

# ANISOTROPY IN PRE-FISSION NEUTRON SPECTRA OF $^{235}\text{U}(n,f)$

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Angular anisotropy of secondary neutrons was evidenced in neutron emission spectra (NES) [1], and prompt fission neutron spectra (PFNS) [2]. In case of NES that is due to pre-equilibrium/semi-direct mechanism of emission of first neutron in  $(n,nX)^1$  reaction, while in case of PFNS by exclusive spectra of pre-fission neutrons of  $(n,xnf)^1$  [3]. In  $^{239}\text{Pu}(n,xnf)^{1,\dots,x}$  and  $^{235}\text{U}(n,xnf)^{1,\dots,x}$  reactions observed PFNS demonstrate differing response to the emission of first pre-fission neutron in forward and backward semi-spheres. Mean energies of  $(n,nf)^1$  neutrons depends on its angle of emission  $\theta$  with respect to the incident beam. The average prompt fission neutron number, fission cross section, TKE depend on  $\theta$  as well. Exclusive spectra of  $(n,xnf)^{1,\dots,x}$  neutrons at  $\theta \sim 90^\circ$  are consistent with  $^{235}\text{U}(n,f)$  ( $^{235}\text{U}(n,xn)$ ) and  $^{239}\text{Pu}(n,f)$  ( $^{239}\text{Pu}(n,xn)$ ) observed cross sections and neutron emission data at  $E_n \sim 0.01-20$  MeV.

The correlations of the angular anisotropy of PFNS with the relative contribution of the  $(n,nf)$  fission chance to the observed fission cross section and angular anisotropy of neutron emission spectra are revealed. The exclusive spectra of  $(n,xnf)^{1,\dots,x}$  and  $(n,n\gamma)$  and  $(n,xn)^{1,\dots,x}$  reactions are calculated alongside with  $(n,f)$  and  $(n,xn)$  cross sections within Hauser-Feshbach formalism, the angular anisotropy of  $(n,nX)^1$  neutrons being included (Fig. 1). The ratios of mean PFNS energies  $\langle E \rangle$  for forward and backward emission of  $^{235}\text{U}(n,xnf)$  pre-fission neutrons (Fig. 2) are consistent with measured data [2].

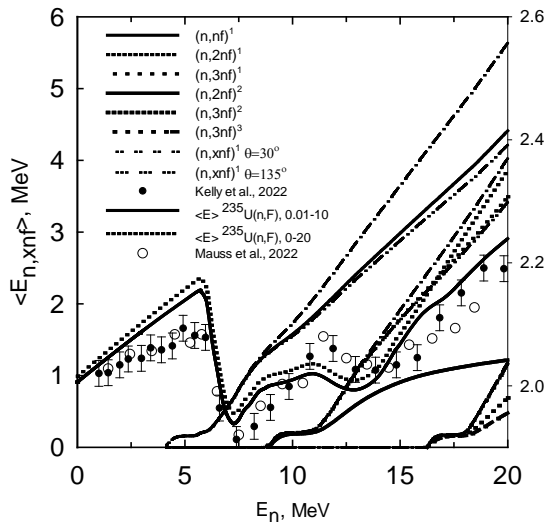


Fig. 1  $\langle E \rangle$  and  $\langle E_{n,xnf} \rangle$  of  $^{235}\text{U}(n,f)$ .

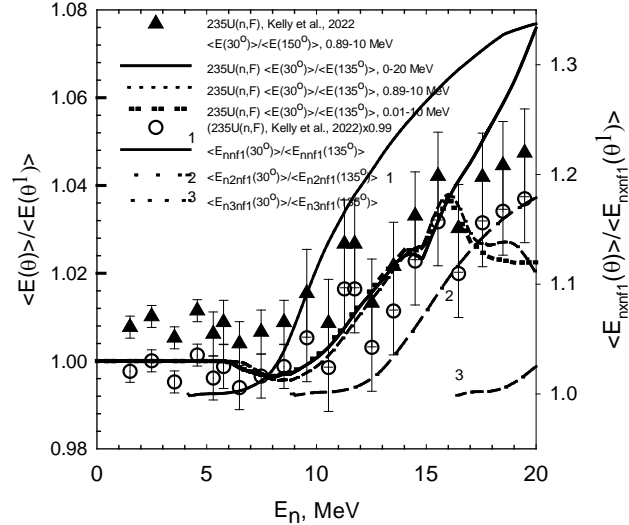


Fig. 2  $^{235}\text{U}(n,f)$  ratios of  $\langle E \rangle$ ,  $\langle E_{n,xnf} \rangle$  at  $30^\circ/135^\circ$ .

1. Kammerdiener J.L., UCRL-51232, 1972.
2. Kelly K. J., Gomez J.A., Devlin M. et al, Phys. Rev. C **105**, 044615 (2022).
3. Maslov V.M., LXXII International Conference “NUCLEUS-2022, Fundamental problems and applications”, Moscow, July, 11—16, 2022, Book of Abstracts, p.168, <https://events.sinp.msu.ru/event/8/attachments/181/875/nucleus-2022-book-of-abstracts-www.pdf>.
4. Mauss B., Taieb J., Laurent B. et al., [https://oecd-nea.org/dbdata/nds\\_jefdoc/jefdoc-2200.pdf](https://oecd-nea.org/dbdata/nds_jefdoc/jefdoc-2200.pdf), Nuclear Data Week, November, 2022, JEFDOC-2200.