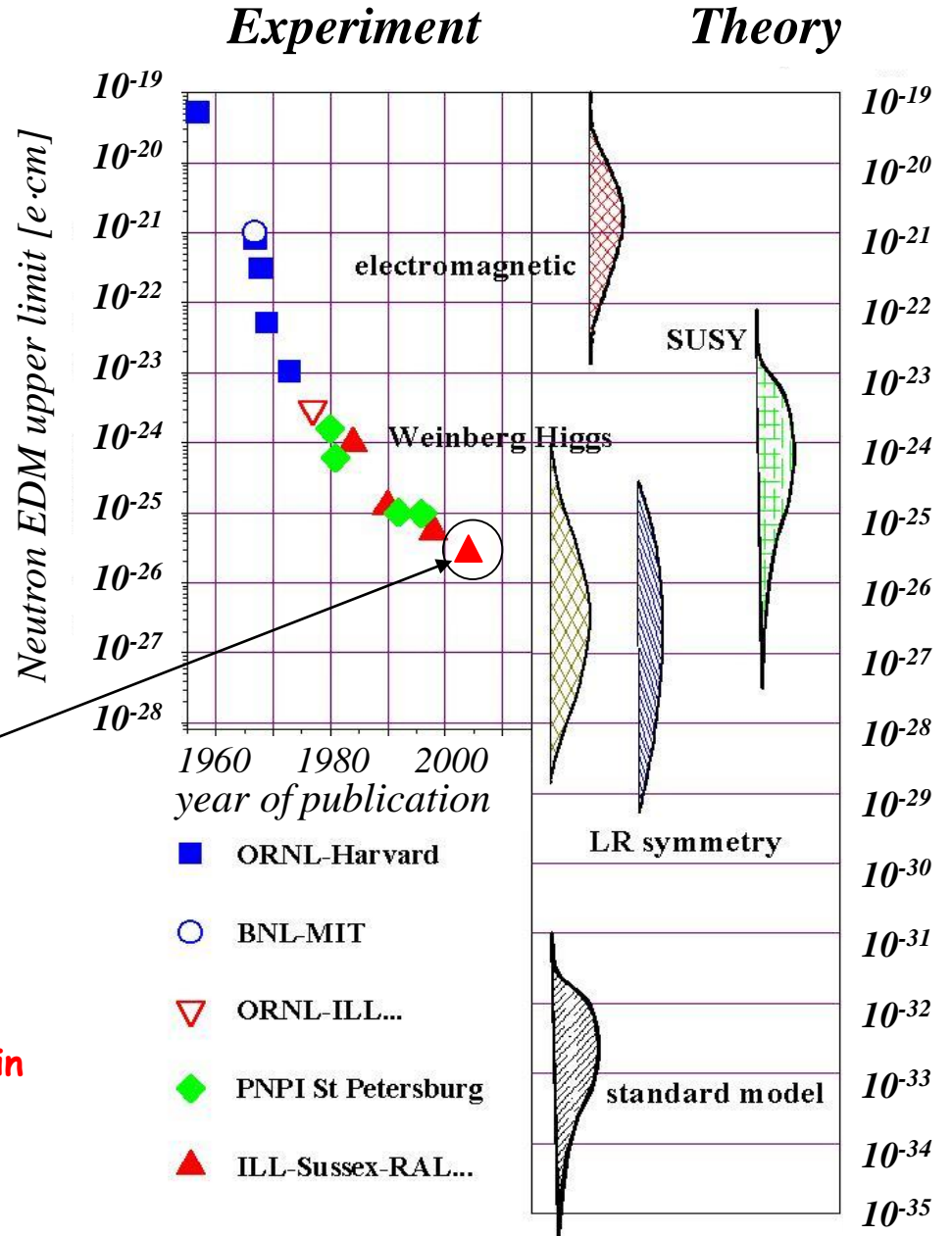
Progress at ~ order of magnitude per decade
 Standard Model out of reach
 Severe constraints on e.g. Super Symmetry



$$d_n = 1 \text{ e}\cdot\text{cm}$$

$$|d_n| < 3 \times 10^{-26} \text{ e}\cdot\text{cm}$$

“It is fair to say that the neutron EDM has ruled out more theories (put forward to explain K_0 decay) than any experiment in the history of physics” R. Golub



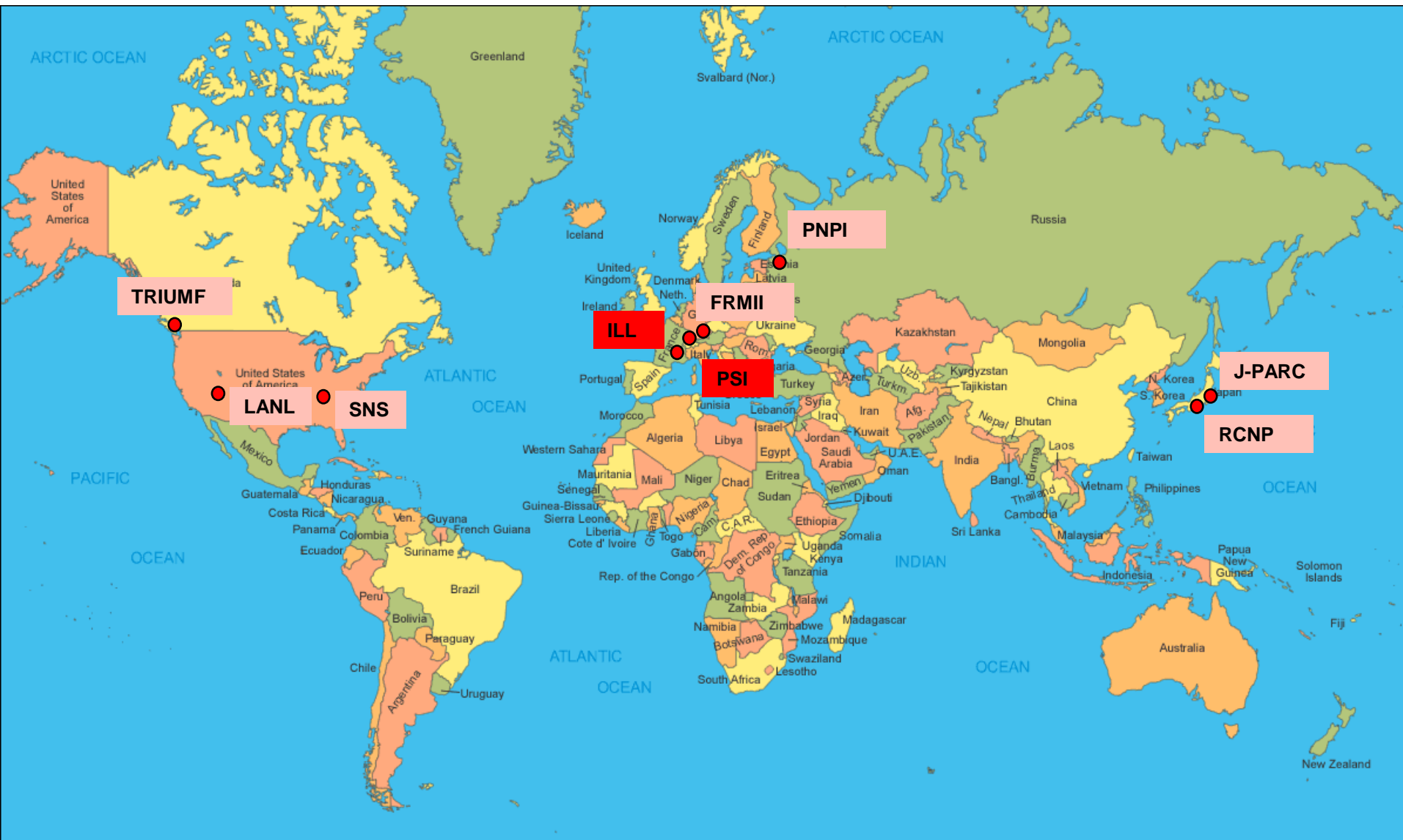
PNPI double-chamber nEDM spectrometer at PF2/MAM



$$|nEDM| \leq 5.5 \cdot 10^{-26} e \cdot cm \quad \text{at 90\% confidence level}$$

A.P. Serebrov et al., Pis'ma v ZhETF 99 (2014) 7

Worldwide nEDM Searches





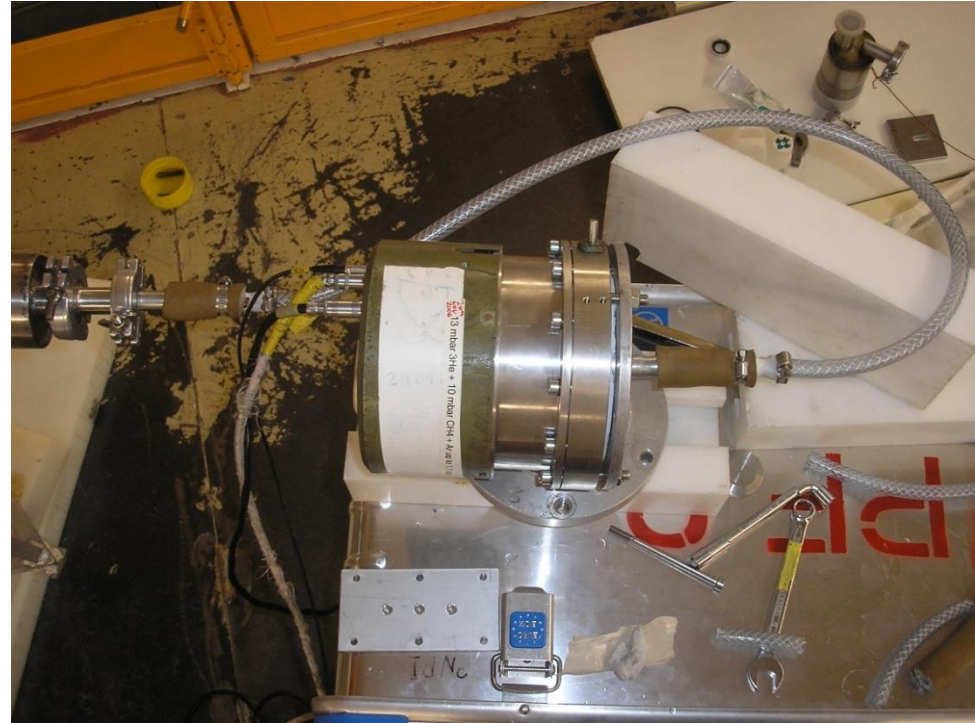
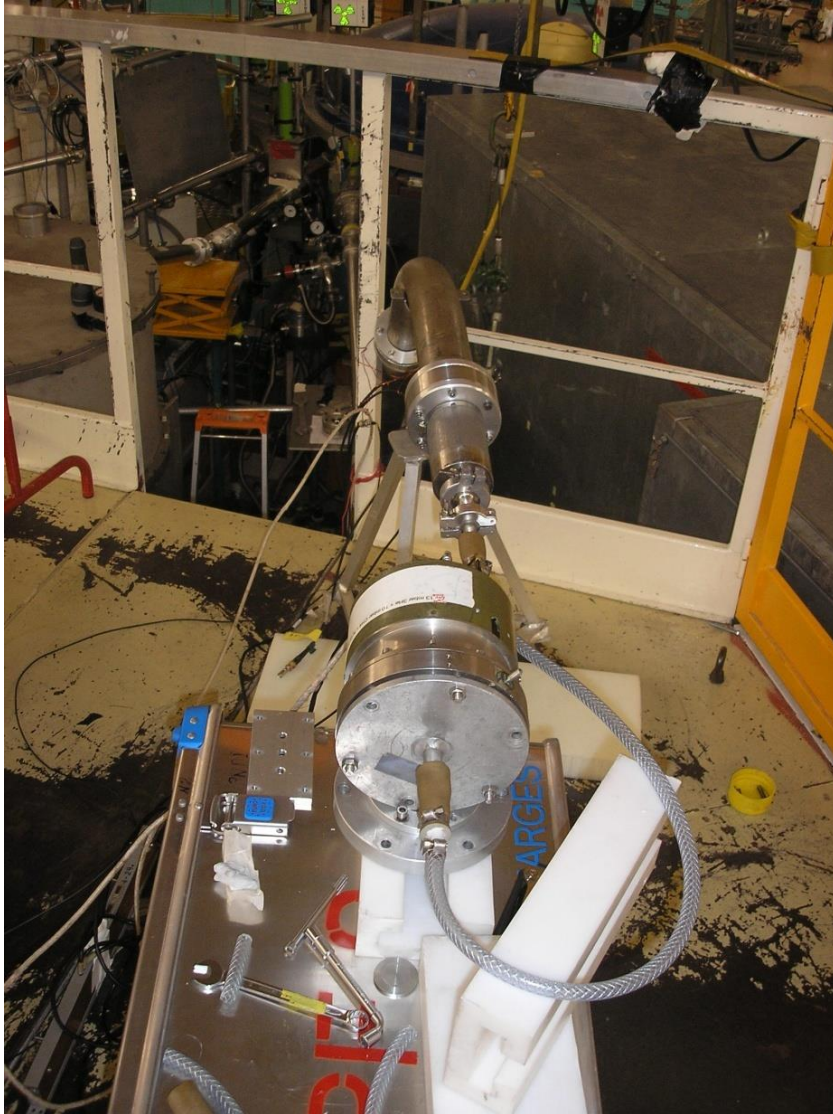
WHAT WOULD
MACGYVER DO?



UCN are always good for a surprise!

Transmission through flexible water hose

Yu. Panin et al., RRC KI Moscow

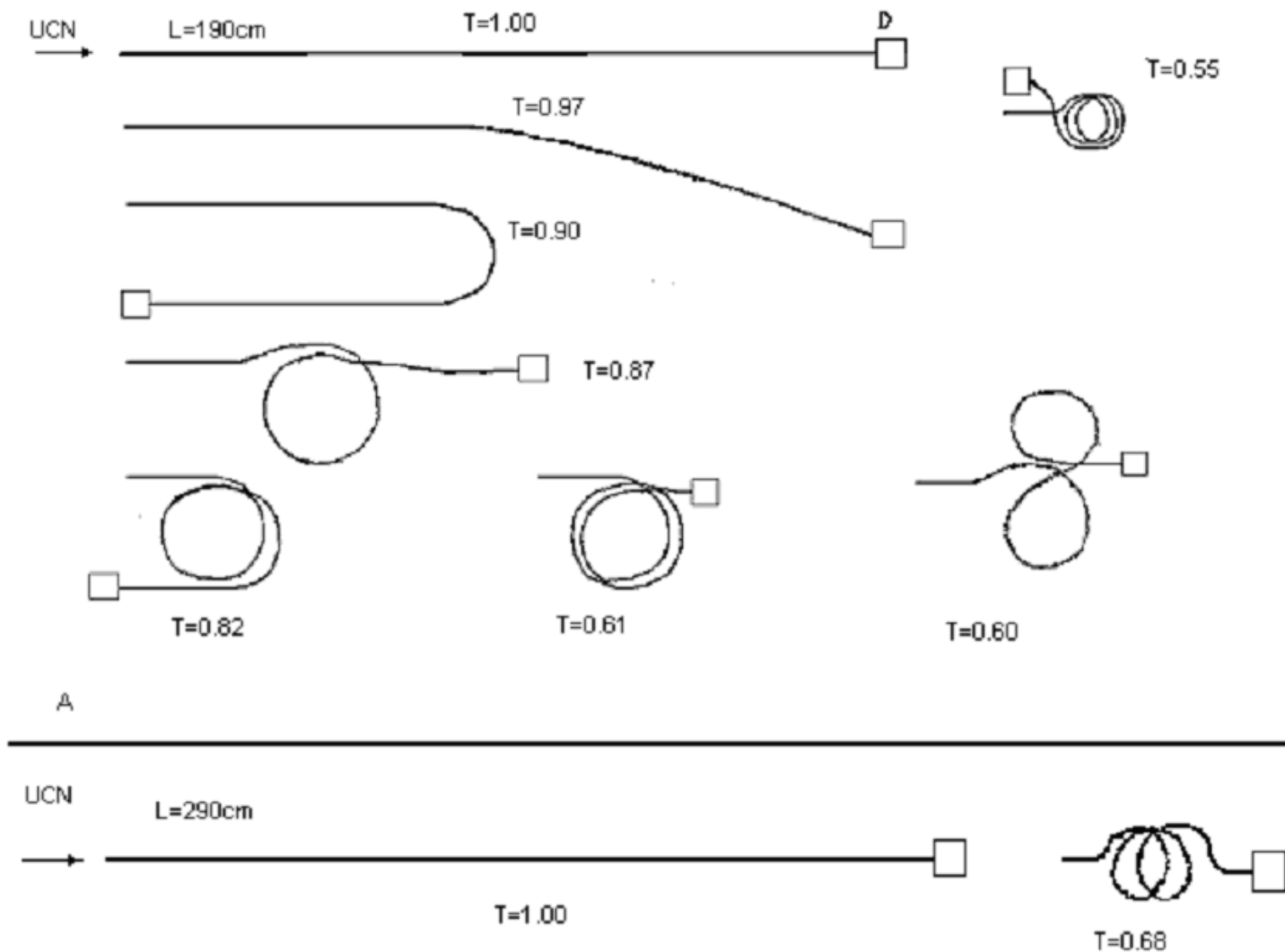


Surprising result

(80 cm hose with 8 mm inner diameter)

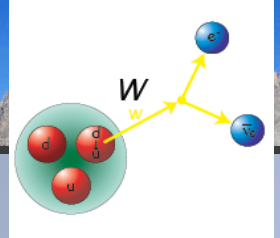
transmission around 85%

Relative Transmission Probability of “fancy guides”



Top view:

- The tube length equals $L=190$ cm.
- The tube length equals $L=290$ cm; the tube is coated inside with thin layer of Fluorine polymer.



The free neutron lifetime: $n \rightarrow p + e^- + \bar{\nu}_e$ (+782 keV)

$$\frac{1}{\tau_n} \propto G_F^2, V_{ud}^2, \lambda^2 \quad \lambda = \frac{g_A}{g_V}$$

$$n \rightarrow p + e^- + \bar{\nu}_e + \gamma \quad BR(15keV) \approx 3 \times 10^{-3}$$

$$n \rightarrow H^0 + \bar{\nu}_e \quad BR \approx 4 \times 10^{-6}$$

Together with measurements of asymmetry coefficients in neutron decay

Weak interaction theory

Neutrino physics

Cosmology

Extraction of g_V, g_A and V_{ud}

Test of Conserved Vector Current (CVC: ' $g_V = 1$ ')
 Test of Unitary of CKM matrix ($V_{ud}^2 + V_{us}^2 + V_{ub}^2 = 1$)

Solar pp-process:

$$p + p \rightarrow d + e^+ + \nu_e \quad \sigma \propto g_A^2$$

Big bang:

Primordial elements' abundances

Neutrino induced reactions:

$$\bar{\nu}_\mu + p \rightarrow \mu^+ + n$$

$$\nu_\mu + n \rightarrow \mu^- + p$$

Neutrino detectors:

$$p + \bar{\nu}_e \rightarrow n + e^+$$

$$\sigma \propto \frac{1}{\tau_n}$$

Important input parameter for tests of the Standard Model of the weak interaction

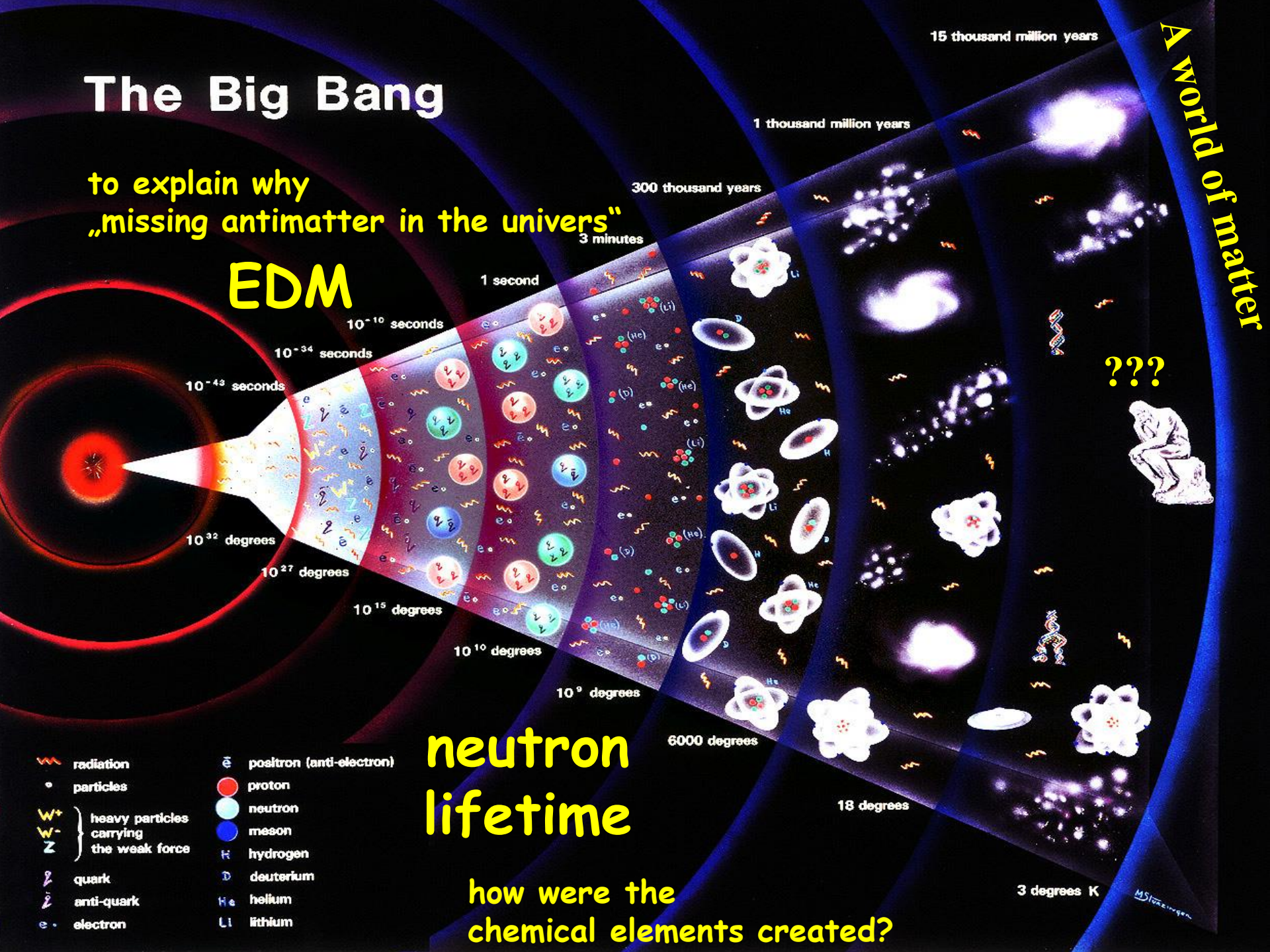
Necessary to understand matter abundance in the Universe

Necessary to calibrate Neutrino Detectors and to predict event rates

The Big Bang

to explain why
„missing antimatter in the univers“

EDM



neutron lifetime

how were the
chemical elements created?

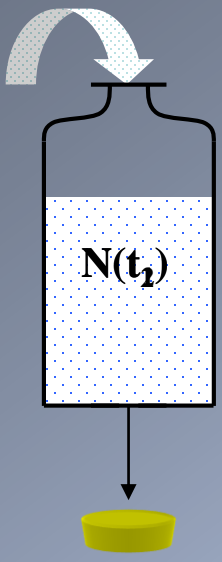
exponential decay law: $N = N_0 e^{-\lambda t}$

or, ultimately, measure the exponential decay directly

Storage experiments with UCN

“counting the surviving neutrons”

“UCN bottle”



$$\frac{1}{\tau_m} = \frac{1}{t_2 - t_1} \cdot \ln \frac{N(t_1)}{N(t_2)}$$

$$\frac{1}{\tau_m} = \frac{1}{\tau_\beta} + \underbrace{\frac{1}{\tau_{\text{wall}}} + \frac{1}{\tau_{\text{leak}}}}_{\rightarrow 0 \text{ (experiment)}} + \underbrace{\frac{1}{\tau_{\text{vacuum}}}}_{\rightarrow 0 \text{ (extrapolation)}} + \dots$$

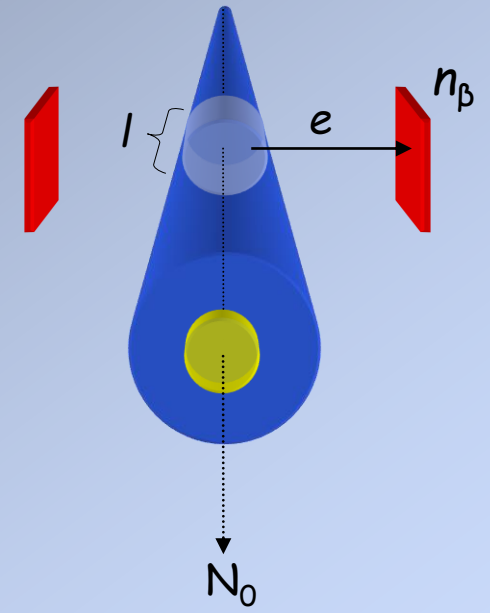
$$\frac{1}{\tau_{\text{wall}}} = \mu \cdot V_{\text{eff}} \rightarrow 0 \text{ (extrapolation)}$$

$$\rightarrow \frac{1}{\tau_m} = \frac{1}{\tau_\beta}$$

Two relative measurements

Beam experiments with cold neutrons

“counting the dead neutrons”



$$n_\beta = \frac{dN}{dt} = -\frac{N_0}{\tau_n} e^{-\frac{l}{v \cdot \tau_n}}$$

Two absolute measurements

Does the neutron lifetime depend on the measuring method?

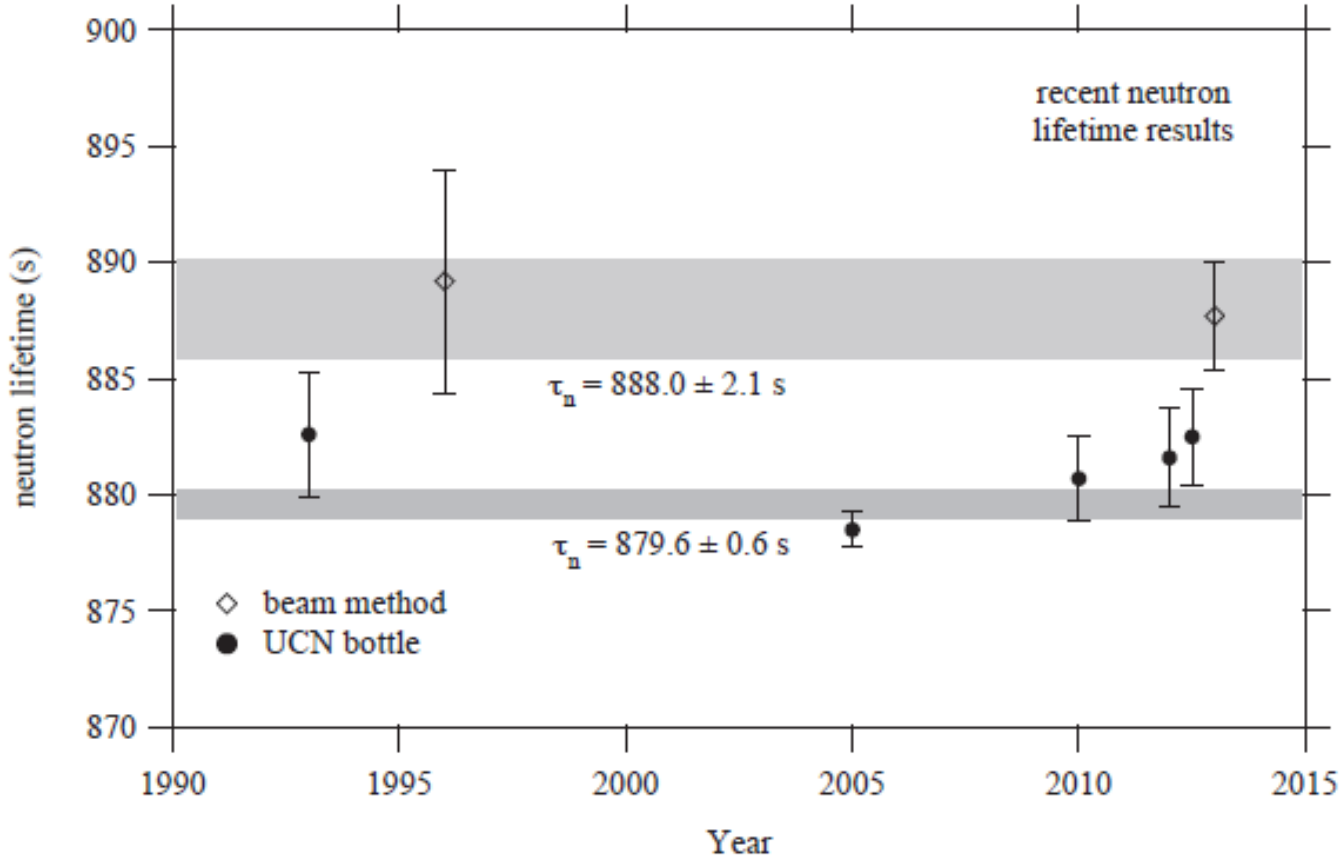
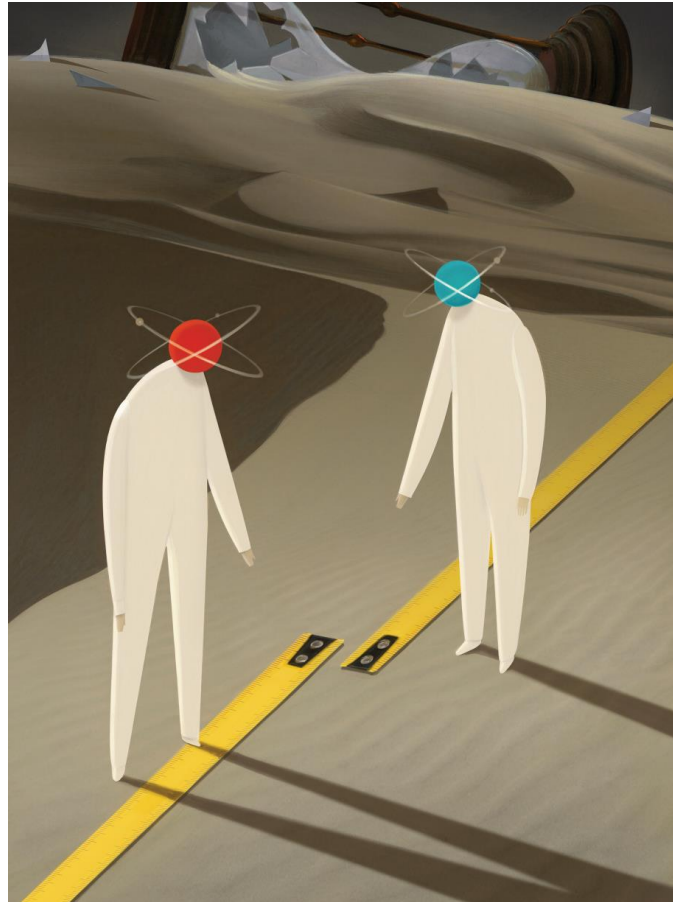


Figure 2: A summary of recent neutron lifetime measurements, showing the five UCN bottle [18, 16, 19, 20, 21] and two neutron beam [12, 15] results used in the 2014 PDG recommended value of $\tau_n = 880.3 \pm 1.1$ s. The shaded regions show the weighted average $\pm 1\sigma$ of each method, which disagree by 3.8σ .

F. Wietfeldt, arXiv:1411.3687v1 [nucl-ex]

For a broader public



PARTICLE PHYSICS

the neutron enigma

Two precision experiments disagree on how long neutrons live before decaying. Does the discrepancy reflect measurement errors or point to some deeper mystery?

By Geoffrey L. Greene and Peter Geltenbort

IN BRIEF

The best experiments in the world cannot agree on how long neutrons live before decaying into other particles. Two main types of experiments are under way; bottle traps count the number of neutrons that survive after various

intervals, and beam experiments look for the particles into which neutrons decay. Resolving the discrepancy is vital to answering a number of fundamental questions about the universe.

Illustration by Bill Mayer

April 2016, ScientificAmerican.com 37

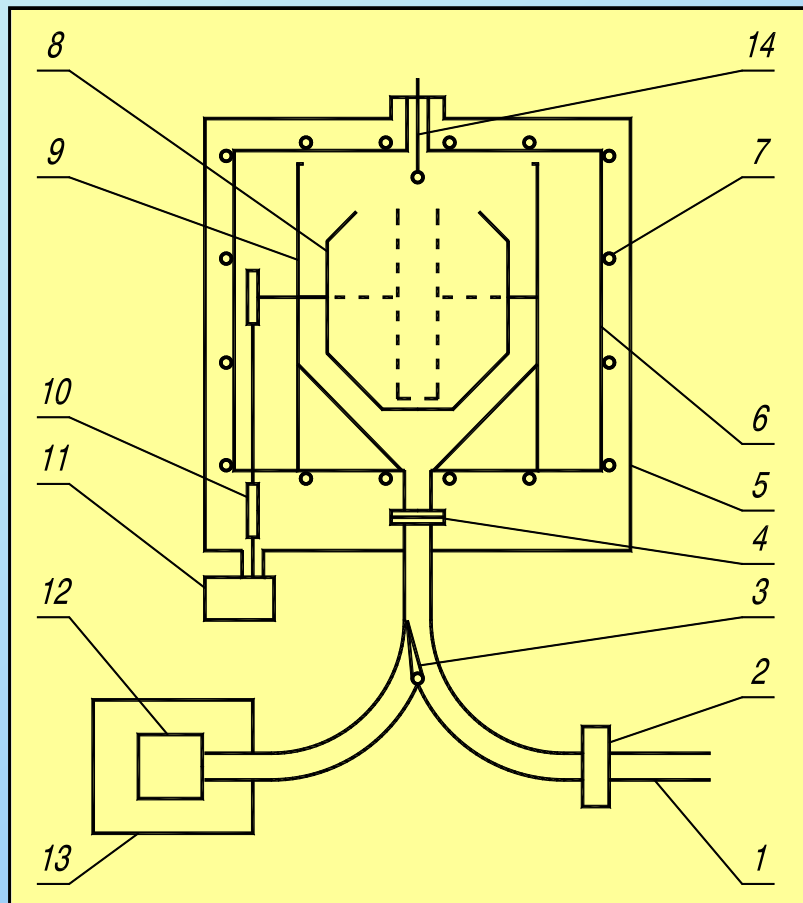
Translated and published in International Editions
of Scientific American:
France, Germany, Italy, Spain
China, Japan,
Russia, Poland, Israel, ...

谜样的 9 秒： 测不准的中子寿命

两个测量中子寿命的精密实验结果存在着差异。这种差异究竟反映了测量的误差，还是预示着一些更深层次的待解之谜？

撰文 杰弗里·L·格林 (Geoffrey L. Greene) 彼得·格尔滕博特 (Peter Geltenbort)
翻译 张淑潮 孙保华

Scheme of “Gravitrap”, the gravitational UCN storage system

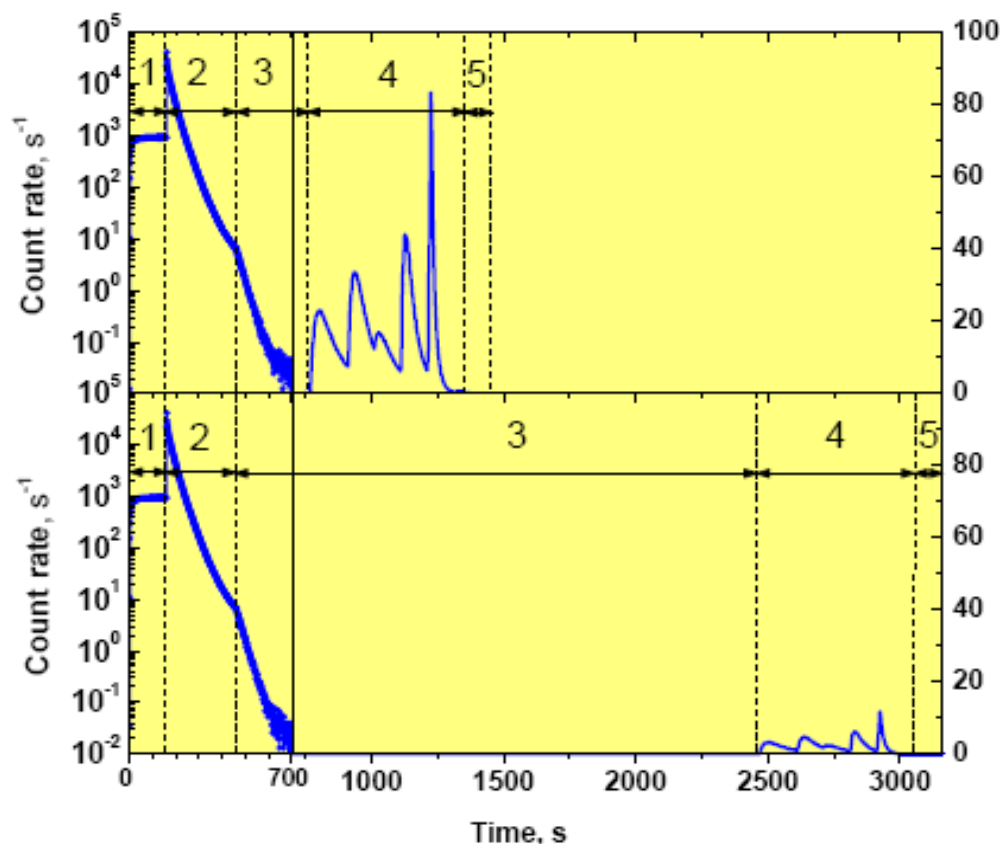


UCN traps are made from copper:

1. quasi-spherical (cylinder + 2 truncated cones) trap, inner
2. narrow (14 cm) cylindrical trap, inner surface - sputtered
3. wide (50 cm) cylindrical trap, inner surface - sputtered tita



Typical measuring cycle

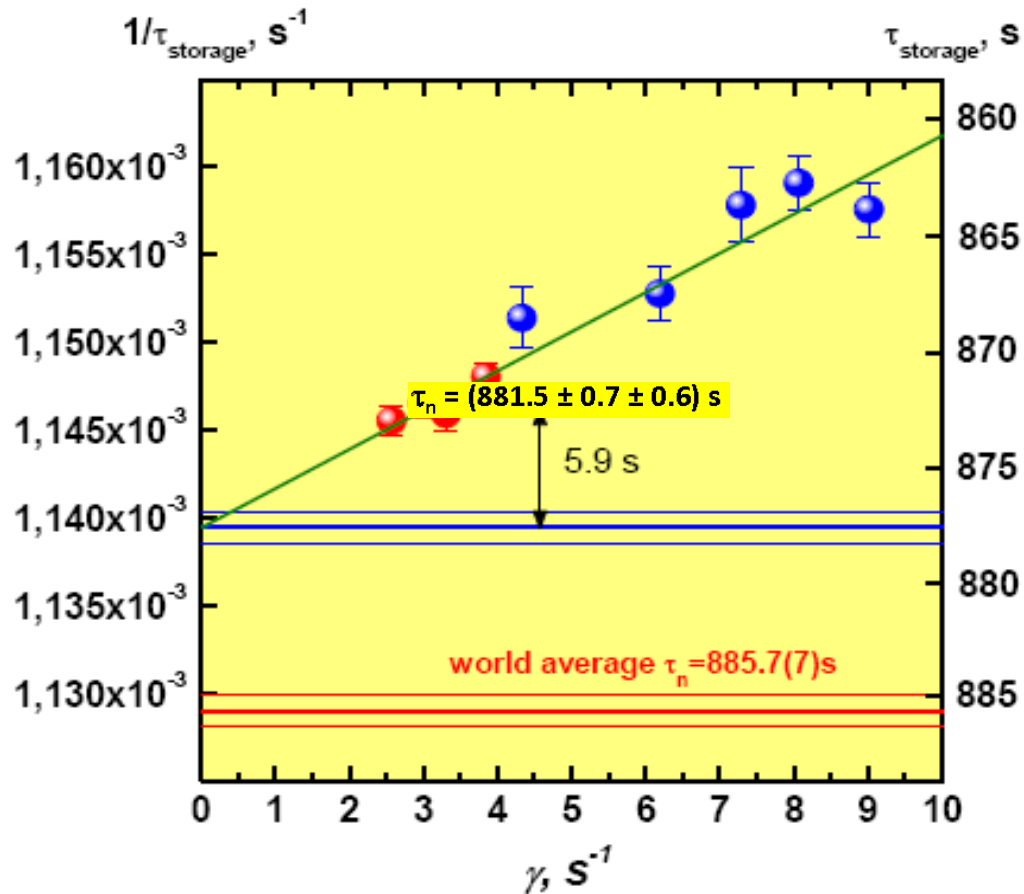


1. filling 160 s (time of trap rotation (35 s) to monitoring position is included);
2. monitoring 300 s;
3. holding 300 s or 2000 s (time of trap rotation (7 s) to holding position is included);
4. emptying has 5 periods 150 s, 100 s, 100 s, 100 s, 150 s (time of trap rotation (2.3 s, 2.3 s, 2.3 s, 3.5 s, 24.5 s) to each position is included);
5. measurement of background 100 s.

$$N(t_2) = N(t_1) \cdot \exp\left(-\frac{t-t_1}{\tau_{st}}\right)$$

$$\tau_{st} = \frac{t_2 - t_1}{\ln(N(t_1)/N(t_2))}$$

Extrapolation to n-lifetime



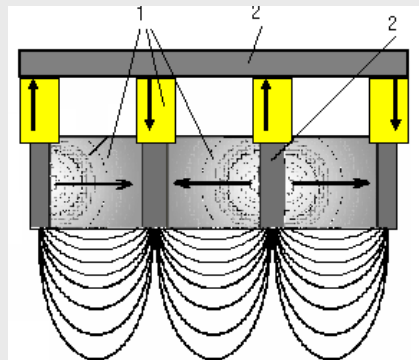
A.P. Serebrov et al. , Phys Lett B 605, (2005) 72-78 : **$(878.5 \pm 0.8) \text{ s}$**

General principle and design

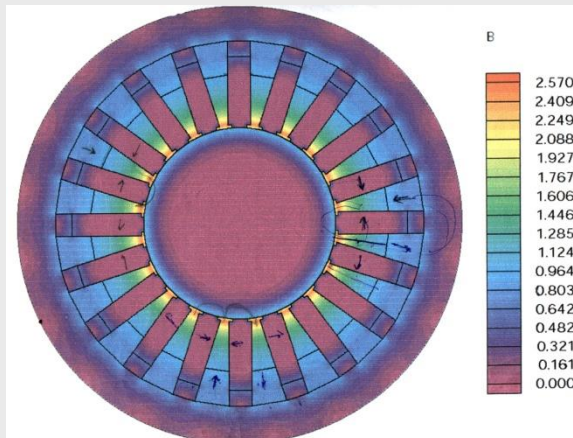
- For $\mu_n = -60.3$ neV/T, a 2T field generates a 120 neV barrier.
- Force due to field gradient, $F = -\mu (dB/dz)$, repels only one spin state.
- Use permanent magnets.

- **Step 1: 1D confinement**

- 1 – permanent magnets
- 2 – magnetic poles

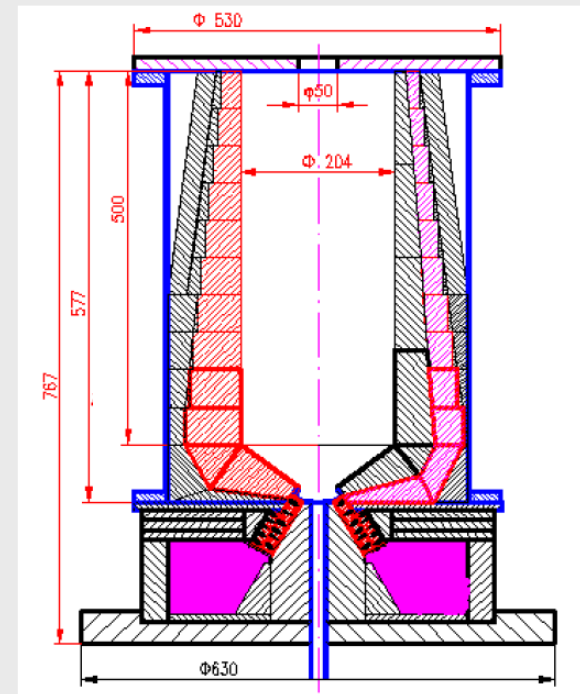


- **Step 2: 2D confinement**



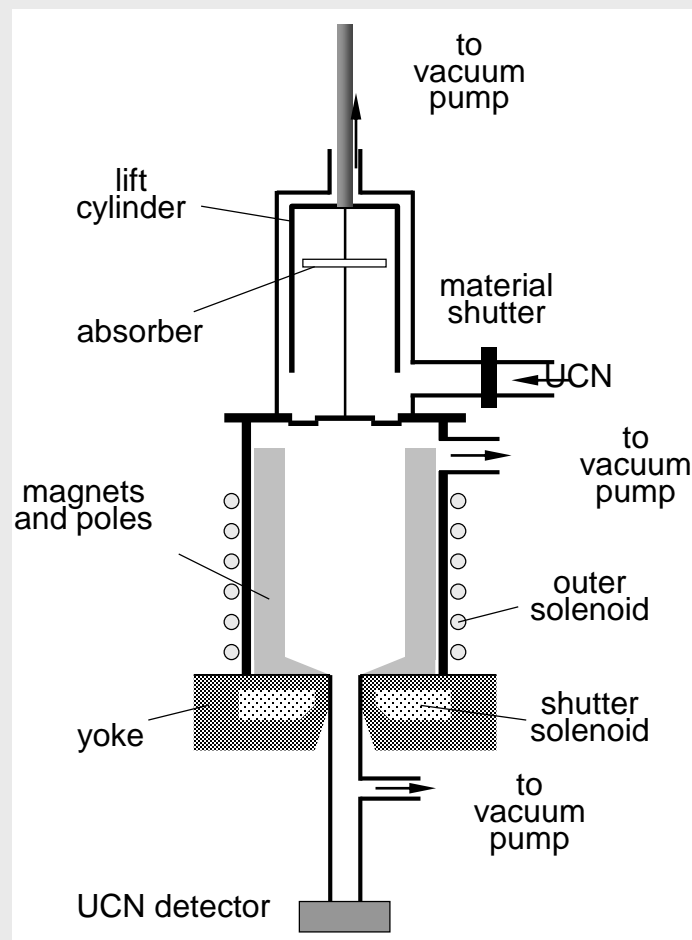
- **Step 3: 3D confinement**

- top (gravity)
- bottom (magnetic shutter)



Magnetic trap for neutron lifetime measurement

main elements: lift, trap, solenoid, shutter, detector



Lift: Fomblin coated Al cylinder + PE disk

$(878.3 \pm 1.9) \text{ s}$

V.F. Ezhov et al., arXiv:1412.7434 (2014)



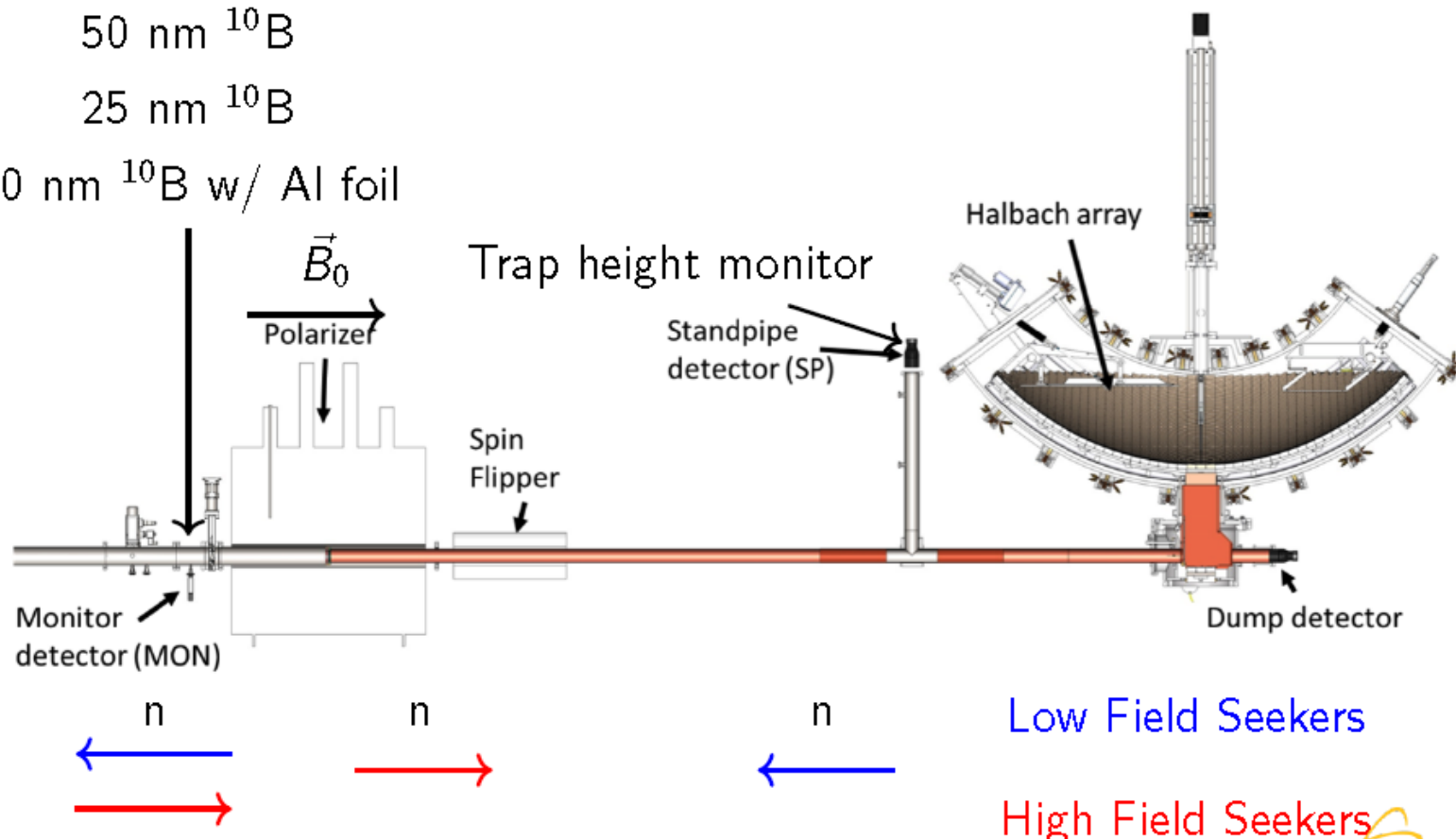
Side view of the experiment

UCN Monitors:

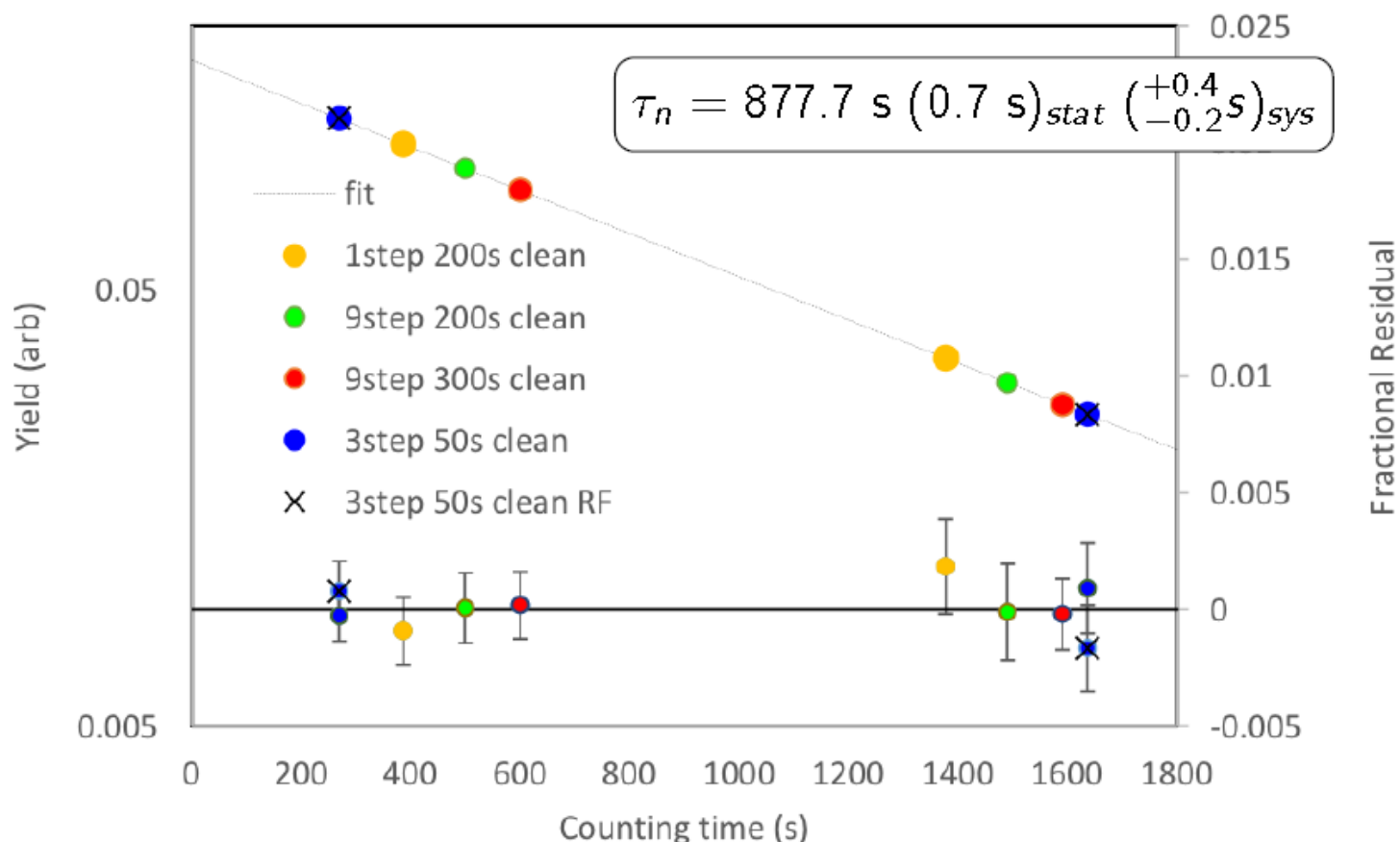
50 nm ^{10}B

25 nm ^{10}B

50 nm ^{10}B w/ Al foil



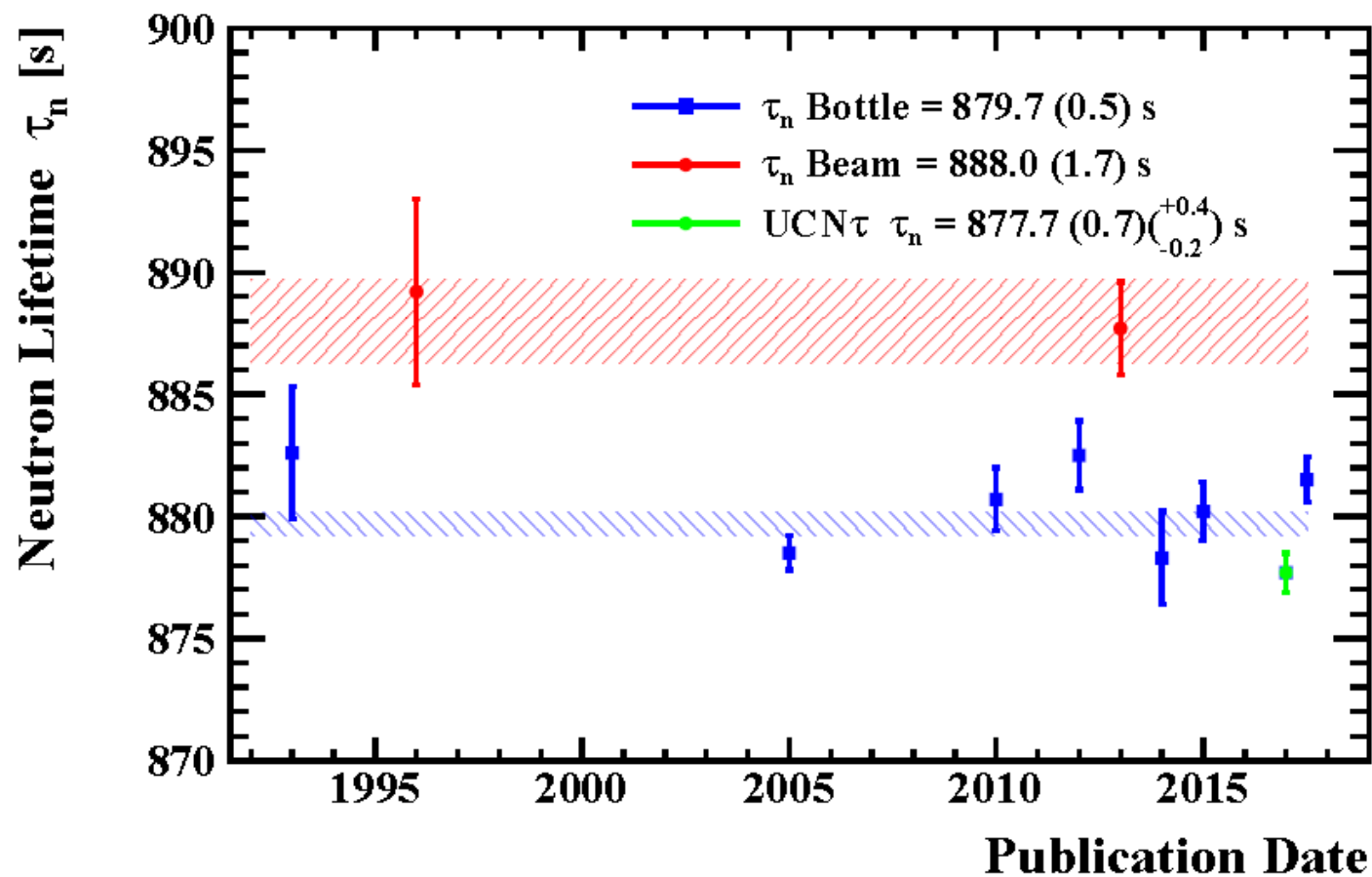
2016-2017 Lifetime Results



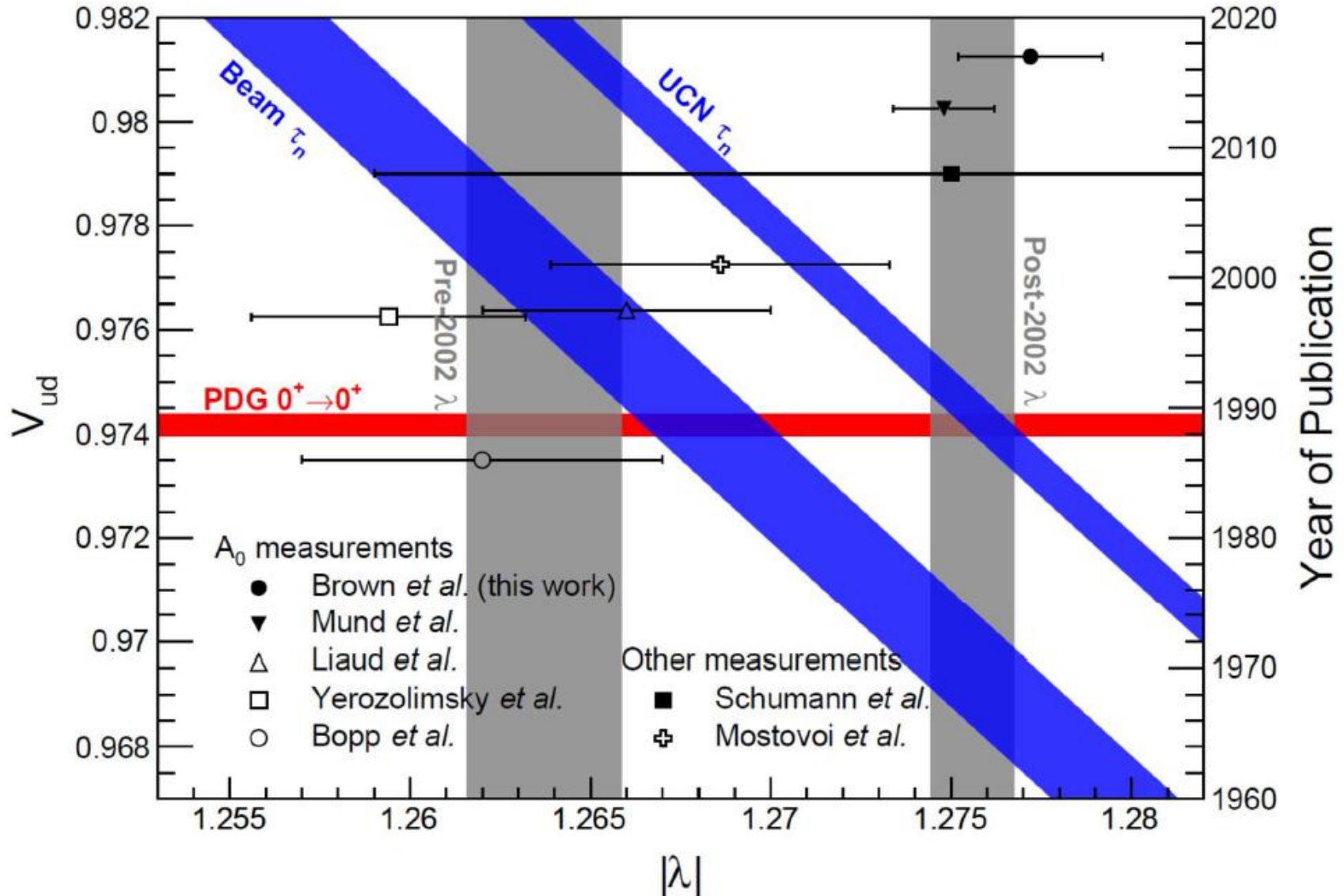
4

⁴R. W. Pattie Jr *et al*, Science **360**, 6389 p.627-632 (2018) (arXiv:1707.01817)

How does this compare?



Experimental Landscape: τ_n, V_{ud}



Neutron dark decay

PHYSICAL REVIEW LETTERS **120**, 191801 (2018)

Editors' Suggestion

Featured in Physics

Dark Matter Interpretation of the Neutron Decay Anomaly

Bartosz Fornal and Benjamín Grinstein

Department of Physics, University of California, San Diego, 9500 Gilman Drive, La Jolla, California 92093, USA



(Received 19 January 2018; revised manuscript received 3 March 2018; published 9 May 2018)

$$\text{Br}(n \rightarrow p + \text{anything}) \approx 99\%$$

Remaining 1% :



$n \rightarrow \text{SM particles (other than } p)$



$n \rightarrow \text{dark particle(s) + SM particle(s)}$



$n \rightarrow \text{dark particles}$



Possible decay channels

$$937.900 \text{ MeV} < M_f < 939.565 \text{ MeV}$$

Neutron \longrightarrow **dark particle + photon**

Neutron \longrightarrow **dark particle + e^+e^-**

Neutron \longrightarrow **two dark particles**

Neutron \longrightarrow **...**

Wishful thoughts

➔ ***“Missing Neutrons May Lead a Secret Life as Dark Matter”,***
C. Moskowitz, Scientific American (January 29, 2018)

➔ ***It would be truly amazing if the good old neutron turned out to be the particle enabling us to probe the dark matter sector of the universe***

Worldwide nLifetime Searches



Neutrons in the gravity field



- Schrödinger equation with linearized gravity

$$\left(-\frac{\hbar^2}{2m} \frac{\partial^2}{\partial z^2} + mgz \right) \varphi_n(z) = E_n \varphi_n(z)$$

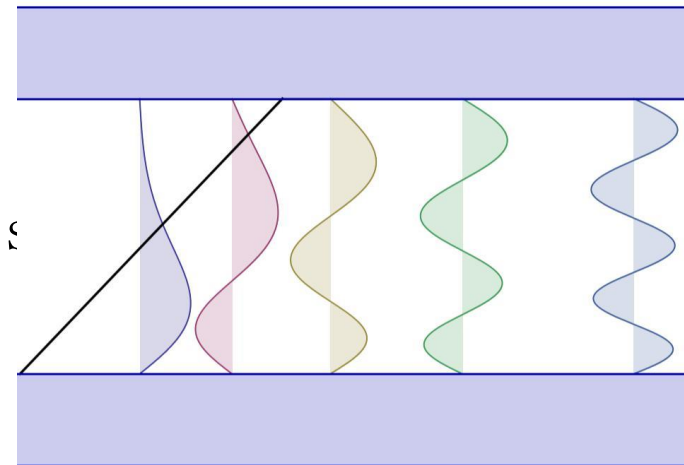
$$\text{bc: } \varphi_n(0) = 0, \quad \varphi_n(l) = 0$$

$$\varphi_n(z) = a_n \text{Ai} \left(\frac{z}{z_0} - \frac{E_n}{E_0} \right) + b_n \text{Bi} \left(\frac{z}{z_0} - \frac{E_n}{E_0} \right)$$

- bound, discrete states
- Non-equidistant energy levels

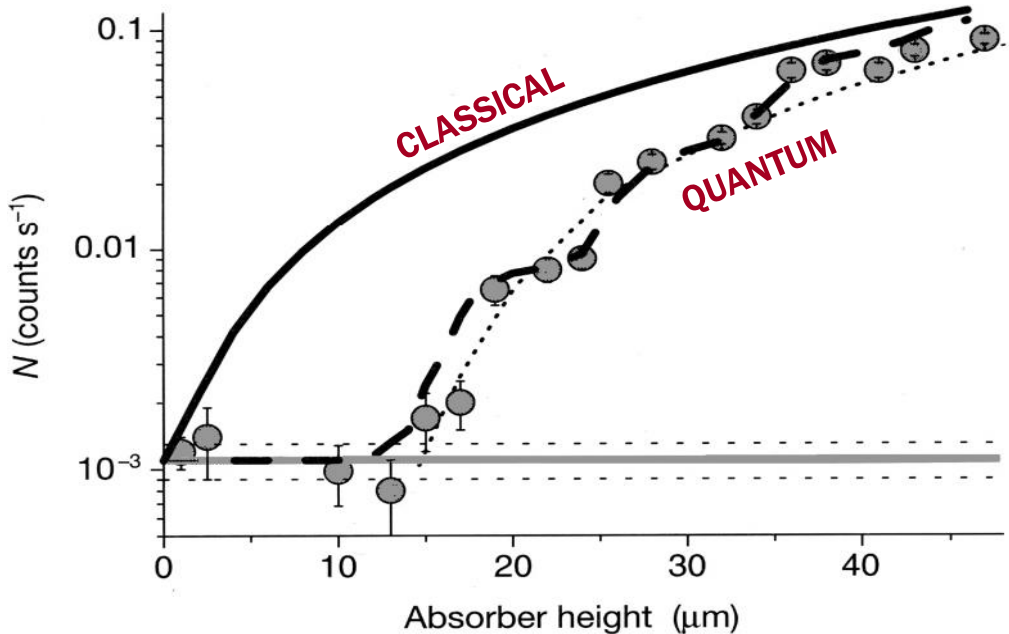
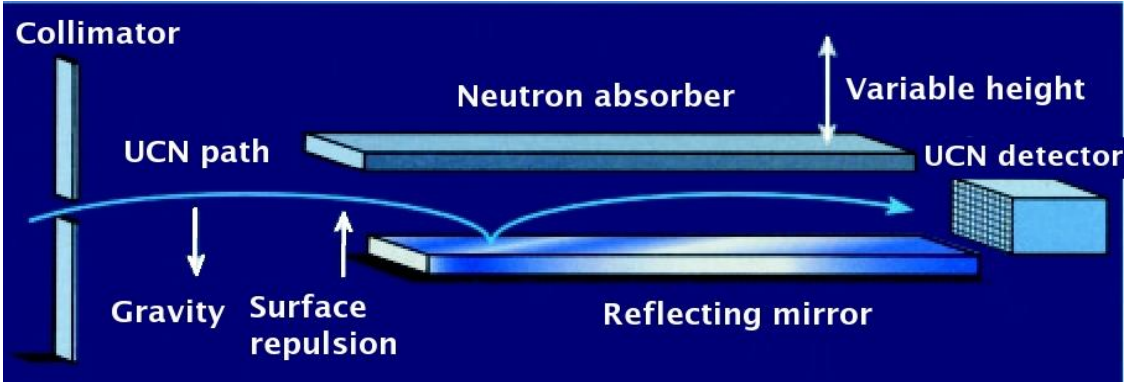
state	energy
1	1.41 peV
2	2.56 peV
3	3.98 peV

Slit width $l = 27 \mu\text{m}$



Discovery of neutron quantum states in 1999

Nesvizhevsky *et al*, Nature 415 (2002)



$$z_0 = \left(\frac{\hbar^2}{2m^2g} \right)^{1/3} = 5.87 \mu m$$

qBounce (H. Abele and his team, ATI Vienna)



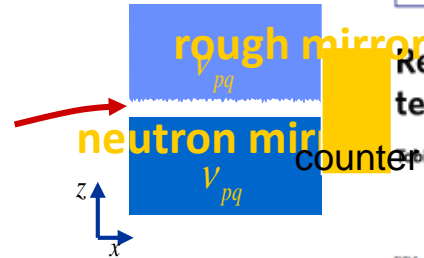
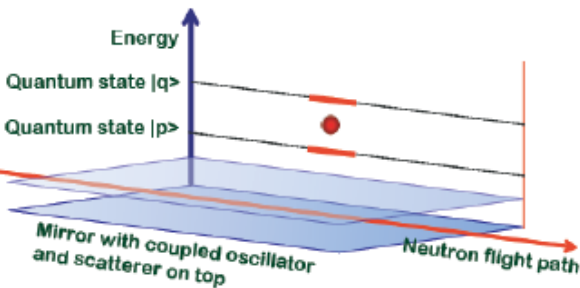
Motivation

- qBounce: quantized gravity bound states of ultra-cold neutrons
- Test of Newton's gravity potential at small distances (microns)
- Detection of new forces
- Tests for chameleons, axions

$$V(r) = -G \frac{m_i m_j}{r} (1 - a e^{-r/l})$$

Arkani-Hamed et al.: Physical Review D 59, 086004
(1999)

Gravity Resonance Spectroscopy



nature physics LETTERS
PUBLISHED ONLINE: 17 APRIL 2011 | DOI:10.1038/NPHYS1970

Realization of a gravity-resonance-spectroscopy technique

T. Jenke¹, Peter Geltenbort², Hartmut Lemmel^{1,2} and Hartmut Abele^{1,3,4*}

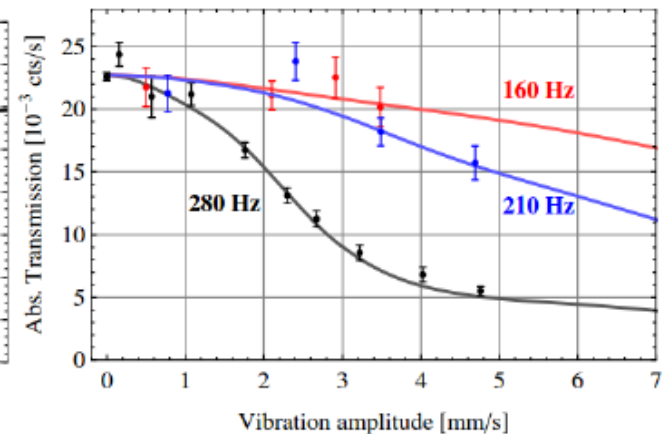
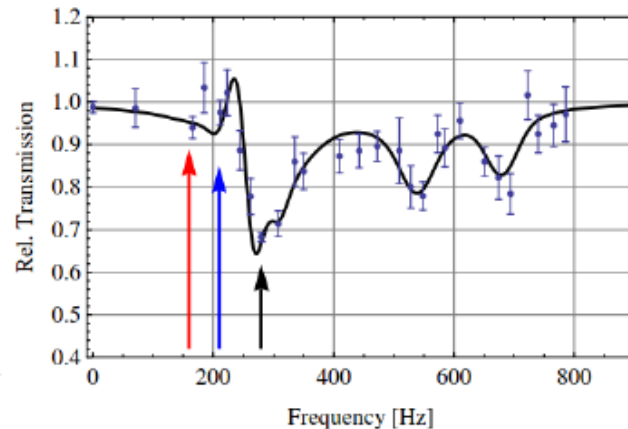
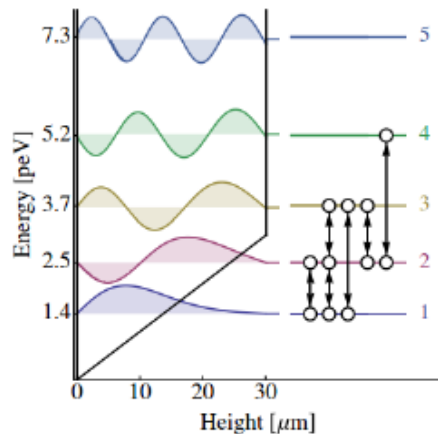
PRL 112, 151105 (2014) Selected for a Viewpoint in Physics PHYSICAL REVIEW LETTERS week ending 18 APRIL 2014

Gravity Resonance Spectroscopy Constrains Dark Energy and Dark Matter Scenarios

T. Jenke,^{1,2} G. Cronenberg,¹ J. Burgdörfer,² L. A. Chizhova,² P. Geltenbort,³ A. N. Ivanov,¹ T. Lauer,⁴ T. Lins,^{1,5} S. Rotter,² H. Saml,^{1,5} U. Schmidt,⁵ and H. Abele^{1,3}

¹Atominstut, Technische Universität Wien, Stadionallee 2, 1020 Wien, Austria
²Institute for Theoretical Physics, Vienna University of Technology, Wiedner Hauptstraße 8-10, 1040 Vienna, Austria
³Institut Laue-Langevin, BP 156, 6 Rue Jules Horowitz, 38042 Grenoble Cedex 9, France
⁴FRM II, Technische Universität München, Lichtenbergstraße 1, 85748 Garching, Germany
⁵Physikalisches Institut, Universität Heidelberg, Im Neuenheimer Feld 226, 69120 Heidelberg, Germany
(Received 26 November 2013; published 16 April 2014)

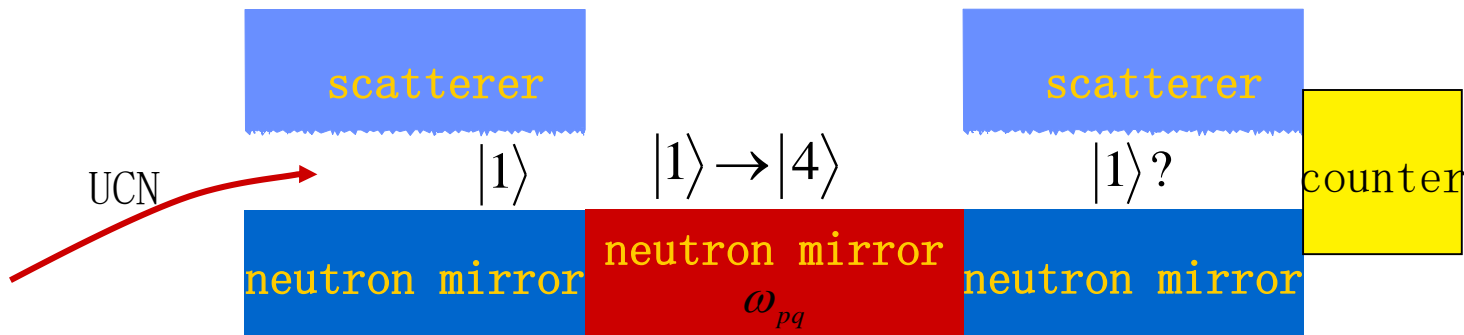
- first-time realization in 2009
- precision measurements in 2010 & 2011
- output:
 - experimental limits on Non-Newtonian gravity
 - study of rough surfaces (quantum transport phenomena)



Gravity Resonance Spectroscopy (GRS)



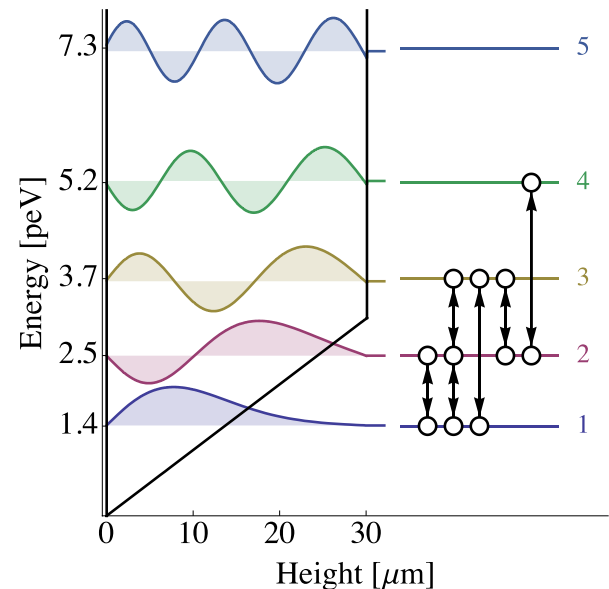
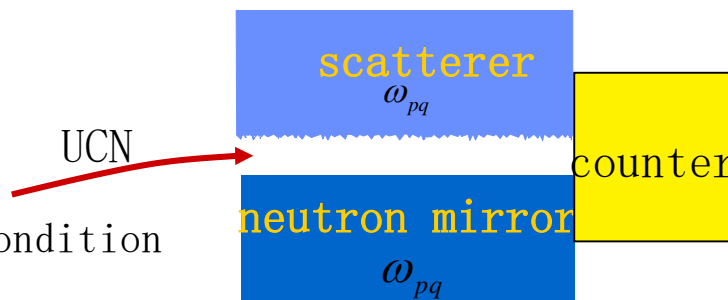
- Rabi setup (2012)



- First realisation (2009, 2010)

Rabi-like experiment with damping

but scatterer introduces 2nd boundary condition



T. Jenke et al.: “Realization of a gravity-resonance-spectroscopy technique” Nature Physics 7, 468 - 472 (2011)

Results

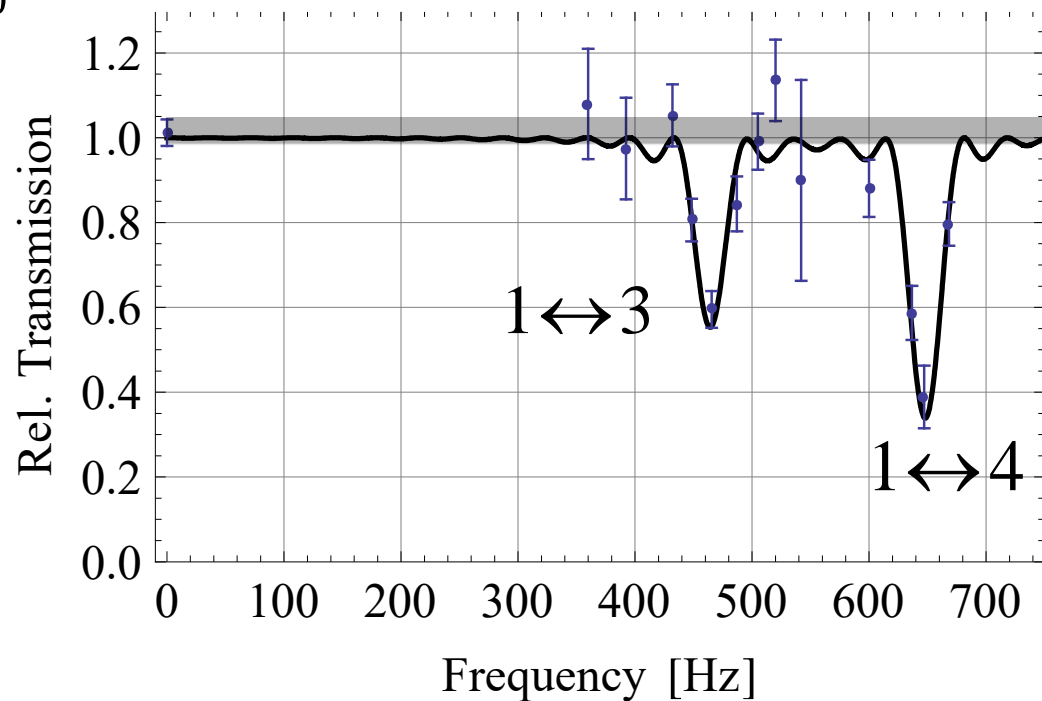


Transitions 1-3 and 1-4 observed

1-3: $(46 \pm 5)\%$ intensity drop

1-4: $(61 \pm 7)\%$

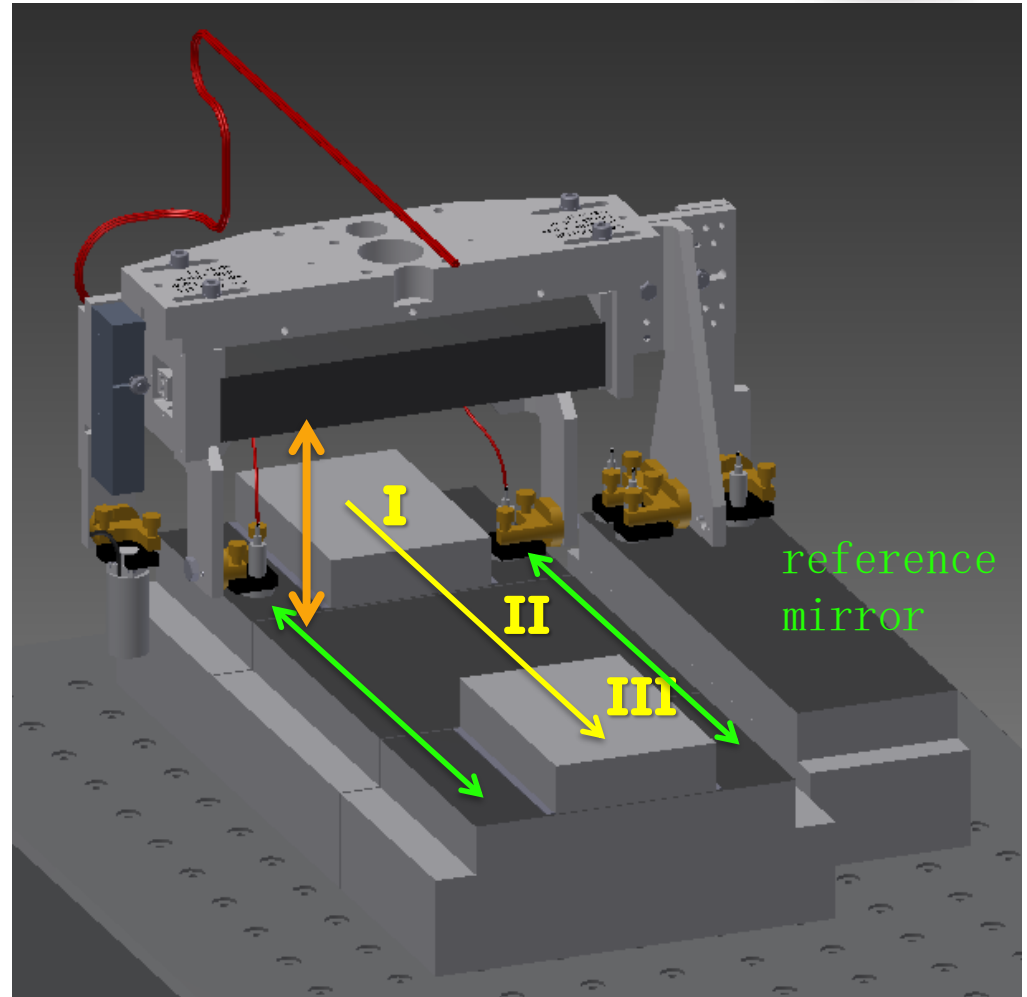
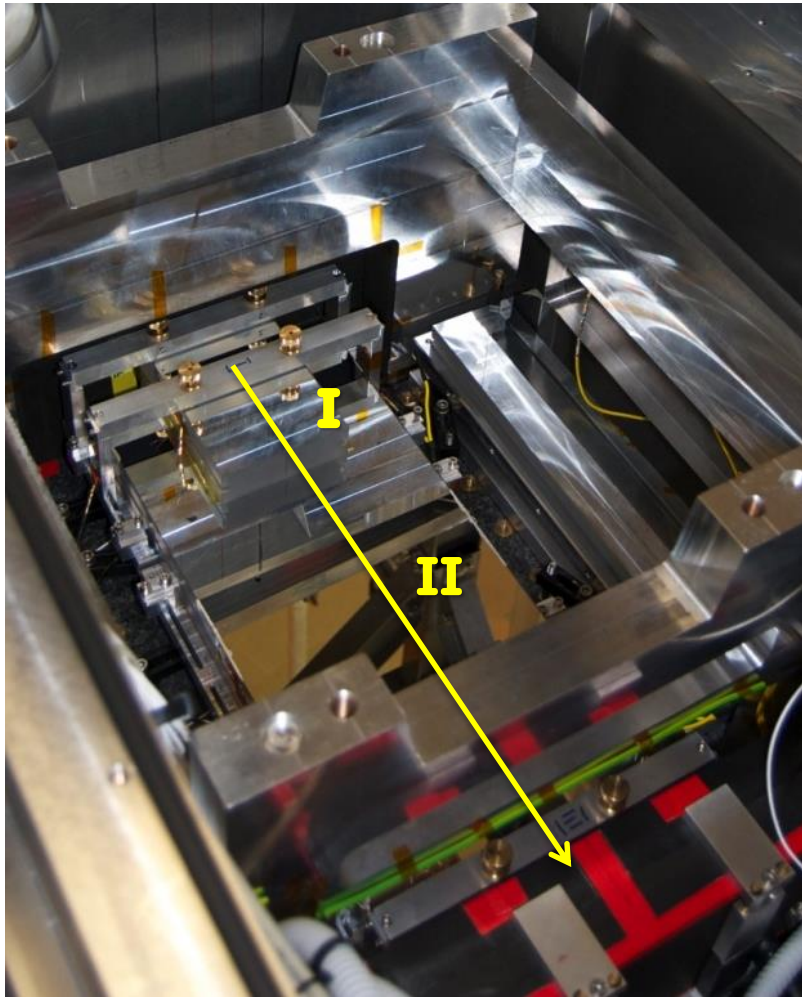
@ 2.1 mm/s



60 measurements

Preliminary, generic
fit

Setup

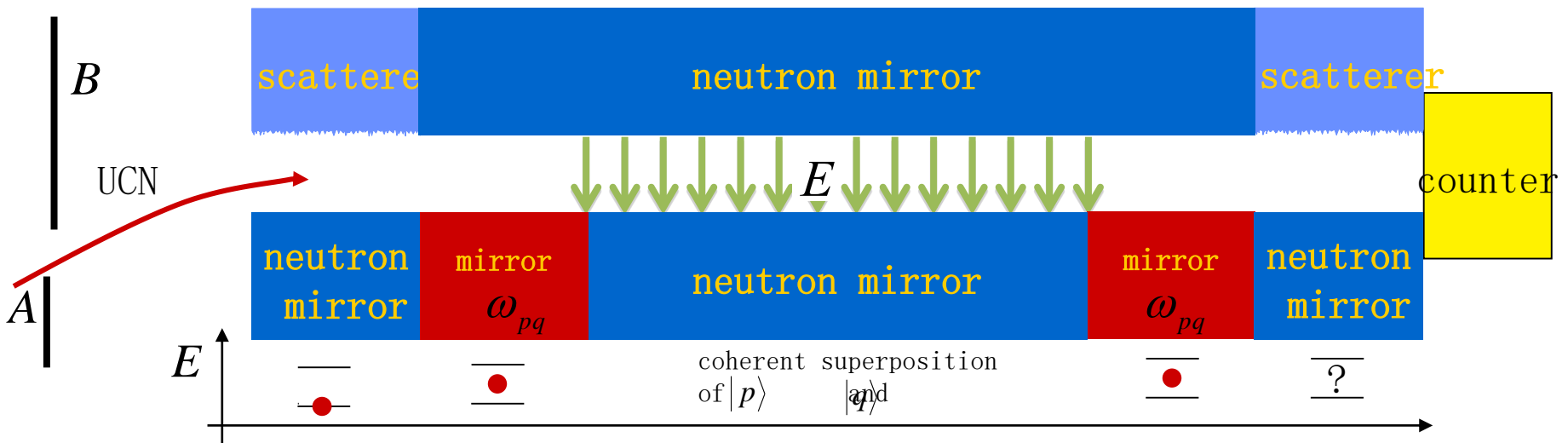


M. Horvath

Outlook: Probing neutrons neutrality

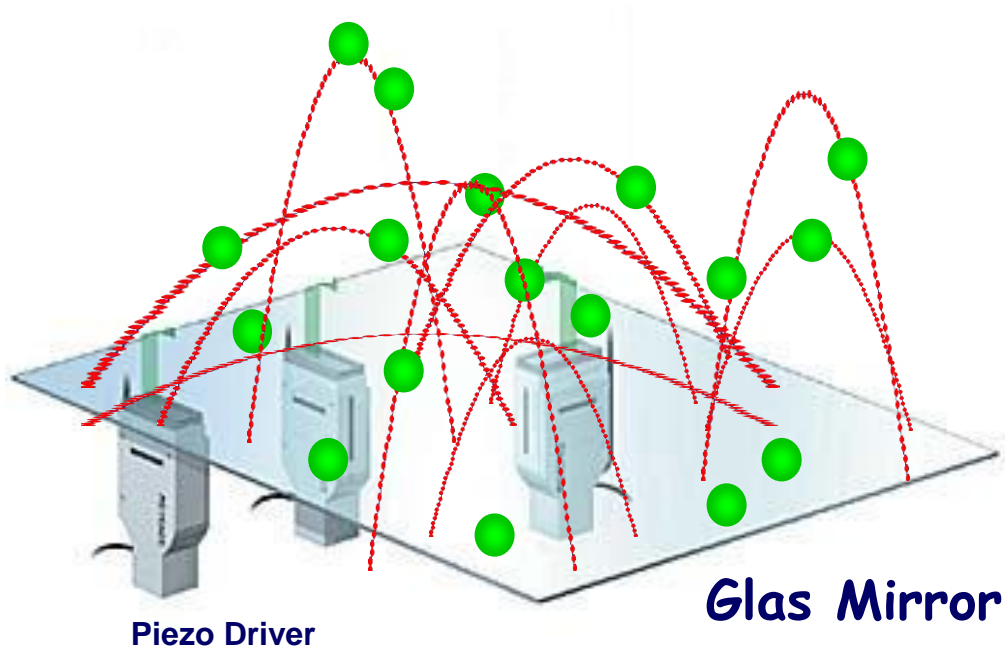


- Electric field modifies detectable phase



Durstberger-Rennhofer, K. et al. PRD 84, 036004 (2011)

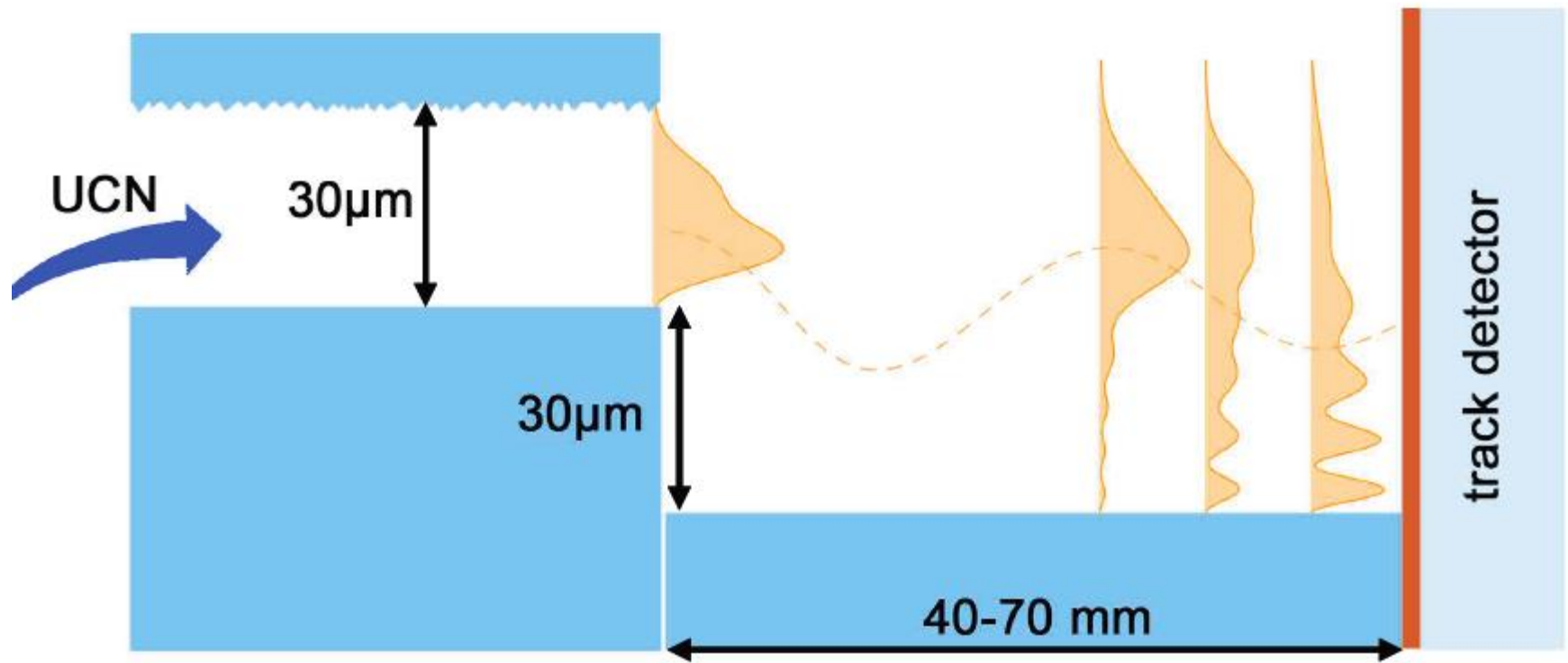
Realization of a Neutron Bouncing Ball Gravity Spectrometer



classical equation of motion for a falling body reflected on a mirror

quantum bouncing ultracold neutrons

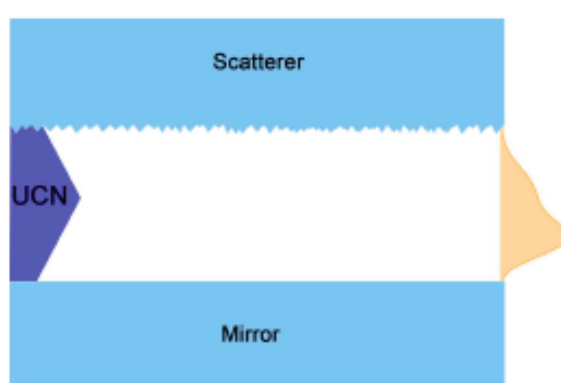
- State Selector
- Snapshots with spatial resolution detectors $\sim 1.5 \mu\text{m}$



Courtesy: M. Thalhammer

L

Preparation $L = 0$



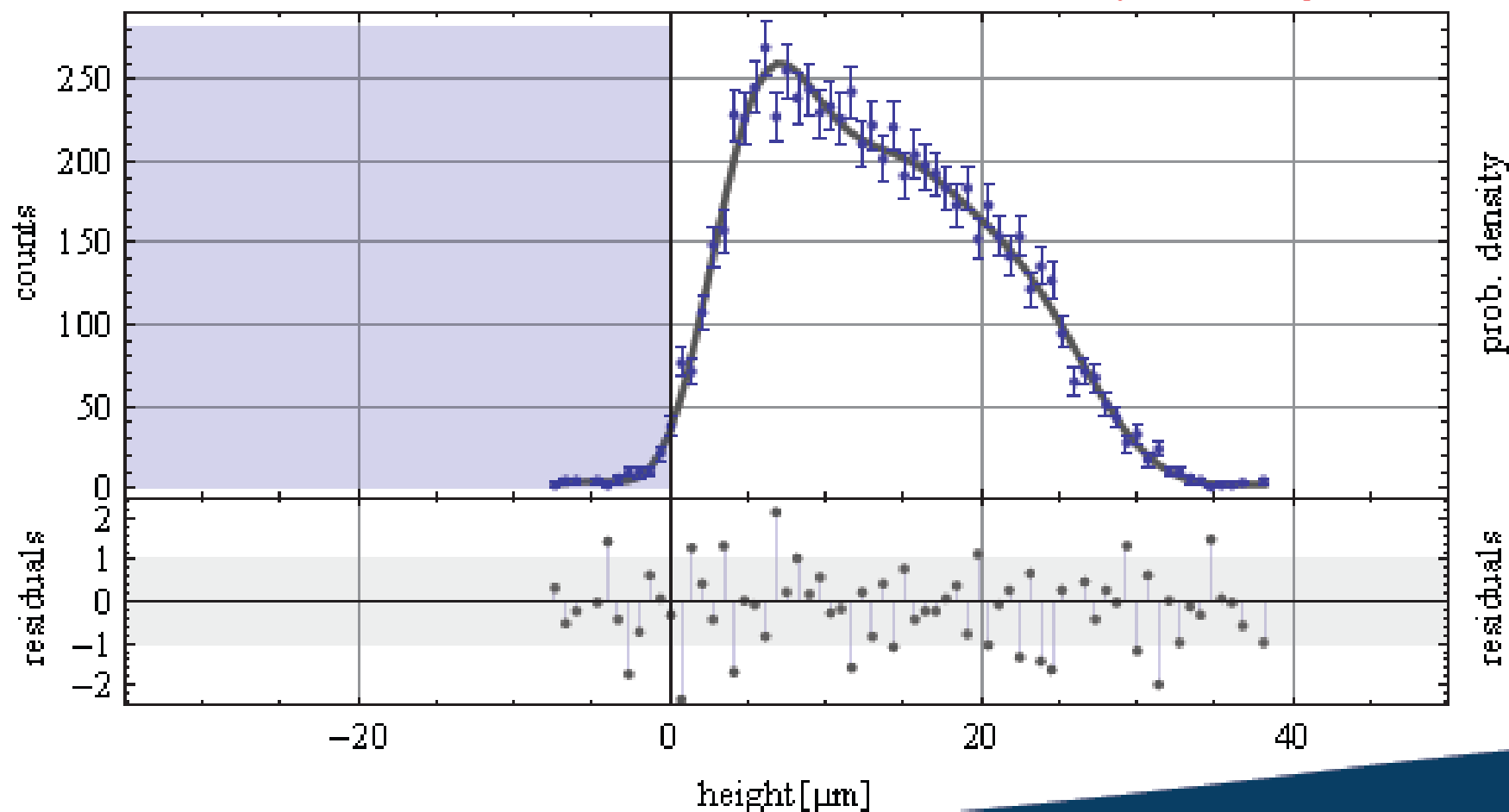
$$|\Psi_I(z, t_1)|^2 = \sum_n |c_n(t_1)|^2 \cdot |\psi_n(z)|^2$$

$$|c_1|^2 = 45\%$$

$$|c_2|^2 = 36\%$$

$$|c_3|^2 = 18\%$$

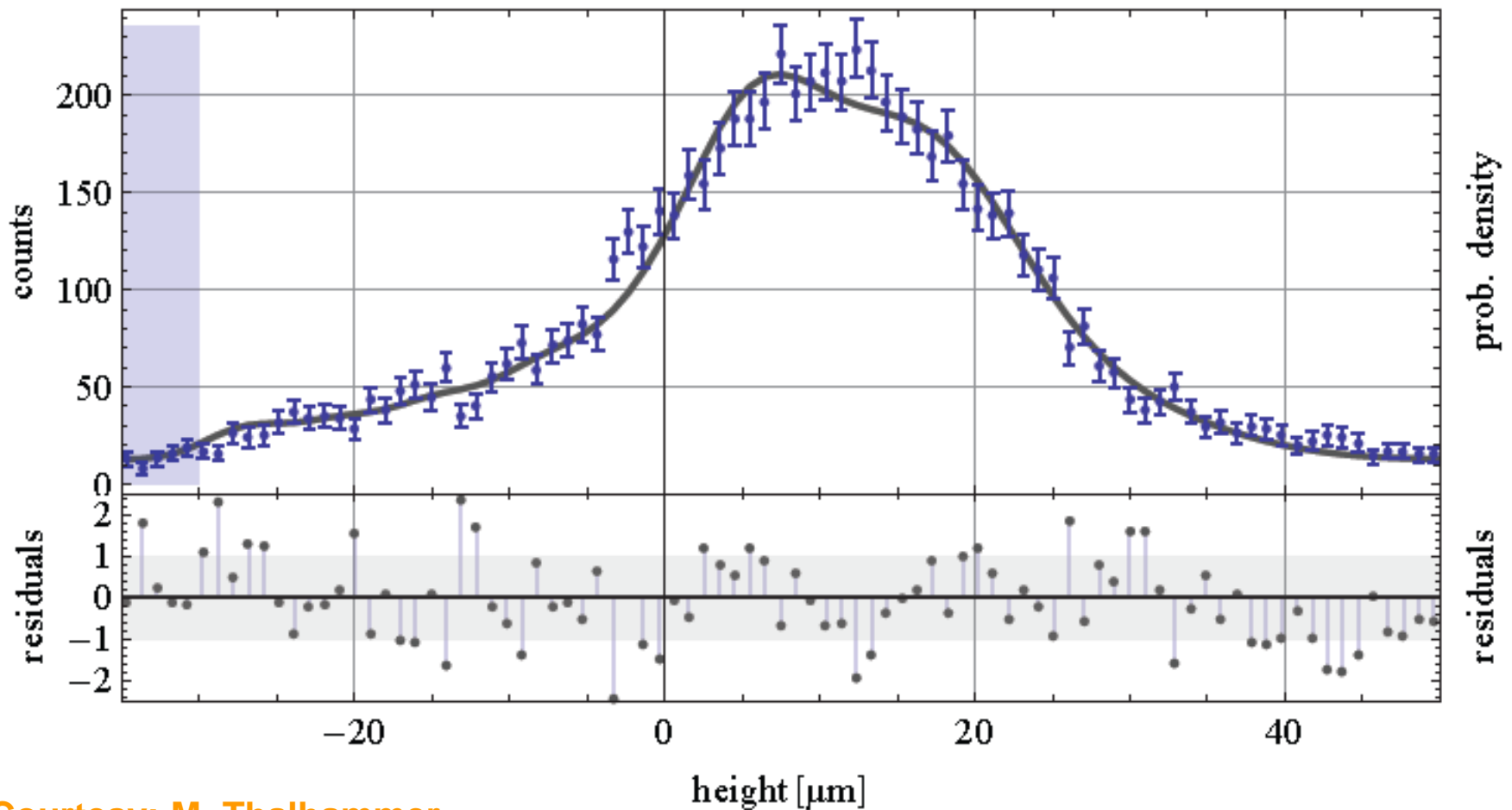
preliminary



Courtesy: M. Thalhammer

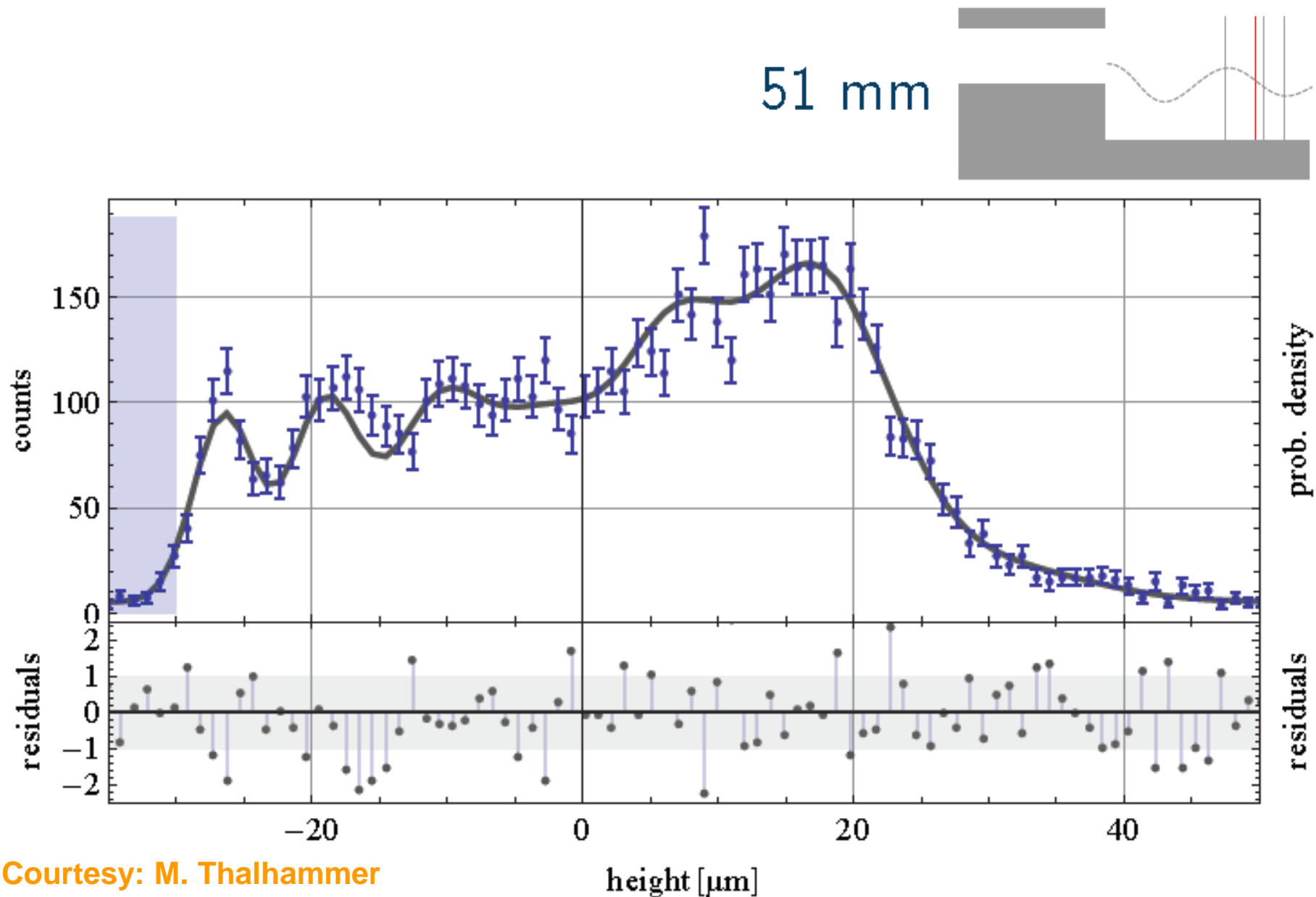
2nd bounce, 2nd turning point, $L = 41 \text{ mm}$

41 mm



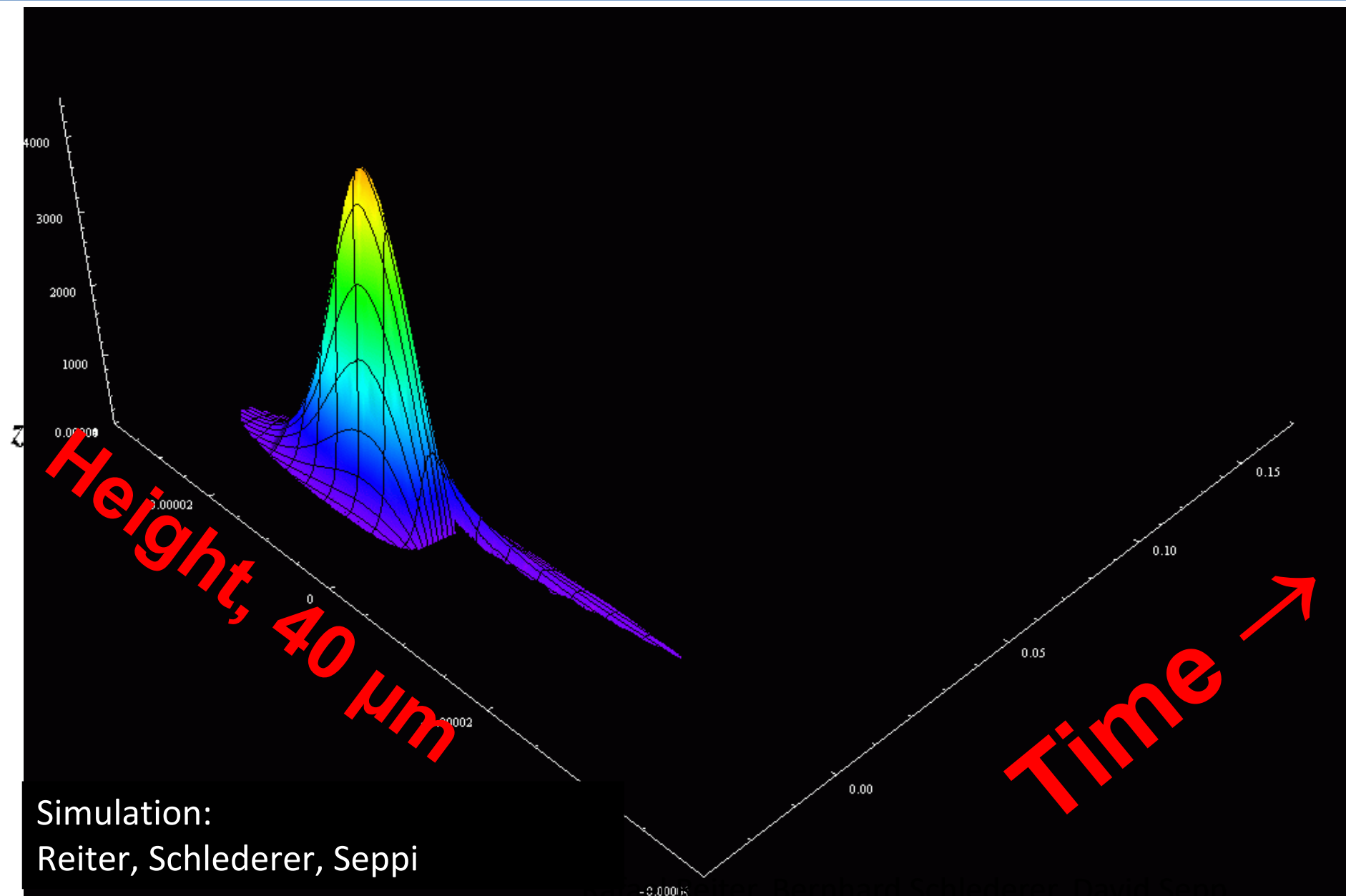
Courtesy: M. Thalhammer

Move downwards, $L = 51$ mm



Courtesy: M. Thalhammer

Show Case: *q*BOUNCE



I hope I could convince you that

ultracold neutrons

- due to the fact that they are storable -
continue to be

a fancy and **powerful tool in fundamental physics**

... and that

ILL's UCN facility PF2

and the other Nuclear and Particle Physics installations

are still very attractive places for fundamental research



Thank you, merci beaucoup, dankeschön for your attention!

International Workshop on Particle Physics at Neutron Sources

24/05/2018 -26/05/2018

Institut Laue-Langevin, Grenoble, France

Main topics:

- Properties of the Neutron
- Fundamental Symmetries and Interactions
- Hadronic Parity Violation
- Search for $e\bar{\nu}$ -Neutrinos
- Gravity tests in the quantum regime
- New Techniques and Ideas

ORGANIZING COMMITTEE

S. Degenklob, P. Geltenbort, T. Jenke,
M. Jentschel, V. Nesvizhevsky, D. Rebreyend,
T. Soldner, A. Stutz, O. Zimmer,
L. Tellier (Workshop Assistant)

Invited review talks by:

- Clare BURRAGE (University of Nottingham, UK)
- Vincenzo CIRIGLIANO (Los Alamos National Lab, USA)
- Carlo GIUNTI (INFN Torino, Italy)
- Martin GONZALE S-ALONSO (CERN, Switzerland)
- Ernst RASEL (Universität Hannover, Germany)

DEADLINE

Registration 2 March 2018
Abstract submission 2 March 2018

The year 2018 also marks the 50th anniversary of the discovery of ultra-cold neutrons. On this occasion, Hartmut ABELE (TU Wien, Austria) will give a commemorative speech.



<https://indico.ill.fr/PPNS2018>

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~140 participants

6 invited talk

35 contributed talks

~60 posters

can be found on

<http://indico.ill.fr/PPNS2018>