# <sup>233</sup>U(n,F) Prompt Fission Neutron Spectra

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Simultaneous analysis of measured data for <sup>235</sup>U(*n*,*F*), <sup>239</sup>Pu(*n*,*F*) and <sup>233</sup>U(*n*,*F*) maintains stronger justification for the predicted prompt fission neutron spectra (PFNS) of <sup>233</sup>U(*n*,*F*). Pre–fission neutrons influence the partitioning of fission energy between excitation energy and total kinetic energy of fission fragments. For the reactions <sup>233</sup>U(*n*,*F*) and <sup>235</sup>U(*n*,*F*) shape of prompt fission neutron spectra strongly depends on relative positions of (*n*,*xnf*) and (*n*,*xn*) reaction thresholds. The correlation of these peculiarities with emissive fission contributions (*n*,*xnf*) to the  $\sigma_{n,F}$  and competition of reactions (*n*,*nγ*) and (*n*,*xn*)<sup>1...x</sup> is established. Exclusive neutron spectra (*n*,*xnf*)<sup>1...x</sup> are consistent with  $\sigma_{n,F}$  of <sup>233</sup>U(*n*,*F*) and <sup>235</sup>U(*n*,*F*) and <sup>235</sup>U(*n*,*F*) and <sup>232</sup>U(*n*,*F*). Initial model parameters for <sup>233</sup>U(*n*,*F*) PFNS are fixed by description of PFNS of <sup>233</sup>U(*n*,*F*). We predict the <sup>233</sup>U(*n*,*xnf*)<sup>1...x</sup> exclusive pre-fission neutron spectra, exclusive neutron spectra of <sup>233</sup>U(*n*,*xn*) <sup>1...x</sup> reactions, total kinetic energy TKE of fission fragments and products, observed and partials of average prompt fission neutron number and observed PFNS of <sup>233</sup>U(*n*,*F*).

#### INTRODUCTION

Fissile nuclides <sup>233</sup>U may build-up in breeder or hybrid reactors. Nuclear data for <sup>233</sup>U+*n* interaction, with the exception for  $\sigma_{n,F}$  data, are scarce, especially as regards PFNS  $S(\varepsilon,E_n)$  of <sup>233</sup>U(*n*,*F*) in the range of <sup>233</sup>U(*n*,*xnf*)<sup>1,..x</sup> reaction. Data on PFNS  $S(\varepsilon,E_n)$  at  $E_n\sim 14.3$  MeV [1] remained the only available before long. Since model analysis in [1] was then over-simplified, relative contributions of pre- and post-fission neutrons in [1] disagree with later predictions [2, 3]. Measured data at  $E_n\sim E_{th}$  [4],  $E_n\sim 0.55$  MeV [5], and data [1] as well, were abandoned in all versions of BROND/ROSFOND, ENDF/B, JEFF and JENDL data libraries. The data of recent PFNS measurements for <sup>233</sup>U(*n*<sub>th</sub>,*f*) [6], <sup>235</sup>U(*n*<sub>th</sub>,*f*) and <sup>239</sup>Pu(*n*<sub>th</sub>,*f*) [6, 7] are discrepant with the data of [4]. Lumping [4, 6, 7] data in a spline fitting procedure of [8] would change predicted PFNS shapes of <sup>233</sup>U(*n*<sub>th</sub>,*f*), <sup>235</sup>U(*n*<sub>th</sub>,*f*) and <sup>239</sup>Pu(*n*<sub>th</sub>,*f*).

In differential PFNS data for <sup>235</sup>U and <sup>239</sup>Pu for  $E_n \sim 1.5-20$  MeV and  $\varepsilon \sim 0.01-10$  MeV [9–11] strong variations of average PFNS energies  $\langle E \rangle$  were observed around (n,xnf) thresholds.  $\langle E \rangle$  is rough signature of PFNS, however it was established in [9–11] that the relative amplitude of  $\langle E \rangle$  variation in case of <sup>239</sup>Pu(n,F) reaction is much weaker than in case of <sup>235</sup>U(n,F). That is due to influence of (n,xnf) reactions on fission observables when fission is preceded by pre-fission neutrons. In case of <sup>233</sup>U(n,F) reactions similar variations of  $\langle E \rangle$  were predicted in [2, 3, 8]. We intend to predict PFNS of <sup>233</sup>U(n,F) at  $E_n \sim E_{th}-20$  MeV.



Fig.1. Prompt fission neutron spectra of  ${}^{233}$ U( $n_{th}$ , f) relative to Maxwellian,  $\langle E \rangle = 2.0564$  MeV.

## <sup>233</sup>U(*n*,*f*) PROMPT FISSION NEUTRONS

Recent PFNS data for <sup>233</sup>U( $n_{th},f$ ) [6] added even more controversy: at range 0.02< $\varepsilon$ <5 MeV they support the evaluations of ENDF/B–VII [12] and JENDL-4.0 [13], while PFNS of both libraries disagree with data [4]. Data [4] are presented as spline approximation of [8], which summons empirical features of consistent analysis of <sup>233</sup>U( $n_{th},f$ ), <sup>235</sup>U( $n_{th},f$ ), <sup>235</sup>U( $n_{th},f$ ), <sup>239</sup>Pu( $n_{th},f$ ) and <sup>252</sup>Cf(*sf*) measured PFNS data [4]. In the energy range 5< $\varepsilon$ <11 MeV data of [6] for <sup>233</sup>U(n,f) support the evaluation of [8], which is based on data [4] fitting at 0.02< $\varepsilon$ <9.3 MeV.

The comparison of PFNS measured data [4, 6, 7, 9–11] for <sup>233</sup>U(*n*,*f*), <sup>235</sup>U(*n*,*f*) and <sup>239</sup>Pu(*n*,*f*) in the range  $E_{th} < E_{n} < E_{nnf}$  shows that enhanced soft neutron yield,  $\varepsilon \leq 1$  MeV is a common feature except PFNS of [6, 7]. In [6, 7] PFNS of <sup>233</sup>U(*n*<sub>th</sub>,*f*), <sup>235</sup>U(*n*,*f*) and <sup>239</sup>Pu(*n*,*f*) were measured relative to spontaneous fission neutron spectra (SFNS) of <sup>252</sup>Cf(*sf*). After various correction are applied to get absolute PFNS values, a number of systematic errors/uncertainties may appear, while the uncorrected cross ratios of various <sup>233</sup>U(*n*<sub>th</sub>,*f*), <sup>235</sup>U(*n*<sub>th</sub>,*f*) PFNS pairs might be quite sterile in that respect.



Fig. 2. Ratio of PFNS of  ${}^{233}U(n_{th},f)$  and  ${}^{235}U(n_{th},f)$  for thermal neutron-induced fission.

The ratios of PFNS for <sup>239</sup>Pu( $n_{th}$ ,f)/<sup>233</sup>U( $n_{th}$ ,f) and <sup>233</sup>U( $n_{th}$ ,f)/<sup>235</sup>U( $n_{th}$ ,f) [6, 7] and [4], contrary to absolute PFNS of <sup>233</sup>U( $n_{th}$ ,f), <sup>235</sup>U( $n_{th}$ ,f) and <sup>239</sup>Pu( $n_{th}$ ,f), quite agree with each other (Fig. 2). At  $E_n \sim E_{th}$  and  $E_n \sim 0.5$  MeV the ratios of calculated PFNS of <sup>239</sup>Pu(n,f) and <sup>235</sup>U(n,f) in the range 0.01< $\varepsilon$  <10 MeV weakly depend on the  $E_n$ . Calculated PFNS ratios of [2, 3, 8, 14], as well as present calculation, at  $E_n \sim E_{th}$  and  $E_n \sim 0.5$  MeV almost coincide with PFNS ratios <sup>239</sup>Pu( $n_{th}$ ,f)/<sup>233</sup>U( $n_{th}$ ,f) and <sup>233</sup>U( $n_{th}$ ,f)/<sup>235</sup>U( $n_{th}$ ,f) of [4, 6, 7]. It might be concluded that the hardest prompt fission neutrons are emitted in <sup>239</sup>Pu( $n_{th}$ ,f) reaction, while the softest PFNS is that of <sup>235</sup>U( $n_{th}$ ,f), PFNS of <sup>233</sup>U( $n_{th}$ ,f) takes intermediate position. Renormalization of model parameters at  $E_n \sim E_{th}$  after fitting data on total kinetic energy of fission fragments [15] amounts to rather small changes of PFNS: for <sup>239</sup>Pu(n, f) a decrease by ~2–3% at  $\varepsilon <1$  MeV, for <sup>235</sup>U(n, f) and <sup>233</sup>U(n,f) PFNS - shifts by ~1–2% [2, 3].

### <sup>233</sup>U(*n*,*xnf*) PROMPT FISSION NEUTRONS

Pre-fission neutrons emerging when  $E_n \gtrsim E_{nnf}$ , influence the PFNS  $S(\varepsilon, E_n)$  shape, total kinetic energy of fission fragments  $E_F^{pre}$  and fission products  $E_F^{post}$ , prompt fission neutron number  $v_p(E_n)$ , mass distributions and other fission observables. PFNS  $S(\varepsilon, E_n)$  is a superposition of exclusive spectra of pre-fission neutrons,  $(n,nf)^1$ ,  $(n,nf)^1$ ,  $(n,2nf)^{1,2}$ ,  $(n,3nf)^{1,2,3}-d\sigma_{nxnf}^k(\varepsilon, E_n)/d\varepsilon$  (x=0, 1, 2, 3; k=1,...,x), index x denotes the fission chance of <sup>234-x</sup>U and spectra of prompt fission neutrons, emitted by fission fragments,  $S_{A+1-x}(\varepsilon, E_n)$ :

$$S(\varepsilon, E_{n}) = \widetilde{S}_{A+1}(\varepsilon, E_{n}) + \widetilde{S}_{A}(\varepsilon, E_{n}) + \widetilde{S}_{A-1}(\varepsilon, E_{n}) + \widetilde{S}_{A-2}(\varepsilon, E_{n}) = v_{p}^{-1}(E_{n}) \cdot \left\{ v_{p1}(E_{n}) \cdot \beta_{1}(E_{n})S_{A+1}(\varepsilon, E_{n}) + v_{p2}(E_{n} - \left\langle E_{nnf} \right\rangle)\beta_{2}(E_{n})S_{A}(\varepsilon, E_{n}) + \beta_{2}(E_{n}) + \left\langle \beta_{2}(E_{n}) \frac{d\sigma_{nnf}^{1}(\varepsilon, E_{n})}{d\varepsilon} + v_{p3}(E_{n} - B_{n}^{A} - \left\langle E_{n2nf}^{1} \right\rangle) - \left\langle E_{n2nf}^{2} \right\rangle)\beta_{3}(E_{n})S_{A-1}(\varepsilon, E_{n}) + \beta_{3}(E_{n}) \cdot \left[ \frac{d\sigma_{n2nf}^{1}(\varepsilon, E_{n})}{d\varepsilon} + \frac{d\sigma_{n2nf}^{2}(\varepsilon, E_{n})}{d\varepsilon} \right] + v_{p4}(E_{n} - B_{n}^{A} - B_{n}^{A-1} - \left\langle E_{n3nf}^{1} \right\rangle - \left\langle E_{n3nf}^{2} \right\rangle - \left\langle E_{n3nf}^{3} \right\rangle) \cdot \left\{ \beta_{4}(E_{n})S_{A-2}(\varepsilon, E_{n}) + \beta_{4}(E_{n}) \left[ \frac{d\sigma_{n3nf}^{1}(\varepsilon, E_{n})}{d\varepsilon} + \frac{d\sigma_{n3nf}^{2}(\varepsilon, E_{n})}{d\varepsilon} + \frac{d\sigma_{n3nf}^{2}(\varepsilon, E_{n})}{d\varepsilon} \right] \right\}.$$

$$(1)$$

In equation (1)  $\tilde{S}_{A+1-x}(\varepsilon, E_n)$  is lumped contribution of *x*-chance fission to the observed PFNS  $S(\varepsilon, E_n)$ ,  $\langle E_{nxnf}^k \rangle$  – average energy of exclusive pre-fission neutron of  $(n, xnf)^{1..x}$  reaction, spectra  $S(\varepsilon, E_n)$ ,  $S_{A+1-x}(\varepsilon, E_n)$  and  $d\sigma_{nxn}^k(\varepsilon, E_n)/d\varepsilon$  are normalized to unity,  $\beta_x(E_n) = \sigma_{n,xnf}(E_n)/\sigma_{n,F}(E_n)$  is the contribution of *x*-th fission chance  $\sigma_{n,xnf}(E_n)$  to  $\sigma_{n,F}$ ,  $v_p(E_n)$  is the average number of prompt fission neutrons,  $v_{px}(E_{nx})$  – average number of prompt fission neutrons,  $v_{px}(E_{nx})$  – average number of prompt fission neutrons,  $\varepsilon_{n,xnf}(\varepsilon, E_n)$ , as proposed in [16], were approximated by the sum of two Watt [17] distributions with different temperatures, the temperature of light fragment being higher.

The differential measured PFNS at  $E_n \gtrsim E_{nnf}$  are also susceptible to systematic errors of various origin. In ratios of PFNS, especially of draft PFNS data, before corrections for the backgrounds, etc., these errors may be partially canceled [18, 19]. Figure 3 shows the  ${}^{239}\text{Pu}(n,F)/{}^{235}\text{U}(n,F)$  and  ${}^{233}\text{U}(n,F)/{}^{235}\text{U}(n,F)$  ratios of PFNS for  $E_n \sim 7 \div 8$  MeV. The averaged  ${}^{233}\text{U}(n,F)/{}^{235}\text{U}(n,F)$  ratio is very much different from that of  ${}^{239}\text{Pu}(n,F)/{}^{235}\text{U}(n,F)$ . The  ${}^{239}\text{Pu}(n,F)/{}^{235}\text{U}(n,F)$  ratios of differential PFNS at  $E_n \sim 7$ , 7.5 and 8 MeV mildly, bur significantly, fluctuate around averaged value. The respective  ${}^{233}\text{U}(n,F)/{}^{235}\text{U}(n,F)$  ratios of exclusive pre-fission (n,nf) spectra and  $\beta_x(E_n)$  values. In the energy range of  $E_n \sim 6\div 7$  [19] the fluctuations of PFNS at  $E_n \sim 6$ , 6.5 and 7 MeV are more pronounced for both  ${}^{239}\text{Pu}(n,F)/{}^{235}\text{U}(n,F)$  and  ${}^{233}\text{U}/{}^{235}\text{U}(n,F)$  ratios, since competition of (n,nf), (n,2n) and  $(n,n\gamma)$  depends on excitation energy. The observed PFNS of  ${}^{233}\text{U}(n,F)$  and  ${}^{233}\text{U}(n,f)$  are similar, as the increase of contribution of  ${}^{233}\text{U}(n,nf)$  is accompanied by decrease of  ${}^{233}\text{U}(n,f)$  reaction contribution  $S_{A+1}(\varepsilon,E_n)$ .



Fig.3. Ratios of PFNS  $^{233}$ U(*n*,*F*)/ $^{235}$ U(*n*,*F*) and  $^{239}$ Pu(*n*,*F*)/ $^{235}$ U(*n*,*F*) at  $E_n \sim 6 \div 7$  MeV.



Fig.4. Ratios of PFNS  ${}^{233}$ U(*n*,*F*)/ ${}^{235}$ U(*n*,*F*) and  ${}^{239}$ Pu(*n*,*F*)/ ${}^{235}$ U(*n*,*F*) at *E<sub>n</sub>*~14÷15 MeV.

At  $E_n \gtrsim E_{n2nf}$  integral emission spectrum of  $(n,nX)^1$  reaction,  $d^2 \sigma_{nnx}^1(\varepsilon, E_n)/d\varepsilon$ , could be represented as a sum of compound and weakly dependent on neutron emission angle preequilibrium components, and phenomenological function, modelling energy and angle dependence of neutron spectra, relevant for the <sup>233</sup>U excitations of 1~6 MeV. Angle-averaged  $\langle \omega(\theta) \rangle_{\theta}$  function,  $\omega(\theta)$  [20], is approximated as  $\langle \omega(\theta) \rangle_{\theta} \approx \omega(90^{\circ})$ , as described in [21]. Figure 4 compares calculated [20] and measured ratios of PFNS <sup>239</sup>Pu(n,F)/<sup>235</sup>U(n,F) [20] and <sup>233</sup>U(n,F)/<sup>235</sup>U(n,F) ratios at  $E_n \sim 13 \div 14$  MeV. The latter calculated present ratio is much discrepant with that of JENDL-4.0 [13], which just follows the shape of <sup>239</sup>Pu(n,F)/<sup>235</sup>U(n,F) of [13].

TKE values of  $E_F^{pre}$  are superposition of TKE for <sup>234-x</sup>U nuclides contributing to the observed fission cross section:

$$E_F^{pre}(E_n) = \sum_{x=0} E_{fx}^{pre}(E_{nx})\beta_x(E_n) .$$
 (2)

The excitation energy  $E_{nx}$  of A,..., A+1-x nuclides, formed after emission of  $(n,xnf)^{1,..x}$  prefission neutrons, depends on their average energies  $\langle E_{nxnf}^k \rangle$ :

$$E_{nx} = E_r - E_{fx}^{pre} + E_n + B_n - \sum_{x=0, 1 \le k \le x} \left( \left\langle E_{nxnf}^k \right\rangle + B_{nx} \right).$$
(3)

Kinetic energy  $E_F^{post}$  of fission products, which emerge after emission of pre-fission neutrons, but before  $\beta^-$ -decay, is defined as

$$E_F^{post} \approx E_F^{pre} \left( 1 - v_{post} / \left( A + 1 - v_{pre} \right) \right). \tag{4}$$

Weak variations of TKE values [15, 22], of both  $E_F^{pre}$  and  $E_F^{post}$ , in the vicinity of  $^{233}$ U(*n*,*xnf*) reaction thresholds are due to the decrease of excitation energy of (A+1-x) fissionning nuclides after emission of *x* pre-fission neutron [23]. Contribution of  $\sigma_{n,nf}$  to the  $\sigma_{n,F}$  of  $^{233}$ U(*n*,*F*), is larger than that of  $^{235}$ U(*n*,*nf*) to the  $\sigma_{n,F}$  of  $^{235}$ U(*n*,*F*) [20, 24], nonetheless the local bumps in TKE around  $^{233}$ U(*n*,*2nf*) and  $^{233}$ U(*n*,*nf*) reaction thresholds are weaker. That might be due to rather flat dependence on excitation energy of TKE for  $^{232,233,234}$ U, opposite to the case of TKE for  $^{235,236,237,238,239}$ U fissionning nuclides. To reproduce the observed dependence of  $E_F^{pre}$  on  $E_n$  in  $^{233}$ U(*n*,*F*) reaction one may assume linear dependence of first-chance fission TKE  $-E_{f0}^{pre}(E_n)$  (Fig. 5).

Average energy of prompt fission neutron spectra is its rather rough signature. Figure 6 evidence that the shapes of  $\langle E \rangle (E_n)$  in cases of  ${}^{233}U(n,F)$  and  ${}^{235}U(n,F)$  [20] are similar. Values of  $\langle E \rangle$  are presented here in the interval  $\varepsilon$ ~0.01–10 MeV. Our estimate of  $\langle E \rangle (E_n)$  for  ${}^{235}U(n,F)$  [20] reproduces the estimate of  $\langle E \rangle$  based on measured PFNS data [9, 26], especially around thresholds of  ${}^{235}U(n,nf)$  and  ${}^{235}U(n,2nf)$  reactions.



Fig. 5. Total kinetic energy TKE of  $^{233}$ U(*n*,*F*).



Fig.6. Average energy  $\langle E \rangle$  of <sup>233</sup>U(*n*,*F*) and <sup>235</sup>U(*n*,*F*) PFNS.

### CONCLUSION

A number of observed peculiarities in PFNS, TKE,  $v_p(E_n)$  correlate with the emission of pre-fission (n,xnf) neutrons, as predicted for the <sup>233</sup>U(n,F) and <sup>233</sup>U(n,xnf) and earlier for <sup>235</sup>U(n,F) and <sup>235</sup>U(n,xnf) [20, 24]. Cross ratios of PFNS of <sup>233</sup>U(n,F), <sup>235</sup>U(n,F) and  $^{239}$ Pu(*n*,*F*) reactions are compatible with measured data [11–13, 18, 19]. The correlation of PFNS shape and emissive ((n, xnf)) fission contribution to the observed fission cross section for  $^{233}U(n,F)$  and  $^{235}U(n,F)$  reactions is established. The net effect of these peculiarities is the occurrence of dips in  $\langle E \rangle$  in the vicinity of (n,nf) and (n,2nf) reaction thresholds and bumps in both  $E_F^{pre}$  and  $E_F^{post}$ . Amplitude of dips in  $\langle E \rangle$  of <sup>233</sup>U(*n*,*F*) PFNS is quite similar to that observed in PFNS of  $^{235}U(n,F)$  reaction, notwithstanding the appreciable differences of  $^{233}$ U(*n*,xnf) and  $^{235}$ U(*n*,xnf) reaction contributions to the observed fission cross sections  $^{233}$ U(*n*,*F*) and  $^{235}$ U(*n*,*F*), respectively. That is explained by relatively large contributions of  $v_{px}(E_{nx})$  as compared with  $v_{pre}(E_n)$  for the reaction <sup>233</sup>U(*n*,*F*). PFNS of <sup>233</sup>U(*n*,*F*) are more hard than those of <sup>235</sup>U(*n*,*F*) PFNS, but softer than those of <sup>239</sup>Pu(*n*,*F*). Difference of average energies of PFNS  $\langle E \rangle$  of <sup>233</sup>U(*n*,*F*) and <sup>235</sup>U(*n*,*F*) amounts to 1~3 %. At incident energies higher than (n, 2nf) reaction threshold the observed PFNS may seem similar, though the partial contributions of  $^{233}$ U(*n*,*xnf*) and  $^{235}$ U(*n*,*xnf*) to the observed PFNS are quite different. It might be argued that correct estimate of the exclusive pre-fission (n, xnf) neutron spectra and modelling of spectra of neutrons emitted from excited fission fragments gives a robust prediction of PFNS for <sup>233</sup>U(*n*,*F*) for incident neutron energies  $E_n \sim E_{th}$ -20 MeV with a precision and reliability comparable to those attained for  $^{235}U(n,F)$  PFNS.

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