

Prompt Fission Neutron Spectra of $^{233}\text{U}(n,F)$

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Prompt fission neutron spectra (PFNS) are produced for incident neutron energies from thermal up to 20 MeV. Simultaneous analysis of measured and calculated data for $^{235}\text{U}(n,F)$, $^{239}\text{Pu}(n,F)$ [1] and $^{233}\text{U}(n,F)$ maintains stronger justification for the predicted PFNS of $^{233}\text{U}(n,F)$. For the latter the reliable measured PFNS data are available at $E_n \sim E_{th}$ only. Pre-fission neutron spectra influence the partitioning of fission energy between excitation energy and total kinetic energy of fission fragments. For the reactions $^{233}\text{U}(n,F)$ and $^{235}\text{U}(n,F)$ shape of prompt fission neutron spectra (PFNS) strongly depends on the fissility of composite and residual nuclides (Fig.1). The correlation of these peculiarities with emissive fission contributions (n,xf) to the observed fission cross section and competition of the reactions (n,γ) and (n,xn) $^{1\dots x}$ is established. Exclusive neutron spectra (n,xf) $^{1\dots x}$ are consistent with fission cross sections of $^{235}\text{U}(n,F)$, $^{234}\text{U}(n,F)$, $^{233}\text{U}(n,F)$ and $^{232}\text{U}(n,F)$ reactions, as well as neutron emissive spectra of $^{235}\text{U}(n,xn)$ at ~ 14 MeV. Initial model parameters for $^{233}\text{U}(n,F)$ PFNS are fixed by description of PFNS of $^{233}\text{U}(n_{th},F)$. We predict the $^{233}\text{U}(n,xf)$ $^{1\dots x}$ exclusive pre-fission neutron spectra, exclusive neutron spectra of $^{233}\text{U}(n,xn)$ $^{1\dots x}$ reactions, total kinetic energy TKE of fission fragments and products, partials of average prompt fission neutron number and observed PFNS of $^{233}\text{U}(n,F)$. PFNS of $^{233}\text{U}(n,F)$ are harder than those of $^{235}\text{U}(n,F)$ PFNS, but softer than those of $^{239}\text{Pu}(n,F)$. Difference of average energies of PFNS $\langle E \rangle$ of $^{233}\text{U}(n,F)$ and $^{235}\text{U}(n,F)$ amounts to 1~3 %. At incident energies higher than ($n,2nf$) reaction threshold the observed PFNS may seem similar, though the partial contributions of $^{233}\text{U}(n,xf)$ and $^{235}\text{U}(n,xf)$ to the observed PFNS are quite different.

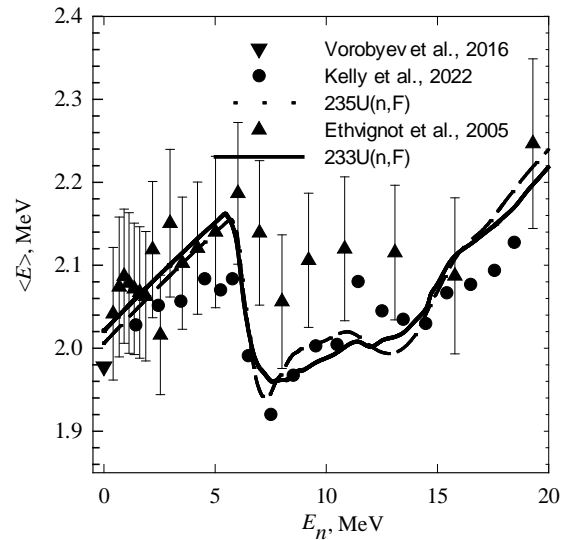
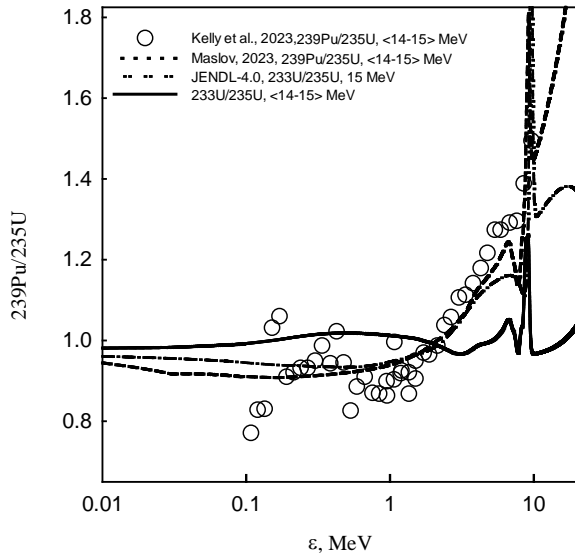


Fig.1. $^{235}\text{U}(n,F)$, $^{233}\text{U}(n,F)$, $^{239}\text{Pu}(n,F)$ PFNS ratios. Fig.2. $^{235}\text{U}(n,F)$ and $^{233}\text{U}(n,F)$ $\langle E \rangle$ of PFNS.

1. V.M. Maslov, Physics of Atomic Nuclei, 2023, vol.86, No. 5, pp. 627–629.